

**MEMORANDUM OF AGREEMENT
SASD DEVELOPMENT PROGRAM**

**1987 Generic R&D Final Report
A Program to Compute the
Aerodynamic Solar Radiation
Forces and Torques on Satellites**

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Prepared by:

V. Carrara

Date:

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Approved by:

J. Soul - Program Manager

B. Berridge - Director Programs

Spar Aerospace Limited

**Satellite and
Aerospace
Systems
Division**

21025 Trans-Canada Highway
Ste-Anne-de-Bellevue, Quebec
Canada H9X 3R2
Telephone (514) 457-2150
Telex 05-822792

SPAR


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Author(s): Valdemir Carrara

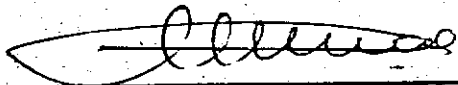
Section: SASD, Systems Engineering

Approval: 
D. Jung (Supervisor's Signature)

Approval: 
J.R.G. Cox (Manager's Signature)



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F. Passier



Technical Report Abstract

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Abstract - Briefly summarize objectives, methods, results, & applications. Type single spaced.

Programs to evaluate forces and torques acting on low Earth orbit satellites have been developed by SPAR since 1981 [1], [2], [3] and [4]. As these perturbations are strongly dependent on the spacecraft's geometry, a computer program that utilizes a standard method to describe the satellite shape becomes important. This work presents the ARFTS (Aerodynamic and solar Radiation Forces and Torques on Satellites) program, which uses a NASTRAN file (and hence the CAD facilities) to inform the program of the position and orientation of each surface of the spacecraft. The program supports spacecraft of varying configurations or with articulated appendages (like rotating solar arrays or dual-spin spacecraft), allowing misalignment studies (for orbit and attitude control system design) and torque-reduction analyses.

Keywords - State identifying terms to a maximum of five.

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INTRODUCTION

The main subject of this work was to develop a computer program that calculates the drag coefficient for a satellite of general shape, as well as the aerodynamic and solar radiation forces and torques. The constraints imposed on the program were the fast running, user's easy interaction and relatively large precision.

The developed program automatically propagates the orbit and calculates the forces and torques for each point. The outputs are the aerodynamic forces and torques as functions of time (the user can choose the reference system in which he wishes the forces and torques to be presented), as well as the linear and angular impulses utilized for orbit and attitude control specifications, respectively. Others applications can include the spacecraft geometry analysis, torque reduction studies, drag and solar pressure coefficient determinations, and numeric orbit and attitude propagations.

The results are presented in several files, including the atmospheric local density and temperature (dynamic model), geocentric coordinates of the spacecraft (latitude, longitude and altitude), direction of incidence of the atmospheric molecules, sun-aspect angles and an orbit resume of the variation in the linear and angular momenta of the satellite.

The spacecraft geometry must be provided by the user, through a data file, assuring independence of the program with the satellite shape and increasing its applicability. The format of the spacecraft geometry data file is compatible with the NASTRAN (NASA STRuctured ANalysis), largely utilized in finite elements calculations, for satellite thermal control analysis and structure design. It makes it easy for the user to change the geometry (and check it) using the CAD facilities [5].

To avoid loss of generality, the program was constructed in a way to support non-rigid satellites, such as dual-spin spacecraft or sun-oriented solar arrays. The geometry of the appendages is also given in the geometry data file, but the position, orientation and angular ratio of each appendage can be changed without changing the spacecraft configuration. It is particularly useful in misalignment and torque reduction studies as well as the analysis of forces and torques in satellites with variable geometry.

As the forces and torques acting on satellites can be difficult or easy to check depending on the coordinate system where they are plotted, Section 2 defines all the systems utilized by the program. Section 3 describes the program structure without the employed theory, which is readily available in the literature ([4], [6] and [7], for example). Section 4 includes an extensive description of the input elements to the program and the output files. The conclusions of the work are presented in Section 5. Finally, the program is applied to the RADARSAT satellite as an example, shown in Appendix A. Appendix B presents the main program and subroutines listing.

2.0

SYSTEM OF COORDINATES

Due to the large number of coordinate systems used by the program, a brief description of each frame is presented here :

- 1- Inertial system. This system has the X axis pointing to the vernal equinox and the Z axis passing through the Earth's North Pole. Its center coincides with the Earth center.
- 2- Ascending node system. By definition, this system is geocentric and inertial too, but it has its X axis passing through the orbit ascending node and the Z axis through the north pole. The name comes from the angle between its X axis and the inertial system X axis, as shown in Figure 2.1.

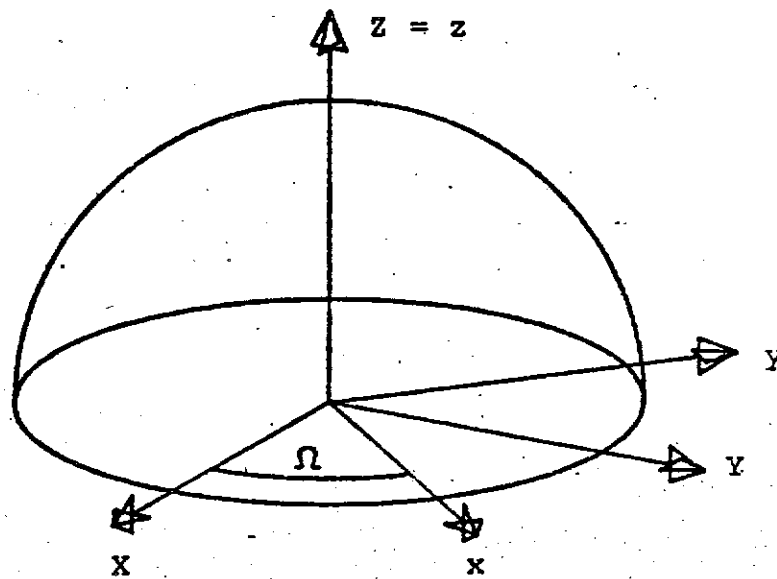


Fig 2.1. Ascending node system (xyz) and inertial system (XYZ). Ω is the orbit right ascension of the ascending node.

- 3- Orbit plane system. Like the last system, this system has its X axis coincident with the orbit ascending node but its Z axis normal to the orbit plane. Its center coincides with the Earth center. The orbit plane system is related to the ascending node system by a rotation about its X axis by an angle equal to the orbit inclination, as can be seen in Figure 2.2.

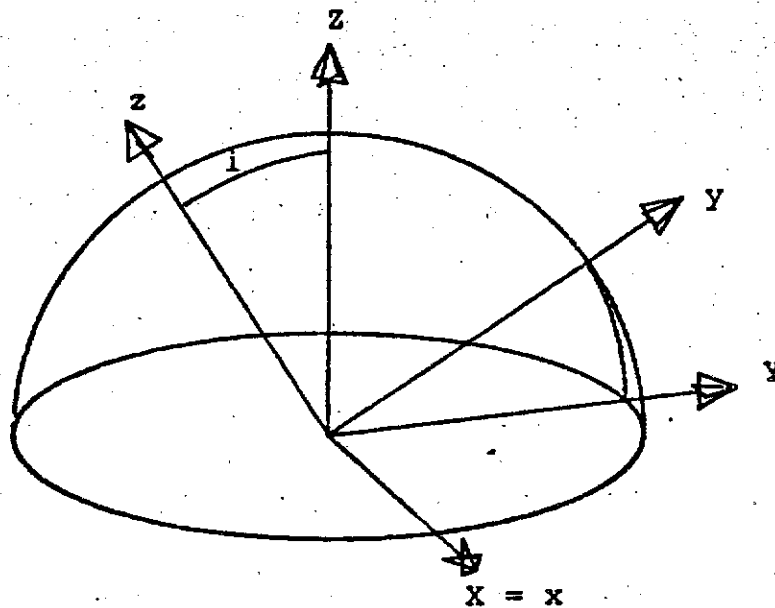


Fig. 2.2. Orbit plane system (xyz) and ascending node system (XYZ). i is the orbit inclination.

- 4- Orbital system. This is a non-inertial system, centered at the Earth center and with its X axis passing through the instantaneous position of the spacecraft. Its Z axis is normal to the orbit plane, as shown in Figure 2.3. The satellite velocity points towards the positive Y direction. The time-dependent true anomaly, f , relates the orbital system to the orbit plane system.

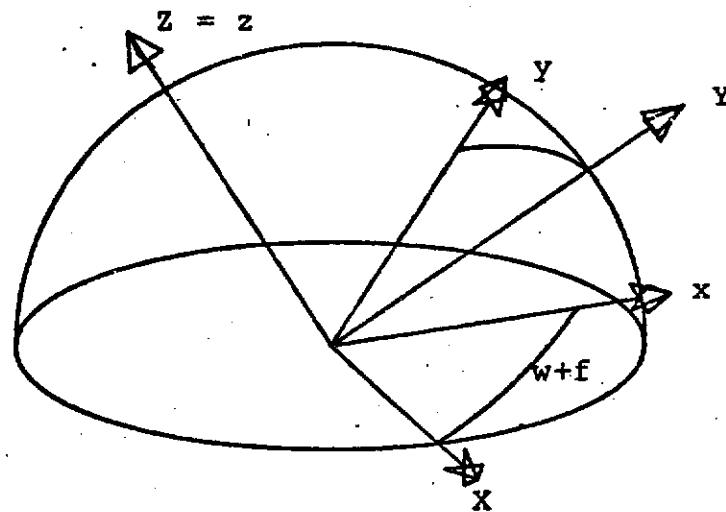


Fig. 2.3. Orbital system (xyz) and the orbit plane system (XYZ). The angle between X and x is the sum of the perigee argument w and the true anomaly f .

- 5- Spacecraft system at $t=0$. This system coincides with the spacecraft system at the origin of the time and remains inertially fixed, despite the satellite angular motion. It can be related to any one of the precedings systems, through a X-Y-Z rotation as in Figure 2.4, with angles provided as input parameters.

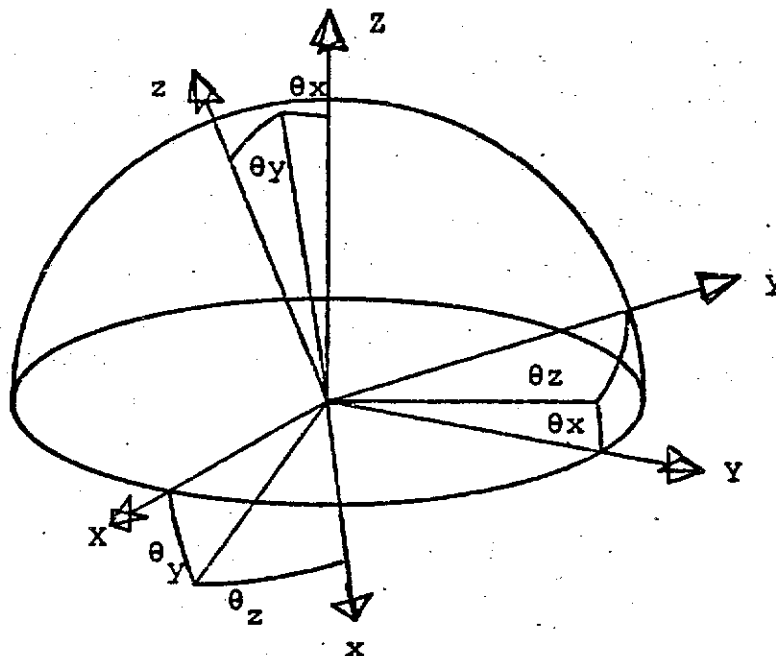


Fig. 2.4. Spacecraft system (xyz) at $T=0$ and the attitude related system (XYZ). The angles θ_x , θ_y and θ_z define the relation among the systems.

- 6- Spacecraft system. This system is fixed to the spacecraft structure and the satellite geometry, as well as the geometry of the appendages, must be described in it. The spacecraft system can also be related to all of the preceding four systems. The relationship is a X-Y-Z rotation as in Figure 2.4, but it doesn't remain inertially fixed since the related system is not inertial (orbital system).

Systems 1 to 4 are useful for the presentation of the forces; systems 5 and 6 are preferred for the torques. The program user must specify the system in which the aerodynamic and solar radiation forces and torques are to be output.

3.0

THE PROGRAM STRUCTURE

The program was structured in a way so as to utilize only data files for both input and output. All the input parameters to the program are given in two files, containing the spacecraft geometry and the orbital data. The outputs are printed in several files, containing the results of the aerodynamic and solar radiation forces and torques, orbit propagation data, atmospheric properties, sun-angles related to the satellite body, spacecraft composition in terms of its surface elements and a resume of the linear and angular impulses over the satellite in each orbit. The inputs and outputs will be described with more detail at Section 4.

Initially, the program reads the spacecraft geometry from the appropriate data files and stores all the surface elements in arrays. Because large memory is consumed by these arrays, the program is limited to a maximum of 50 surface elements. All the necessary rotation matrices are calculated and the orbit propagation is initialized. The program uses an analytical orbit propagator which takes into account only the Earth oblateness effects. Due to the normally short time period of the orbit propagation in a typical run, this simplification of the orbit perturbations is acceptable.

For each point, the program computes the sun position, atmospheric density, mean molecular weight, and the velocity of the atmosphere. The sun position is obtained by the SUN subroutine (Kuga [8]), that employs an analytical Earth orbit propagator without corrections for the Earth-sun distance due to eccentricity. This causes an error less than 7% in the radiation forces.

The atmospheric properties are calculated by the ASDAMO subroutine (from Analytical Static and Dynamic Atmosphere Models), an analytical version (Lafontaine [9]) of the Jacchia's J77 model [10]. All the necessary parameters for the atmospheric model are calculated internally with exception of the exospheric temperature, a function of the daily 10.7 cm wavelength solar flux. The user must provide this data through the input file, as the other input parameters.

The sun position and the satellite velocity vectors are then rotated to the spacecraft system and a summation of the aerodynamic and solar radiation forces is made over the spacecraft main body. These vectors are rotated again to each appendage system and additional summations are performed over each appendage. The output values for this point in time are then evaluated and printed. Once a whole orbit is completed, the program outputs the net impulse per orbit. The program then continues calculations for the next orbit until the desired time period has been processed.

Care must be taken with the aerodynamic and solar radiation shadowing on the spacecraft, since the program doesn't automatically eliminate the shadowed areas from the summation. This is particularly important on concave-shape spacecrafts (i. e. when one spacecraft external surface sees others). The expressions for the elementary aerodynamic force are obtained from the assumption of total surface exposure by the flow, meaning that any other element can be in its field of view (the hemisphere above the surface), independent of the flux direction. Nevertheless, the resulting force is affected more with shadowing in the flux direction than in any other direction. For that reason it is a good approximation to consider the aerodynamic shadowing like a solar radiation shadowing, removing the surfaces which are behind the others in the flow direction from integration. It can be done by properly eliminating the shadowed surfaces in the spacecraft geometry data file: those in the satellite velocity direction when computing the aerodynamic forces and those in the sun direction when calculating the solar radiation forces.

The main program is stored in the ARFTS.FTN file (from Aerodynamic and solar Radiation Forces and Torques on Satellites). Almost all the auxiliary routines are in the AUROT.FTN file. The atmospheric model routines are stored in the ANATM.FTN file (ANalytical ATMospheric Model). The program also utilizes some orbit propagation routines, provided by the ORBIT.FTN file and described in Kuga [8].

The default names of the input files and the extension of the output file names can be changed during the running command:

ARFTS,<INPUT.FILE1>,<INPUT.FILE2>,<OUTPUT EXT>

INPUT.FILE1 is a NASTRAN file containing all the spacecraft definition surfaces. Its default name is SATGE.DAT, from SATEllite GEometry. The INPUT.FILE2 is an input file with several parameters, like orbital elements, date, the composition and properties of the surfaces in the spacecraft, the position of the satellite centre of mass, etc. Its default name is INELE.DAT (from Input ELEments). The output file extension has the DAT string as default. Nevertheless, it is useful sometimes to change the output extension, to avoid overwriting previous results when running the program. The new extension string is passed to the program by the <OUTPUT EXT> parameter of the runstring.

Figure 3.1 represents the subroutines hierarchy of the ARFTS program. Some of these routines (CONSTP, DJM, VARELK, KEPLER and SUN) came from the ORBIT.FTN file [8]. The atmospheric model comprises the ASDAMO, ASMADE, DIVARA, GEOACA, SEALAT, SEMIAN, TEMLO and AMOWEI subroutines.

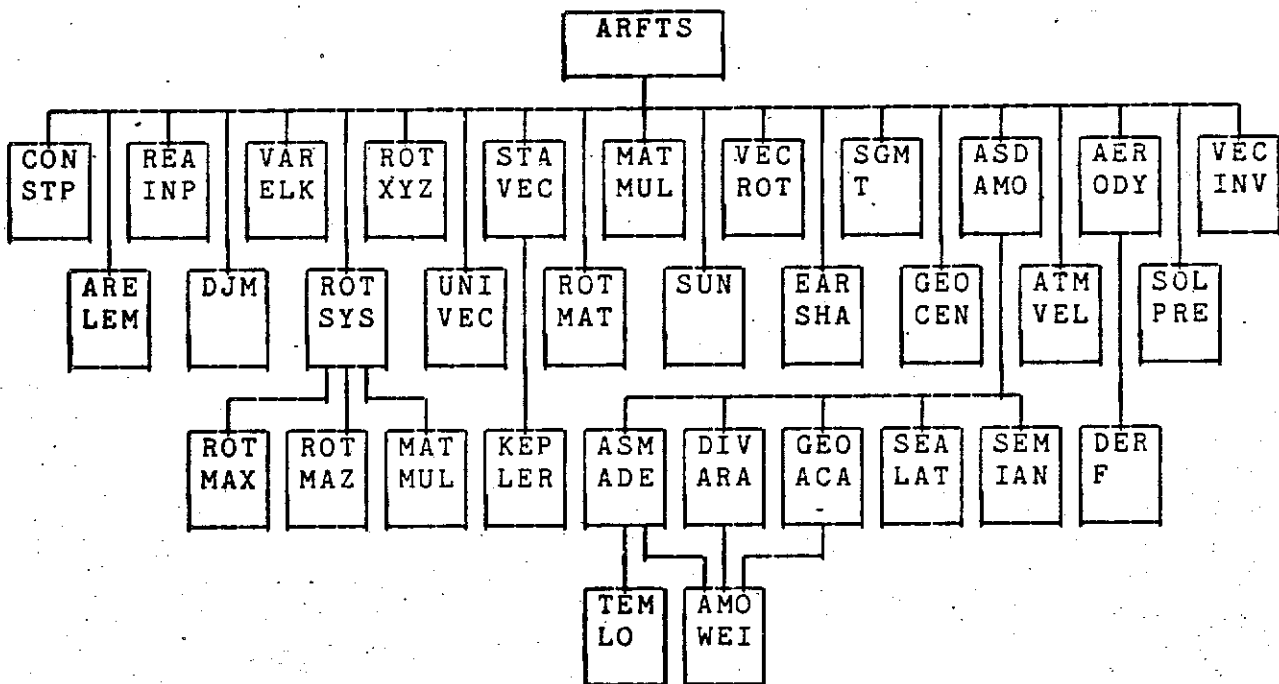


Figure 3.1. ARFTS structure.

4.0 INPUT AND OUTPUT DATA FILES

4.1 INPUT FILES DESCRIPTION

The inputs for the program are stored in two files: SATGE.DAT and INELE.DAT are the default names.

The SATGE.DAT file contains a NASTRAN description of the spacecraft geometry. The NASTRAN is a widely used language to perform structural analysis. It was chosen because it is easy to employ and it decreases the user's spent time to create a geometry description file or subroutine (as a NASTRAN file is often already available or can easily be created using CAD utilities).

As the program is interested only in the spacecraft geometry description, there are few NASTRAN commands (or cards) recognized by the program: GRID, CTRIA3, CQUAD4 and ENDDATA - the remainder are ignored. If the file contains surface definition cards other than the CTRIA3 and CQUAD4 commands, the user must modify the ARELEM subroutine in such a way as to accept the new cards.

GRID A GRID card gives the program information about the position of a spacecraft node, i. e. the three coordinates of a characteristic point of the satellite structure. A GRID card is identified by the program in the format: A8, 2I8, 3F8.2. The last variable in the 2I8 field and all the variables after the 3F8.2 field positions will be ignored, as they have no useful meaning to this program (only to the NASTRAN program). The first field, A8, must contain the "GRID" character string, starting on column 1 and filled with blanks until column 8, inclusive. The first integer in the 2I8 field is the identification number of the node, and will be used by the surface definition cards, CTRIA3 and CQUAD4. The 3F8.2 position contains the rectangular XYZ coordinates in millimeters of the node given in the spacecraft system. All the node elements are stored in an array by the program. Although the program capacity has been limited to 50 nodes, it is possible to alter this value by changing the NGRD variable of the PARAMETER statement in the ARELEM subroutine. A message error will be printed if there are more than the maximum number of nodes in the data file.

CTRIA3 A card with the string "CTRIA3" starting at column 1 is recognized by the program as a triangular surface definition. It's format must be A8 (for the alphanumeric string) and 5I8. The first integer is the identification surface number, used in the BODYAP cards of the INELE.DAT file. The second is unused. The 3 remaining values are the node identification numbers of a GRID definition card, which fix the corners of the triangular surface. Attention should be given to the sequence in which the numbers appear: the program calculates the external normal by the right-hand rule of the first, second and third node points.

CQUAD4 This card defines a quadrilateral surface of the spacecraft structure, and has the same format of a CTRIA3 card, except than a fourth node identification number must be given. The read format is then A8, 6I8. The second integer number and the numbers after the last integer field (which sometimes appears in NASTRAN files) are ignored. The surface external normal is defined, as a CTRIA3 card, by the right-hand rule of the first, second and third node points.

ENDDATA This card, with the string "ENDDATA" starting on column 1, informs the program that the end of data has been reached. It causes the program to end the surface computations or rewind the file if there are more data to be read.

Appendix A lists a sample of a SATGE.DAT file, with a NASTRAN definition geometry for the RADARSAT configuration. Nevertheless it is not a typical NASTRAN file: NASTRAN has many other commands not given there.

The surface elements are also stored in arrays as the node points. Consequently, the program capacity has been limited to 50 surface elements maximum. If more than 50 surfaces are needed to describe the spacecraft geometry, the program can be altered so as to accept additional nodes by changing the NSUR variable of the PARAMETER statement to the desired value.

The second input file, INELE.DAT, defines several input variables, described below. To make it easy for the user, the INELE.DAT definition cards have total compatibility with the NASTRAN cards. It was separated into a different file, however, as the SATGE.DAT can be the actual NASTRAN file (without change). The file comprises 10 different command cards: ORBIT, DATE, SOLFLUX, BARICEN, PROAREA, MAINSYS, APPSYS, OUTSYS, BODYAP and ENDDATA. All these cards, except the APPSYS and BODYAP cards, may appear only once in the file. If more than one card is encountered by the program, subsequent cards will be ignored. The APPSYS card, which informs the existence of spacecraft appendages to the program, appears once for each appendage. BODYAP cards define the spacecraft surface composition and properties, having as many cards as necessary to define the satellite geometry. All the cards without these names are ignored by the program, allowing the insertion of comments to the file as well as to permit data to be added or deleted by simply changing a character in the card name.

Unlike the SATGE.DAT file, all the data in the INELE.DAT file are in free format. Therefore, it isn't necessary to comment on the format of the data, only its meaning to the program. Nevertheless, the command name must be in A8 format (starting on column 1), as in SATGE.DAT file.

ORBIT This command contains the orbital elements of the satellite. They are, respectively: perigee and apogee altitudes, in km and range of 100 to 2000 km (limit imposed by the atmospheric model constraints); orbit inclination, in degrees; right ascension of the ascending node, in degrees; perigee argument, in degrees and mean anomaly at the date, in degrees. The semi-major axis is calculated with the perigee and apogee values, using an Earth's radius of 6,378.16 km.

DATE This card gives to the program the time for the orbital elements: month, day, year, hour, minute (integers) and second (decimal). The two last values are the time step size and the total time for the orbit propagation, both in seconds.

SOLFLUX The values that this card must contain are the daily solar flux at a 10.7 cm wavelength (adjusted for the correct Earth-sun distance), the averaged flux over six solar rotations (162 days approximately) and the geomagnetic activity index Kp. These values are utilized in the atmospheric properties computations. More details of how to get and utilize this data can be found in refs. [10] and [11]. In case of simulation, the same value for both the daily and the averaged solar flux must be used. Mean values are: 70 ($\times 10^{-22}$ W/m²/Hz) at minimum solar activity and 240 ($\times 10^{-22}$ W/m²/Hz) at maximum solar activity. The geomagnetic activity index depends on the activity in the solar corona, increasing during the solar storms (which are more frequent near the maximum activity periods). It can vary from 0.0 (during quiet days) to 9.7 (during strong solar storms), in steps of 1/3.

BARICEN This command gives the three rectangular coordinates of the spacecraft's center of mass, in meters, in the spacecraft system.

PROAREA This card must contain a spacecraft characteristic area in square meters to compute the aerodynamic drag coefficients. If the area is equal to the projected satellite area in the velocity direction, then the drag coefficient will be approximately 2 for high speed ratio values (low altitudes).

MAINSYS The spacecraft system must be related to one of the frames described on Section 2, to find the spacecraft's attitude by the program. The related spacecraft frame is the first parameter in this card, corresponding to the number of the system described on Section 2. As the related frame can't be the spacecraft system itself, it is restricted to the first four systems:

- 1- Inertial system.
- 2- Ascending node system.
- 3- Orbit plane system.
- 4- Orbital system.

The next parameters are, respectively: the three components of a X-Y-Z rotation relating the frame to the spacecraft system (see Fig. 2.4), in degrees and the three components of the spacecraft angular velocity relative to that system, in revolutions per minute, in the spacecraft system.

APPSYS

These cards provide the relation between the spacecraft body and the appendages. For each appendage (corresponding one APPSYS card), the command must contain: the identification structure number (corresponding to that in the BODYAP cards), the three coordinates of the position, in meters, of the origin of the appendage-fixed system; the three components of the angular velocity of the appendage, in revolutions per minute and the three components of a X-Y-Z rotation, in degrees, relating the spacecraft system to the appendage orientation.

The position of the appendage-fixed system origin informs the program that the appendage is articulated at this point, with three degrees of freedom (the orientation of the appendage). As the program understands that the articulation point is fixed in both the appendage and the spacecraft's body, care must be taken if one wants to change the appendage position by simply altering the articulation point position in the APPSYS card. Sometimes is necessary to alter also the geometry description data file (SATGE.DAT) as the program adopts that geometry at the initial time (T=0).

The angular velocity of each appendage tells the program that the appendage must rotate with time. This information must be given when the spacecraft is composed by rotating counter-parts (as dual-spin) or when an Earth-pointing satellite has sun-oriented solar arrays, for example.

The last three components of the APPSYS instruct the program to reorient the appendage through a X-Y-Z rotation applied to the articulation point as shown in Figure 4.1. These angles can be used to direct the solar arrays to the sun direction before starting calculations (as a phase angle), to introduce misalignments in the appendage positions or even to reorient them for a new spacecraft's geometry.

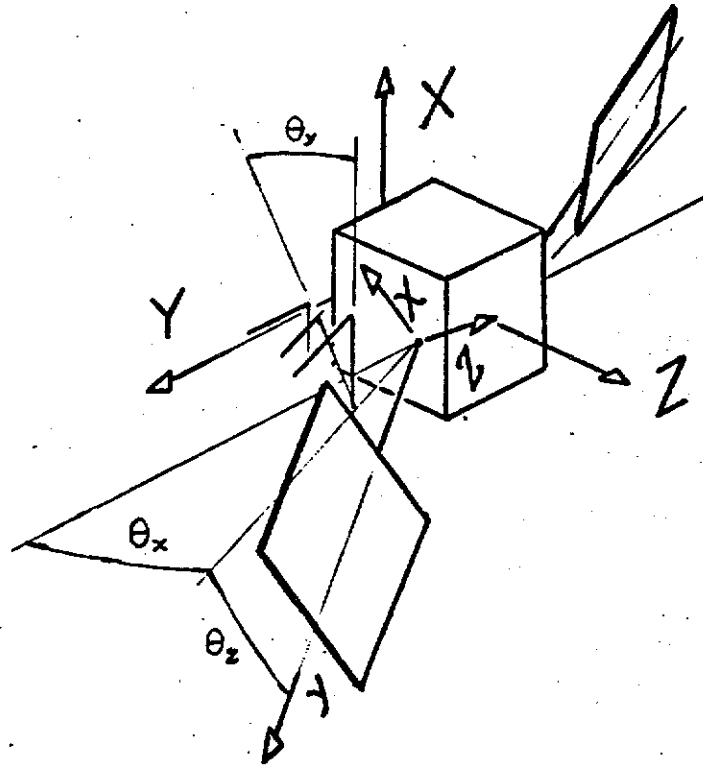


Fig. 4.1 Spacecraft body-fixed system X-Y-Z and appendage-fixed frame x-y-z with origin at the articulation point.

These parameters define a new system of coordinates (x-y-z in Figure 4.1) fixed at each appendage, with origin at the articulation point. Both the position of the articulation point and the angular velocity of the appendage must be in spacecraft system coordinates. The appendage-fixed system is then found by a translation of the spacecraft system to the articulation point, a rotation about the angular velocity direction until the calculated appendage direction and, finally, a X-Y-Z rotation to the desired appendage orientation.

Note that the appendage position is fixed in the spacecraft body, when the SATGE.DAT file is created. However, it can be oriented in other positions and rotated without changing the geometry file, by simply altering the APPSYS commands. It is useful to simulate the appendage rotation in geo-pointing satellites with sun-oriented arrays. Other applications include appendage misalignment studies and

analysis of the appendage and structure composition, with the aim of minimizing disturbance torques.

OUTSYS The OUTSYS card defines the coordinate system in which the user wants the results printed. Its parameters are, respectively, the system number for the aerodynamic force, aerodynamic torque, solar radiation force and solar radiation torque. The output systems can be (see Section 2 for frame definitions):

- 1- Inertial system.
- 2- Ascending node system.
- 3- Orbit plane system.
- 4- Orbital system.
- 5- Spacecraft system at $t=0$.
- 6- Spacecraft system.

BODYAP These cards give the surface composition of the spacecraft to the program. The first parameter is the identification structure number (0 is reserved for the main body or the spacecraft itself, 1 is for the first appendage, 2 for the second and so on until 10 appendages, maximum). The second is the surface element identification number, as described in the CTRIA3 and CQUAD4 commands of the SATGE.DAT file. If the number is negative, then the surface normal will be considered to be in the opposite direction to that given in the SATGE.DAT file. The back-side of some structures like solar arrays can be composed easily in this manner. Note that each BODYAP command corresponds to one satellite surface, meaning that the surface identification number can't appear more than once in BODYAP card, except for a back-side composition (one positive and one negative). The next 7 values are the surface properties, respectively: the tangential and normal aerodynamic coefficients (see refs. [4], [6] and [7] for more details); the specular reflectivity, diffuse reflectivity and the absorption coefficient in the solar spectrum; the infra-red emissivity; and the surface temperature, in Kelvin degrees.

There is no fixed sequence for the command cards. They can be placed anywhere in the INELE.DAT file. However, the name must start in column 1 and be completed with blanks up to column 9.

4.2

OUTPUT FILES DESCRIPTION

All the most important results are printed in separated files, in order to simplify the user's analysis. The files are: "PLANE.DAT", "EXTSE.DAT", "AEROF.DAT", "AEROT.DAT", "SORAF.DAT", "SORAT.DAT", "ORBPR.DAT", "ADRAG.DAT", "SUNAN.DAT" and "ORBMO.DAT". The first and second files are the internal results computed from the SATGE.DAT file.

After reading the surface definition commands (CTRIA3 and CQUAD4), the program creates the PLANE.DAT file, storing values of the surface identification number, unit external normal, position of the center and area.

The second file, EXTSE.DAT, shows the spacecraft composition, after the BODYAP commands (see the INELE.DAT file description, in Section 4.1). The printed values are the structure number (spacecraft body must always have the 0 number), element number, unit external normal, position of the center of pressure and the element area. These results can be used to check the correct understanding of the input data.

The AEROF.DAT, AEROT.DAT, SORAF.DAT and SORAT.DAT files are, respectively, results of the: aerodynamic forces, aerodynamic torques, solar radiation forces and solar radiation torques. Just after the title, the program prints the coordinate system where the forces, torques and impulses were calculated. The first column is the time, in seconds, for all the files. The next three are the force (in N) or the torque (in Nm) and the last three are the linear (in Ns) or angular (in Nms) impulse components.

ORBPR.DAT contains data from the orbit propagation as function of the time: the orbit mean anomaly, the spacecraft geocentric distance and velocity, the longitude and latitude of the sub-satellite point, the geocentric height, and the right ascension and declination of the sun. The sun position is calculated from the date and hour of the day, using the SUN subroutine, implemented by Kuga [8,12]. This file contains also printed values of the ORBIT, DATE, SOLFLUX, BARICEN, MAINSYS, APPSYS and PROAREA commands.

The ADRAG.DAT is a table with some intermediate results of the aerodynamic force computations: the local atmospheric temperature and density, the speed ratio (the ratio between the spacecraft velocity and the most probable speed of the molecules of the atmosphere). The

direction of the satellite velocity relative to the velocity of the atmosphere in the spacecraft system coordinates is also printed. Here, ALPHA is the angle between the YZ plane and the velocity direction and BETA the angle between the Y axis and the projection of the velocity direction in the plane YZ. These angles are useful to check the satellite attitude. The other values are: the spacecraft system components of the drag coefficient and the drag coefficient itself. This file also contains a resume of the input elements, as in the ORBPR.DAT file.

The SUNAN.DAT file has the sun aspect angles for each structure of the spacecraft, as a function of the time. These angles are defined in the same way as in the ADRAG.DAT file (with ALPHA being the angle between the YZ plane and the sun direction and BETA the angle between the Y axis and the projection of the sun direction in the YZ plane). These angles are in spacecraft system coordinates. They are useful to check the satellite attitude and also the appendage orientation. As for the two last files, the SUNAN.DAT contains the printed values of the input parameters.

The last output file, ORBMO.DAT is a per orbit evaluation of the forces, torques and momenta variations (the accumulated impulse over one orbit). It contains these values for each completed orbit during the propagation time, identified by the orbit number and the time when the results were calculated. For the aerodynamic and solar radiation, the minimum and maximum force and torque intensities encountered at this orbit, as well as its mean values, are printed in the coordinate system given by the OUTSYS command card. The same system is used to print the minimum, maximum, the difference between the minimum and maximum, the build-up and the total accumulated values of the linear and angular momentum variations (impulses). The minimum and maximum are the lower and the higher impulse values encountered at this orbit. Note that the impulse is always made equal zero at the beginning of a new orbit. Its difference is the maximum variation of the momentum during each orbit. The build-up values are the impulse at the end of the orbit. A small value for the build-up for impulse (relative to the difference value) means a symmetrical force or torque on this axis and, consequently, a small change in the spacecraft attitude about this axis. The accumulated value is total momentum change since the start of the simulation.

The output files are listed in Appendix A, an example of application to RADARSAT.

5.0

CONCLUSION

The aerodynamic and solar radiation are the most important low Earth orbit perturbations acting on satellites. However, depending on the spacecraft geometry, the other perturbations - gravity gradient torque, magnetic torque, etc. - can supplant the first ones. The user should compare the magnitude of these forces before making any judgement, in order to avoid neglecting important results.

As the program gives instantaneous values of the forces and torques, sometimes it will be necessary to change some of the input parameters (like the starting date, solar flux, orbital elements, etc.) so as to assure a reasonable sample of the perturbations during the mission duration.

The inability of the program to automatically remove the shadowed surfaces from integration precludes its application to some highly convex satellites. Also, the NASTRAN is not a good format to implement a shadow analysis, because it takes too much computer time to evaluate the forces for a single orbit point (it could be better to describe the spacecraft geometry by means of a subroutine, where the shadowed areas were removed from integration by the user, during the subroutine construction).

6.0 REFERENCES

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APPENDIX A

APPLICATION TO RADARSAT

The program was applied to the dawn-dusk configuration of RADARSAT satellite [13]. It has two large solar arrays fixed on the spacecraft structure, as shown in Figure A-1, drawn with points extracted from the SATGE.DAT file. For that configuration the main origin of the torques are the spacecraft assymetry due to the Synthetic Aperture Radar (SAR) positioning, the solar array misalignments and changes in the spacecraft center of mass due to fuel consumption.

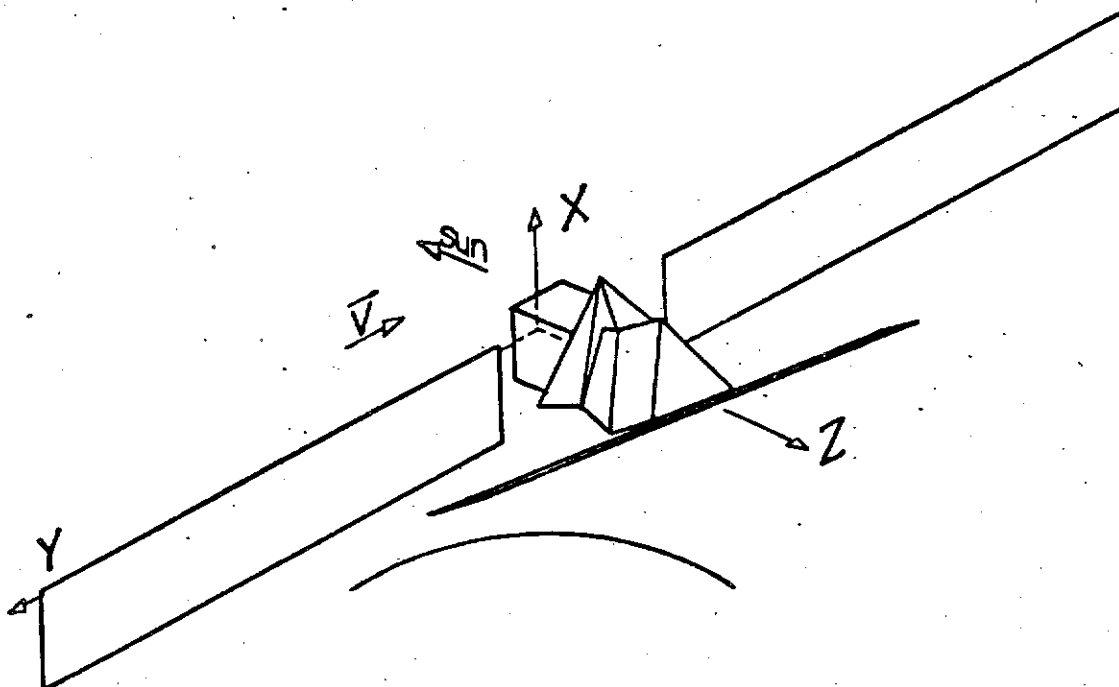


Figure A-1. RADARSAT geometry.

The input files, SATGE.DAT and INELE.DAT, are listed in Tables A-1 and A-2, respectively. The SATGE.DAT is an original NASTRAN file for RADARSAT intended for calculating solar fluxes on external surfaces. Note that the program doesn't use all the spacecraft surfaces defined in the SATGE.DAT file. In fact, as can be seen in the INELE.DAT file, the surface identification numbers 1101, 1041, 1141 and 1161 don't appear in the BODYAP commands, as they are internal surfaces. Note also that it is necessary to define both sides of the solar arrays and the SAR (the 1121 and 1131 surfaces are negatives because its normal was defined as internal in the NASTRAN file). It was assumed that all the surfaces have a normal and tangential aerodynamic coefficients of 0.9 and, with exception of the solar array, a specular reflectivity of 0.8, no diffuse reflection, absorption coefficient equals to 0.2 (no transmitted radiation) and infra-red emissivity equals to 1. The specular and diffuse reflectivity of the solar array were, respectively, 0.21 and 0.0 [4]. The surface temperature was admitted to be equal to 350 K for all the surfaces (solar arrays including). Although they are fixed in the spacecraft, the solar arrays were considered as appendages, for misalignment study purposes.

Table A-1. SATGE.DAT file (obtained from a NASTRAN file).

```
BEGIN BULK
GRID      1000      01400.000-2875.001328.190      0
GRID      1010      0-1400.00-2875.001328.190      0
GRID      1020      0-1400.00-18915.01328.190      0
GRID      1030      01400.000-18915.01328.190      0
GRID      1040      01400.0002875.0001328.190      0
GRID      1050      01400.00018915.001328.190      0
GRID      1060      0-1400.0018915.001328.190      0
GRID      1070      0-1400.002875.0001328.190      0
GRID      1080      0-949.405-7432.445490.793      0
GRID      1090      0174.47787510.4814822.949      0
GRID      1100      0-693.9557510.4813361.504      0
GRID      1110      0-1817.84-7432.444029.347      0
GRID      1120      01050.000875.00001807.190      0
GRID      1130      01050.000875.0000-162.810      0
GRID      1140      0-1050.00875.0000-162.810      0
GRID      1150      0-1050.00875.00001807.190      0
GRID      1160      01050.000-875.0001807.190      0
GRID      1170      0-1050.00-875.0001807.190      0
GRID      1180      01050.000-875.000-162.810      0
GRID      1190      0-1050.00-875.000-162.810      0
GRID      1200      02358.972+0.0E+002057.190      0
GRID      1210      0-413.5072338.6892057.190      0
GRID      1220      0-1050.00875.00002057.190      0
GRID      1230      0-1050.00-875.0002057.190      0
GRID      1240      0-413.507-2338.692057.190      0
GRID      1250      0863.3638+0.0E+002057.190      0
GRID      1260      0-1007.671412.7943548.807      0
GRID      1270      0893.6788-135.8622418.970      0
GRID      1280      0-1209.97-1276.933669.018      0
GRID      1290      0-341.537-1276.935130.464      0
GRID      1300      01762.112-135.8623880.416      0
GRID      1310      0-139.2381412.7945010.252      0
GRID      1320      0-1050.00850.00002057.190      0
GRID      1330      0-1050.00-850.0002057.190      0
```

Table A-1. SATGE.DAT file (cont'd).

CQUAD4	1001	1	1000	1010	1020	1030+0.0E+00	+10001
+10001			+0.0E+00				
CQUAD4	1011	1	1040	1050	1060	1070+0.0E+00	+10002
+10002			+0.0E+00				
CQUAD4	1021	1	1080	1090	1100	1110+0.0E+00	+10003
+10003			+0.0E+00				
CQUAD4	1031	1	1120	1130	1140	1150+0.0E+00	+10004
+10004			+0.0E+00				
CQUAD4	1041	1	1160	1120	1150	1170+0.0E+00	+10005
+10005			+0.0E+00				
CQUAD4	1051	1	1160	1180	1130	1120+0.0E+00	+10006
+10006			+0.0E+00				
CQUAD4	1061	1	1130	1180	1190	1140+0.0E+00	+10007
+10007			+0.0E+00				
CQUAD4	1071	1	1180	1160	1170	1190+0.0E+00	+10008
+10008			+0.0E+00				
CQUAD4	1081	1	1140	1190	1170	1150+0.0E+00	+10009
+10009			+0.0E+00				
CTRIA3	1091	1	1200	1210	1220+0.0E+00		+10010
+10010			+0.0E+00				
CTRIA3	1101	1	1200	1220	1230+0.0E+00		+10011
+10011			+0.0E+00				
CTRIA3	1111	1	1230	1240	1200+0.0E+00		+10012
+10012			+0.0E+00				
CQUAD4	1121	1	1280	1290	1300	1270+0.0E+00	+10013
+10013			+0.0E+00				
CQUAD4	1131	1	1270	1300	1310	1260+0.0E+00	+10014
+10014			+0.0E+00				
CQUAD4	1141	1	1310	1260	1280	1290+0.0E+00	+10015
+10015			+0.0E+00				
CTRIA3	1151	1	1300	1310	1290+0.0E+00		+10016
+10016			+0.0E+00				
CTRIA3	1161	1	1270	1280	1260+0.0E+00		+10017
+10017			+0.0E+00				
CQUAD4	1171	1	1320	1260	1270	1250+0.0E+00	+10018
+10018			+0.0E+00				
CQUAD4	1181	1	1250	1270	1280	1330+0.0E+00	+10019
+10019			+0.0E+00				
CQUAD4	1191	1	1330	1280	1260	1320+0.0E+00	+10020
+10020			+0.0E+00				
PSHELL	1		11.000000		11.000000	1.8333330+0.0E+00	
ENDDATA							

The RADARSAT orbit is circular and sunsynchronous (98.5 degrees inclination) at 777.5 km altitude. An equator crossing time of 18:00 hs was adopted, with resulting right ascension of the ascending node of 11.6 degrees, at 0:00 of 1/1/1993. The RADARSAT attitude is Earth pointing, so that the orbit system best relates to the spacecraft attitude (MAINSYS command equals to 4). Nevertheless, the spacecraft system must be rotated 180 degrees about the X axis, to direct the spacecraft Y axis in the opposite direction of its velocity (the Y direction of the orbit system) as Figure A-1 shows. As there is no relative motion between the spacecraft and the orbital system, the last three components of the MAINSYS card are nulls.

For the APPSYS commands, a frame translated to the point where the solar arrays are articulated during deployment was used so that it would be possible to impose the solar array misalignments at these points. Two cases were analysed: first a nominal geometry and secondly a misaligned solar array geometry. Note that the solar arrays are fixed (i. e. non-rotating) due to the orbit geometry and hence the components of the appendage angular velocity (5th, 6th and 7th parameters of the APPSYS command) are also nulls.

Tables A-3 to A-13 list the output files for the nominal geometry. Table A-3 is the PLASE.DAT file, which contains the coordinates and size of each surface defined in the SATGE.DAT file. Table A-4, EXTSE.DAT, gives the number of each surface defined by the BODYAP commands. The 4 next tables are the aerodynamic and solar radiation forces and torques, as well as the impulses (linear for the forces and angular for the torques) equal to the summation of the product of the forces or torques and the time step.

The input variables in the INELE.DAT file are printed in a common header for the ORBPR, ADRAG and SUNAN files. As the header is strictly the same in all the files, it is shown only once, in Table A-9. ORBPR.DAT contains data from the orbit propagation, as the mean anomaly, spacecraft geocentric coordinates, etc., as indicated in Table A-10. The ADRAG.DAT output file is shown in Table A-11 with the local atmospheric properties and the drag coefficient. The sun-angles are printed in the SUNAN.DAT file, for the main body and each solar array (Table A-12). In this case the angles are the same, as the spacecraft body and the appendages coordinate system have the same orientation. Finally, Table A-13 presents the orbit resume file, ORBMO.DAT, with the maximum and minimum values for the forces, torques and impulses.

Table A-3. PLASE.DAT file.

INT NBR	IDENT NUMBR	EXT. NORMAL VECTOR			ELEMENT POSITION (M)			AREA (M2)
		X	Y	Z	X	Y	Z	
1	1001	0.0000	0.0000	1.0000	.0000E+00	-.1090E+02	.1328E+01	.4491E+02
2	1011	0.0000	0.0000	1.0000	.0000E+00	.1090E+02	.1328E+01	.4491E+02
3	1021	-.8564	.0872	-.5089	-.8217E+00	.3902E-01	.4426E+01	.2550E+02
4	1031	0.0000	1.0000	0.0000	.0000E+00	.8750E+00	.8222E+00	.4137E+01
5	1041	0.0000	0.0000	1.0000	.0000E+00	.0000E+00	.1807E+01	.3675E+01
6	1051	1.0000	0.0000	0.0000	.1050E+01	.0000E+00	.8222E+00	.3448E+01
7	1061	0.0000	0.0000	-1.0000	.0000E+00	.0000E+00	-.1628E+00	.3675E+01
8	1071	0.0000	-1.0000	0.0000	.0000E+00	-.8750E+00	.8222E+00	.4137E+01
9	1081	-1.0000	0.0000	0.0000	-.1050E+01	.0000E+00	.8222E+00	.3448E+01
10	1091	0.0000	0.0000	1.0000	.2985E+00	.1071E+01	.2057E+01	.2773E+01
11	1101	0.0000	0.0000	1.0000	.8632E-01	.0000E+00	.2057E+01	.2983E+01
12	1111	0.0000	0.0000	1.0000	.2985E+00	-.1071E+01	.2057E+01	.2773E+01
13	1121	-.3633	.9063	-.2159	.2761E+00	-.7064E+00	.3775E+01	.4590E+01
14	1131	-.4931	-.8192	.2930	.3772E+00	.6385E+00	.3715E+01	.4590E+01
15	1141	-.8564	.0872	-.5089	-.6746E+00	.6793E-01	.4340E+01	.4590E+01
16	1151	.5108	.0000	.8597	.4271E+00	.6667E-06	.4674E+01	.3157E+01
17	1161	-.5108	.0000	-.8597	-.4413E+00	.6667E-06	.3212E+01	.3157E+01
18	1171	.4752	.8187	-.3224	-.7516E-01	.5317E+00	.2521E+01	.2158E+01
19	1181	.2717	-.8932	-.3582	-.1257E+00	-.5657E+00	.2551E+01	.2169E+01
20	1191	-.9943	.0712	-.0798	-.1079E+01	.3397E-01	.2833E+01	.3423E+01

Table A-5. AEROF.DAT file.

AERODYNAMIC FORCES (N) AND IMPULSES (N*S)
ORBITAL SYSTEM

TIME (S)	FORCES			IMPULSES		
	X	Y	Z	X	Y	Z
0	-.4644E-05	-.1088E-03	-.1518E-04	-.2786E-03	-.6530E-02	-.9109E-03
60	-.4508E-05	-.1052E-03	-.1471E-04	-.5491E-03	-.1284E-01	-.1794E-02
120	-.4345E-05	-.1010E-03	-.1411E-04	-.8098E-03	-.1890E-01	-.2640E-02
180	-.4159E-05	-.9627E-04	-.1341E-04	-.1059E-02	-.2468E-01	-.3445E-02
240	-.3958E-05	-.9126E-04	-.1262E-04	-.1297E-02	-.3015E-01	-.4202E-02
300	-.3746E-05	-.8610E-04	-.1179E-04	-.1522E-02	-.3532E-01	-.4909E-02
360	-.3531E-05	-.8093E-04	-.1093E-04	-.1733E-02	-.4018E-01	-.5565E-02
420	-.3319E-05	-.7593E-04	-.1007E-04	-.1933E-02	-.4473E-01	-.6169E-02
480	-.3119E-05	-.7130E-04	-.9251E-05	-.2120E-02	-.4901E-01	-.6724E-02
540	-.2941E-05	-.6723E-04	-.8496E-05	-.2296E-02	-.5304E-01	-.7234E-02
600	-.2793E-05	-.6390E-04	-.7829E-05	-.2464E-02	-.5688E-01	-.7704E-02
660	-.2679E-05	-.6139E-04	-.7258E-05	-.2624E-02	-.6056E-01	-.8139E-02
720	-.2600E-05	-.5973E-04	-.6780E-05	-.2780E-02	-.6415E-01	-.8546E-02
780	-.2553E-05	-.5884E-04	-.6379E-05	-.2934E-02	-.6768E-01	-.8929E-02
840	-.2534E-05	-.5861E-04	-.6036E-05	-.3086E-02	-.7119E-01	-.9291E-02
900	-.2537E-05	-.5889E-04	-.5725E-05	-.3238E-02	-.7473E-01	-.9634E-02
960	-.2552E-05	-.5950E-04	-.5424E-05	-.3391E-02	-.7830E-01	-.9960E-02
1020	-.2573E-05	-.6026E-04	-.5112E-05	-.3545E-02	-.8191E-01	-.1027E-01
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.
5580	-.4349E-05	-.1042E-03	-.1309E-04	-.1252E-01	-.3159E+00	-.1273E-01
5640	-.4515E-05	-.1079E-03	-.1388E-04	-.1279E-01	-.3224E+00	-.1356E-01
5700	-.4650E-05	-.1109E-03	-.1456E-04	-.1307E-01	-.3290E+00	-.1443E-01
5760	-.4751E-05	-.1130E-03	-.1509E-04	-.1336E-01	-.3358E+00	-.1534E-01
5820	-.4812E-05	-.1142E-03	-.1547E-04	-.1365E-01	-.3427E+00	-.1627E-01
5880	-.4832E-05	-.1143E-03	-.1567E-04	-.1394E-01	-.3495E+00	-.1721E-01
5940	-.4811E-05	-.1134E-03	-.1569E-04	-.1423E-01	-.3563E+00	-.1815E-01
6000	-.4750E-05	-.1116E-03	-.1553E-04	-.1451E-01	-.3630E+00	-.1908E-01
6060	-.4653E-05	-.1088E-03	-.1520E-04	-.1479E-01	-.3695E+00	-.1999E-01
6120	-.4523E-05	-.1053E-03	-.1473E-04	-.1506E-01	-.3759E+00	-.2088E-01

Table A-6. AEROT.DAT file.

TIME (S)	TORQUES			ANGULAR IMPULSES		
	X	Y	Z	X	Y	Z
0	.3471E-04	-.4901E-05	-.8766E-05	.2083E-02	-.2941E-03	-.5260E-03
60	.3375E-04	-.4765E-05	-.8496E-05	.4108E-02	-.5800E-03	-.1036E-02
120	.3256E-04	-.4601E-05	-.8172E-05	.6062E-02	-.8560E-03	-.1526E-02
180	.3121E-04	-.4414E-05	-.7806E-05	.7934E-02	-.1121E-02	-.1994E-02
240	.2973E-04	-.4211E-05	-.7410E-05	.9717E-02	-.1374E-02	-.2439E-02
300	.2816E-04	-.3997E-05	-.6997E-05	.1141E-01	-.1613E-02	-.2859E-02
360	.2657E-04	-.3778E-05	-.6579E-05	.1300E-01	-.1840E-02	-.3254E-02
420	.2499E-04	-.3563E-05	-.6169E-05	.1450E-01	-.2054E-02	-.3624E-02
480	.2351E-04	-.3359E-05	-.5784E-05	.1591E-01	-.2255E-02	-.3971E-02
540	.2218E-04	-.3177E-05	-.5439E-05	.1724E-01	-.2446E-02	-.4297E-02
600	.2108E-04	-.3027E-05	-.5151E-05	.1851E-01	-.2628E-02	-.4606E-02
660	.2024E-04	-.2913E-05	-.4925E-05	.1972E-01	-.2802E-02	-.4902E-02
720	.1967E-04	-.2837E-05	-.4765E-05	.2090E-01	-.2973E-02	-.5188E-02
780	.1935E-04	-.2796E-05	-.4663E-05	.2206E-01	-.3140E-02	-.5467E-02
840	.1925E-04	-.2786E-05	-.4612E-05	.2322E-01	-.3308E-02	-.5744E-02
900	.1933E-04	-.2799E-05	-.4597E-05	.2438E-01	-.3476E-02	-.6020E-02
960	.1953E-04	-.2827E-05	-.4606E-05	.2555E-01	-.3645E-02	-.6296E-02
1020	.1980E-04	-.2862E-05	-.4624E-05	.2674E-01	-.3817E-02	-.6574E-02
.
.
.
5580	.3188E-04	-.4612E-05	-.8157E-05	.1081E+00	-.1408E-01	-.2246E-01
5640	.3321E-04	-.4777E-05	-.8493E-05	.1101E+00	-.1437E-01	-.2297E-01
5700	.3432E-04	-.4912E-05	-.8768E-05	.1122E+00	-.1467E-01	-.2350E-01
5760	.3516E-04	-.5012E-05	-.8972E-05	.1143E+00	-.1497E-01	-.2404E-01
5820	.3572E-04	-.5073E-05	-.9098E-05	.1164E+00	-.1527E-01	-.2458E-01
5880	.3595E-04	-.5093E-05	-.9139E-05	.1186E+00	-.1558E-01	-.2513E-01
5940	.3587E-04	-.5072E-05	-.9096E-05	.1207E+00	-.1588E-01	-.2568E-01
6000	.3548E-04	-.5011E-05	-.8973E-05	.1229E+00	-.1618E-01	-.2622E-01
6060	.3481E-04	-.4914E-05	-.8776E-05	.1250E+00	-.1648E-01	-.2674E-01
6120	.3388E-04	-.4785E-05	-.8515E-05	.1270E+00	-.1676E-01	-.2725E-01

Table A-7. SORAF.DAT file

RADIATION FORCES (N) AND IMPULSES (N*S)
ORBIT PLANE SYSTEM

TIME (S)	FORCES			IMPULSES		
	X	Y	Z	X	Y	Z
0	-.4066E-04	.9045E-04	-.5276E-03	-.2440E-02	.5427E-02	-.3165E-01
60	-.3886E-04	.8786E-04	-.5260E-03	-.4771E-02	.1070E-01	-.6322E-01
120	-.3693E-04	.8547E-04	-.5245E-03	-.6987E-02	.1583E-01	-.9469E-01
180	-.3492E-04	.8328E-04	-.5231E-03	-.9082E-02	.2082E-01	-.1261E+00
240	-.3283E-04	.8131E-04	-.5216E-03	-.1105E-01	.2570E-01	-.1574E+00
300	-.3070E-04	.7954E-04	-.5203E-03	-.1289E-01	.3047E-01	-.1886E+00
360	-.2854E-04	.7798E-04	-.5189E-03	-.1461E-01	.3515E-01	-.2197E+00
420	-.2638E-04	.7662E-04	-.5177E-03	-.1619E-01	.3975E-01	-.2508E+00
480	-.2423E-04	.7544E-04	-.5165E-03	-.1764E-01	.4428E-01	-.2818E+00
540	-.2212E-04	.7446E-04	-.5153E-03	-.1897E-01	.4874E-01	-.3127E+00
600	-.2005E-04	.7364E-04	-.5142E-03	-.2017E-01	.5316E-01	-.3435E+00
660	-.1804E-04	.7298E-04	-.5132E-03	-.2126E-01	.5754E-01	-.3743E+00
720	-.1611E-04	.7247E-04	-.5123E-03	-.2222E-01	.6189E-01	-.4051E+00
780	-.1426E-04	.7208E-04	-.5114E-03	-.2308E-01	.6621E-01	-.4358E+00
840	-.1250E-04	.7182E-04	-.5105E-03	-.2383E-01	.7052E-01	-.4664E+00
900	-.1084E-04	.7165E-04	-.5098E-03	-.2448E-01	.7482E-01	-.4970E+00
960	-.9283E-05	.7157E-04	-.5091E-03	-.2503E-01	.7912E-01	-.5275E+00
1020	-.7823E-05	.7155E-04	-.5086E-03	-.2550E-01	.8341E-01	-.5580E+00
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5580	-.4956E-04	.1159E-03	-.5405E-03	.1660E-01	.6091E+00	-.2999E+01
5640	-.4918E-04	.1120E-03	-.5388E-03	.1364E-01	.6159E+00	-.3031E+01
5700	-.4854E-04	.1082E-03	-.5370E-03	.1073E-01	.6224E+00	-.3064E+01
5760	-.4767E-04	.1046E-03	-.5353E-03	.7872E-02	.6286E+00	-.3096E+01
5820	-.4657E-04	.1012E-03	-.5335E-03	.5078E-02	.6347E+00	-.3128E+01
5880	-.4529E-04	.9788E-04	-.5318E-03	.2361E-02	.6406E+00	-.3160E+01
5940	-.4382E-04	.9478E-04	-.5301E-03	-.2687E-03	.6463E+00	-.3191E+01
6000	-.4221E-04	.9186E-04	-.5284E-03	-.2801E-02	.6518E+00	-.3223E+01
6060	-.4046E-04	.8915E-04	-.5268E-03	-.5229E-02	.6571E+00	-.3255E+01
6120	-.3860E-04	.8665E-04	-.5253E-03	-.7545E-02	.6623E+00	-.3286E+01

Table A-8. SORAT.DAT file

TIME (S)	TORQUES			ANGULAR IMPULSES		
	X	Y	Z	X	Y	Z
0	-.8945E-04	-.4507E-04	.1891E-05	-.5367E-02	-.2704E-02	.1135E-03
60	-.8944E-04	-.4812E-04	.1860E-05	-.1073E-01	-.5592E-02	.2251E-03
120	-.8908E-04	-.5128E-04	.1817E-05	-.1608E-01	-.8668E-02	.3341E-03
180	-.8838E-04	-.5452E-04	.1761E-05	-.2138E-01	-.1194E-01	.4398E-03
240	-.8732E-04	-.5783E-04	.1695E-05	-.2662E-01	-.1541E-01	.5415E-03
300	-.8592E-04	-.6120E-04	.1620E-05	-.3178E-01	-.1908E-01	.6387E-03
360	-.8417E-04	-.6460E-04	.1538E-05	-.3683E-01	-.2296E-01	.7310E-03
420	-.8208E-04	-.6801E-04	.1448E-05	-.4175E-01	-.2704E-01	.8178E-03
480	-.7966E-04	-.7142E-04	.1352E-05	-.4653E-01	-.3132E-01	.8990E-03
540	-.7690E-04	-.7479E-04	.1251E-05	-.5114E-01	-.3581E-01	.9740E-03
600	-.7384E-04	-.7812E-04	.1146E-05	-.5557E-01	-.4050E-01	.1043E-02
660	-.7046E-04	-.8136E-04	.1038E-05	-.5980E-01	-.4538E-01	.1105E-02
720	-.6680E-04	-.8450E-04	.9272E-06	-.6381E-01	-.5045E-01	.1161E-02
780	-.6286E-04	-.8753E-04	.8139E-06	-.6758E-01	-.5570E-01	.1210E-02
840	-.5866E-04	-.9040E-04	.6988E-06	-.7110E-01	-.6113E-01	.1252E-02
900	-.5422E-04	-.9312E-04	.5825E-06	-.7435E-01	-.6671E-01	.1286E-02
960	-.4958E-04	-.9564E-04	.4682E-06	-.7733E-01	-.7245E-01	.1315E-02
1020	-.4475E-04	-.9797E-04	.3533E-06	-.8001E-01	-.7833E-01	.1336E-02
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5580	-.7939E-04	-.2531E-04	.1593E-05	-.9500E-03	-.3089E+00	-.1191E-01
5640	-.8170E-04	-.2757E-04	.1688E-05	-.5852E-02	-.3105E+00	-.1181E-01
5700	-.8373E-04	-.2995E-04	.1768E-05	-.1088E-01	-.3123E+00	-.1170E-01
5760	-.8547E-04	-.3244E-04	.1833E-05	-.1600E-01	-.3143E+00	-.1159E-01
5820	-.8690E-04	-.3505E-04	.1878E-05	-.2122E-01	-.3164E+00	-.1148E-01
5880	-.8803E-04	-.3777E-04	.1904E-05	-.2650E-01	-.3186E+00	-.1137E-01
5940	-.8883E-04	-.4060E-04	.1910E-05	-.3183E-01	-.3211E+00	-.1125E-01
6000	-.8930E-04	-.4351E-04	.1901E-05	-.3719E-01	-.3237E+00	-.1114E-01
6060	-.8946E-04	-.4651E-04	.1877E-05	-.4255E-01	-.3265E+00	-.1103E-01
6120	-.8928E-04	-.4961E-04	.1840E-05	-.4791E-01	-.3294E+00	-.1092E-01

Table A-9. Header of the ORBPR.DAT, ADRAG.DAT
and SUNAN.DAT files.

ORBITAL ELEMENTS:			
SEMI MAJOR AXIS (M)	7155660.000		
ECCENTRICITY	0.00000000		
INCLINATION (DEG)	98.500		
ASCENDING NODE (DEG)	11.600		
PERIGEE ARG. (DEG)	0.000		
MEAN ANOMALY (DEG)	0.000		
ORBITAL PERIOD (MIN)	100.400		
START TIME:			
MONTH	1		
DAY	1		
YEAR	1993		
MODIFIED JULIAN DATE	48988.5		
HOUR	0		
MINUTES	0		
SECONDS	0.000		
TIME STEP (SECONDS)	60.000		
CALCULATION TIME (S)	6120.000		
SOLAR FLUX DATA:			
SOLAR FLUX F10.7	180.00		
AVERAGED SOLAR FLUX	180.00		
GEOMAG. ACTIVITY KP	5.00		
SATELLITE CENTER OF MASS IN BODY FRAME COORDINATES:			
X AXIS (METERS)	0.00000		
Y AXIS (METERS)	0.00000		
Z AXIS (METERS)	2.56000		
REFERENCE AREA FOR DRAG COEFFICIENT:			
SATELLITE AREA (M2)	11.90700		
ATTITUDE RELATED FRAME:			
FRAME NUMBER	4 - SPACECRAFT SYSTEM		
X-Y-Z ROTATION (DEG)	180.00000	0.00000	0.00000
ANGULAR VELOCITY (RPM)	0.00000	0.00000	0.00000
APPENDAGE NUMBER 1			
APPENDAGE POSITION (M)	0.00000	-.87500	1.32819
ANGULAR VELOCITY (RPM)	0.00000	0.00000	0.00000
X-Y-Z ROTATION (DEG)	0.00000	0.00000	0.00000
APPENDAGE NUMBER 2			
APPENDAGE POSITION (M)	0.00000	.87500	1.32819
ANGULAR VELOCITY (RPM)	0.00000	0.00000	0.00000
X-Y-Z ROTATION (DEG)	0.00000	0.00000	0.00000

Table A-10. ORBPR.DAT file.

ORBIT TIME (S)	MEAN ANOM. (DEG)	GEOCENTR. DISTANCE (KM)	SATEL. VELOC. (KM/S)	ORBIT DATA		GEOCENTR. ALTITUDE (KM)	SUN R. ASCEN. (DEG)	SUN DECLI. (DEG)
				LONGI- TUDE (DEG)	LATI- TUDE (DEG)			
0	0.00	7155.660	7.464	270.95	0.00	777.500	281.59	-23.01
60	3.58	7155.660	7.464	270.17	3.54	777.582	281.59	-23.01
120	7.17	7155.660	7.464	269.38	7.08	777.827	281.59	-23.01
180	10.75	7155.660	7.464	268.59	10.62	778.231	281.59	-23.01
240	14.33	7155.660	7.464	267.79	14.17	778.787	281.59	-23.01
300	17.92	7155.660	7.464	266.96	17.70	779.487	281.59	-23.01
360	21.50	7155.660	7.464	266.12	21.24	780.319	281.60	-23.01
420	25.08	7155.660	7.464	265.24	24.78	781.271	281.60	-23.01
480	28.67	7155.660	7.464	264.33	28.31	782.328	281.60	-23.01
540	32.25	7155.660	7.464	263.37	31.84	783.472	281.60	-23.01
600	35.83	7155.660	7.464	262.36	35.36	784.687	281.60	-23.01
660	39.42	7155.660	7.464	261.28	38.88	785.952	281.60	-23.01
720	43.00	7155.660	7.464	260.11	42.39	787.248	281.60	-23.01
780	46.59	7155.660	7.464	258.83	45.90	788.555	281.60	-23.01
840	50.17	7155.660	7.464	257.41	49.39	789.852	281.60	-23.01
900	53.75	7155.660	7.464	255.81	52.87	791.119	281.60	-23.01
960	57.34	7155.660	7.464	253.98	56.34	792.337	281.60	-23.01
1020	60.92	7155.660	7.464	251.83	59.77	793.486	281.60	-23.01
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5580	333.26	7155.660	7.464	251.99	-25.61	781.807	281.66	-23.01
5640	336.85	7155.660	7.464	251.10	-23.07	780.799	281.66	-23.01
5700	340.43	7155.660	7.464	250.24	-19.54	779.903	281.66	-23.01
5760	344.02	7155.660	7.464	249.40	-16.00	779.132	281.66	-23.01
5820	347.60	7155.660	7.464	248.59	-12.46	778.500	281.66	-23.01
5880	351.18	7155.660	7.464	247.79	-8.92	778.017	281.66	-23.01
5940	354.77	7155.660	7.464	247.00	-5.38	777.689	281.66	-23.01
6000	358.35	7155.660	7.464	246.22	-1.84	777.522	281.67	-23.01
6060	1.93	7155.660	7.464	245.44	1.71	777.519	281.67	-23.01
6120	5.52	7155.660	7.464	244.66	5.25	777.680	281.67	-23.01

Table A-11. ADRAG.DAT file.

ORBIT TIME (S)	ATM. TEMP (K)	LOCAL DENSITY (KG/M3)	SPEED RATIO	ATMOSPHERIC DATA		DRAG COMPONENTS			DRAG COEFF.
				ANG. ALPHA (DEG)	ANG. BETA (DEG)	X	Y	Z	
0	1161	.762E-13	5.92	-.00	-176.1	-.179	4.202	.586	4.232
60	1161	.732E-13	5.84	-.00	-176.1	-.181	4.224	.591	4.255
120	1161	.699E-13	5.76	-.00	-176.1	-.183	4.247	.593	4.277
180	1161	.663E-13	5.67	.00	-176.2	-.184	4.268	.594	4.299
240	1161	.626E-13	5.58	.00	-176.2	-.186	4.288	.593	4.318
300	1161	.588E-13	5.51	-.00	-176.3	-.187	4.306	.590	4.335
360	1161	.551E-13	5.44	0.00	-176.4	-.188	4.320	.583	4.348
420	1161	.516E-13	5.38	-.00	-176.5	-.189	4.329	.574	4.357
480	1161	.484E-13	5.34	.00	-176.6	-.190	4.334	.562	4.360
540	1161	.457E-13	5.31	.00	-176.7	-.190	4.332	.547	4.357
600	1161	.435E-13	5.29	.00	-176.8	-.189	4.326	.530	4.349
660	1161	.419E-13	5.28	.00	-177.0	-.188	4.315	.510	4.336
720	1161	.409E-13	5.28	.00	-177.1	-.187	4.301	.488	4.320
780	1161	.405E-13	5.28	.00	-177.3	-.186	4.284	.465	4.301
840	1161	.405E-13	5.29	-.00	-177.5	-.184	4.266	.439	4.281
900	1161	.409E-13	5.30	0.00	-177.7	-.183	4.249	.413	4.262
960	1161	.415E-13	5.30	.00	-177.9	-.182	4.232	.386	4.244
.
1020	1161	.421E-13	5.30	.00	-178.1	-.180	4.219	.358	4.228
5580	1161	.759E-13	6.36	.00	-176.5	-.169	4.037	.507	4.060
5640	1161	.782E-13	6.32	-.00	-176.4	-.170	4.060	.522	4.084
5700	1161	.799E-13	6.27	.00	-176.3	-.171	4.082	.536	4.108
5760	1161	.810E-13	6.22	0.00	-176.2	-.173	4.103	.548	4.131
5820	1161	.814E-13	6.17	.00	-176.2	-.174	4.125	.559	4.153
5880	1161	.811E-13	6.11	-.00	-176.1	-.175	4.146	.568	4.175
5940	1161	.800E-13	6.04	-.00	-176.1	-.177	4.168	.576	4.198
6000	1161	.783E-13	5.96	.00	-176.1	-.178	4.190	.583	4.220
6060	1161	.759E-13	5.88	0.00	-176.1	-.180	4.213	.589	4.244
6120	1161	.731E-13	5.80	-.00	-176.1	-.182	4.237	.592	4.267

Table A-12. SUNAN.DAT file.

ORBIT TIME (S)	SPACECRAFT		SUN ASPECT ANGLE (DEGREES)							
	MAIN	BODY	APPENDAGE NUMBER 1		APPENDAGE NUMBER 2		APPENDAGE NUMBER 3		APPENDAGE NUMBER 4	
			ALPHA	BETA	ALPHA	BETA	ALPHA	BETA	ALPHA	BETA
0	-0.0	-75.5	-0.0	-75.5	-0.0	-75.5				
60	-0.9	-75.5	-0.9	-75.5	-0.9	-75.5				
120	-1.8	-75.6	-1.8	-75.6	-1.8	-75.6				
180	-2.7	-75.7	-2.7	-75.7	-2.7	-75.7				
240	-3.6	-75.9	-3.6	-75.9	-3.6	-75.9				
300	-4.4	-76.2	-4.4	-76.2	-4.4	-76.2				
360	-5.3	-76.5	-5.3	-76.5	-5.3	-76.5				
420	-6.1	-76.8	-6.1	-76.8	-6.1	-76.8				
480	-6.9	-77.2	-6.9	-77.2	-6.9	-77.2				
540	-7.7	-77.7	-7.7	-77.7	-7.7	-77.7				
600	-8.4	-78.2	-8.4	-78.2	-8.4	-78.2				
660	-9.2	-78.7	-9.2	-78.7	-9.2	-78.7				
720	-9.8	-79.3	-9.8	-79.3	-9.8	-79.3				
780	-10.5	-79.9	-10.5	-79.9	-10.5	-79.9				
840	-11.1	-80.6	-11.1	-80.6	-11.1	-80.6				
900	-11.7	-81.3	-11.7	-81.3	-11.7	-81.3				
960	-12.2	-82.0	-12.2	-82.0	-12.2	-82.0				
1020	-12.6	-82.8	-12.6	-82.8	-12.6	-82.8				
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5580	6.5	-77.0	6.5	-77.0	6.5	-77.0				
5640	5.7	-76.6	5.7	-76.6	5.7	-76.6				
5700	4.9	-76.3	4.9	-76.3	4.9	-76.3				
5760	4.0	-76.0	4.0	-76.0	4.0	-76.0				
5820	3.1	-75.8	3.1	-75.8	3.1	-75.8				
5880	2.2	-75.7	2.2	-75.7	2.2	-75.7				
5940	1.4	-75.6	1.4	-75.6	1.4	-75.6				
6000	.5	-75.5	.5	-75.5	.5	-75.5				
6060	-.4	-75.5	-.4	-75.5	-.4	-75.5				
6120	-1.3	-75.6	-1.3	-75.6	-1.3	-75.6				

Table A-13. ORBMO.DAT file.

ORBIT NUMBER 1 T = 6060 S

AERODYNAMIC FORCES AND TORQUES:

ORBITAL SYSTEM SPACECRAFT SYSTEM

AERODYNAMIC FORCE		TORQUE	
MINIMUM	.294183E-04 N	.119103E-04	N*M
MAXIMUM	.115465E-03 N	.374445E-04	N*M
X MEAN	-.244060E-05 N	.206209E-04	N*M
Y MEAN	-.609816E-04 N	-.271885E-05	N*M
Z MEAN	-.329892E-05 N	-.441317E-05	N*M

AEROD. IMPULSE LINEAR		ANGULAR	
X MINIMUM	-.147900E-01 N*S	.000000E+00	N*M*S
X MAXIMUM	.000000E+00 N*S	.124962E+00	N*M*S
X DIFF.	.147900E-01 N*S	.124962E+00	N*M*S
X BUILD-UP	-.147900E-01 N*S	.124962E+00	N*M*S
X ACCUM.	-.147900E-01 N*S	.124962E+00	N*M*S
Y MINIMUM	-.369548E+00 N*S	-.164762E-01	N*M*S
Y MAXIMUM	.000000E+00 N*S	.000000E+00	N*M*S
Y DIFF.	.369548E+00 N*S	.164762E-01	N*M*S
Y BUILD-UP	-.369548E+00 N*S	-.164762E-01	N*M*S
Y ACCUM.	-.369548E+00 N*S	-.164762E-01	N*M*S
Z MINIMUM	-.199915E-01 N*S	-.267438E-01	N*M*S
Z MAXIMUM	.000000E+00 N*S	.000000E+00	N*M*S
Z DIFF.	.199915E-01 N*S	.267438E-01	N*M*S
Z BUILD-UP	-.199915E-01 N*S	-.267438E-01	N*M*S
Z ACCUM.	-.199915E-01 N*S	-.267438E-01	N*M*S

Table A-13. ORBMO.DAT file (cont'd).

SOLAR RADIATION FORCES AND TORQUES:
 ORBIT PLANE SYSTEM SPACECRAFT SYSTEM

SOLAR RADIATION		FORCE	TORQUE	
MINIMUM	.511612E-03	N	.115285E-04	N*M
MAXIMUM	.585884E-03	N	.107771E-03	N*M
X MEAN	-.862856E-06	N	-.702226E-05	N*M
Y MEAN	.108437E-03	N	-.538721E-04	N*M
Z MEAN	-.537071E-03	N	-.181943E-05	N*M
S. RAD. IMPULSE		LINEAR	ANGULAR	
X MINIMUM	-.267705E-01	N*S	-.901591E-01	N*M*S
X MAXIMUM	.560858E-01	N*S	.526704E-01	N*M*S
X DIFF.	.828563E-01	N*S	.142830E+00	N*M*S
X BUILD-UP	-.522891E-02	N*S	-.425549E-01	N*M*S
X ACCUM.	-.522891E-02	N*S	-.425549E-01	N*M*S
Y MINIMUM	.000000E+00	N*S	-.326465E+00	N*M*S
Y MAXIMUM	.657130E+00	N*S	.000000E+00	N*M*S
Y DIFF.	.657130E+00	N*S	.326465E+00	N*M*S
Y BUILD-UP	.657130E+00	N*S	-.326465E+00	N*M*S
Y ACCUM.	.657130E+00	N*S	-.326465E+00	N*M*S
Z MINIMUM	-.325465E+01	N*S	-.124062E-01	N*M*S
Z MAXIMUM	.000000E+00	N*S	.135757E-02	N*M*S
Z DIFF.	.325465E+01	N*S	.137637E-01	N*M*S
Z BUILD-UP	-.325465E+01	N*S	-.110257E-01	N*M*S
Z ACCUM.	-.325465E+01	N*S	-.110257E-01	N*M*S

Except the spacecraft's geometry files PLASE.DAT, EXTSE.DAT and the orbit resume output ORBMO.DAT all the other files are difficult to analyse, due to the large amount of printed results. Therefore, as they are in a table in function of the time, it is preferable to plot those values. They are shown in Figures A-2 to A-11.

Figure A-2 shows the atmospheric density as printed in the ADRAG.DAT file. The dynamic model and the satellite polar orbit are responsible for the high variation in the density with time. The drag coefficient suffers the same effect, as can be seen in Figure A-3. The high value of the drag coefficient is caused by the large solar arrays, which contribute to a significant drag but have no projected area in the velocity direction.

Figures A-4 and A-5 are plots of the aerodynamic and solar radiation forces, respectively. Note that the forces aren't in the same coordinate system. The aerodynamic are weaker than the solar radiation forces, basically due to the orbit geometry: the sun radiation impinges on the entire solar arrays while the atmospheric molecules do not. The orbit inclination and the date chosen for the simulation are responsible for the lack of an eclipse period during the orbit, and thus the non-pulsating shape of the solar radiation forces (and torques) that are typical of these type of perturbations.

The torques, shown in Figures A-6 and A-7 show a similar behavior as the forces, with the aerodynamic and solar radiation effects causing a torque principally about the spacecraft's X and Y axes, respectively. Both are due to the SAR positioning.

Figures A-8 and A-9 show the linear impulse during the propagation period. Again, the solar radiation is the most important effect, giving an impulse of almost 3.3 Ns in the orbit normal direction per orbit period. The angular impulses (Figures A-10 and A-11), present high values in the X and Y axis, spacecraft's yaw and roll axis respectively.

ATMOSPHERIC DENSITY
DYNAMIC MODEL

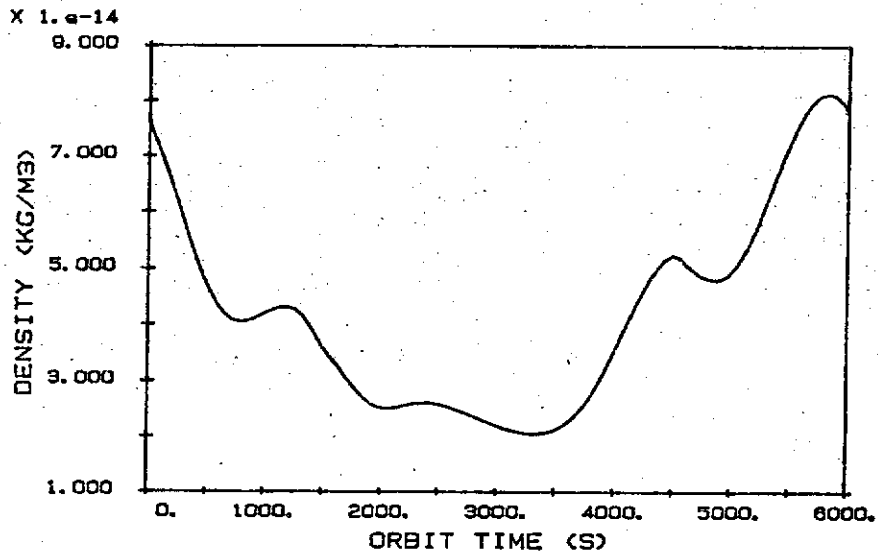


Fig. A-2. Local atmospheric density for the RADARSAT orbit.

RADARSAT AERODYNAMIC DRAG
WITH NO MISALIGNMENT

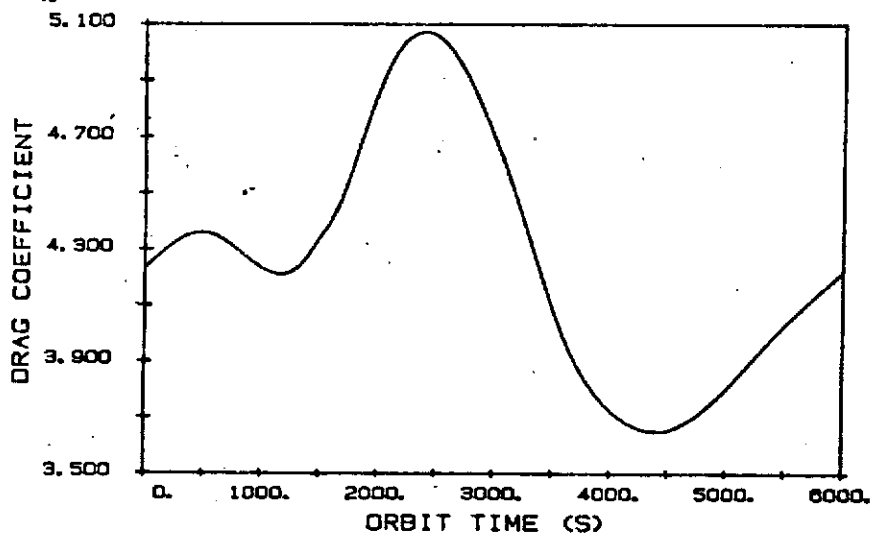


Fig A-3. RADARSAT drag coefficient variation during the orbit.

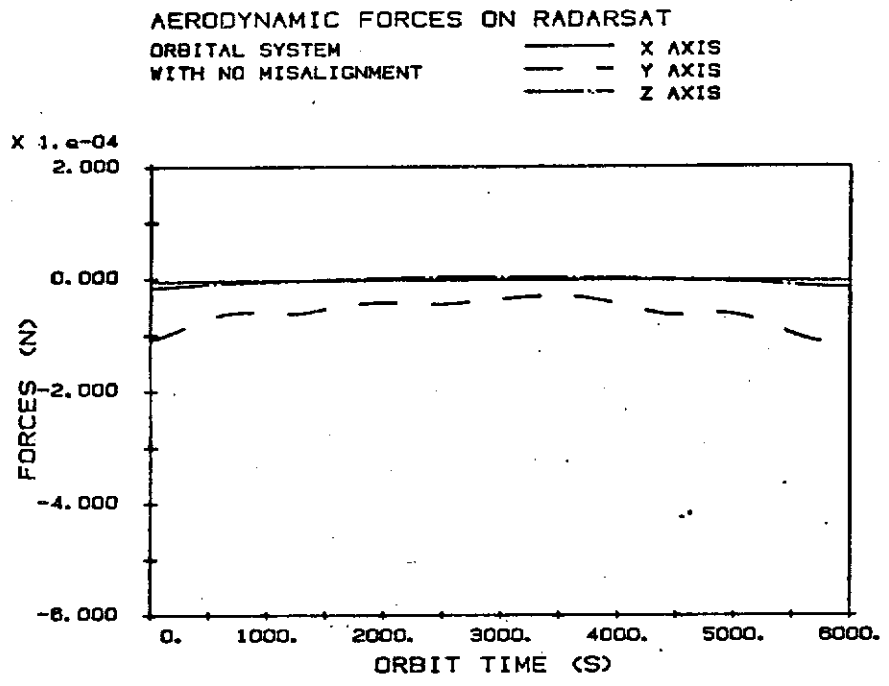


Fig. A-4. Aerodynamic forces on RADARSAT without misalignment.

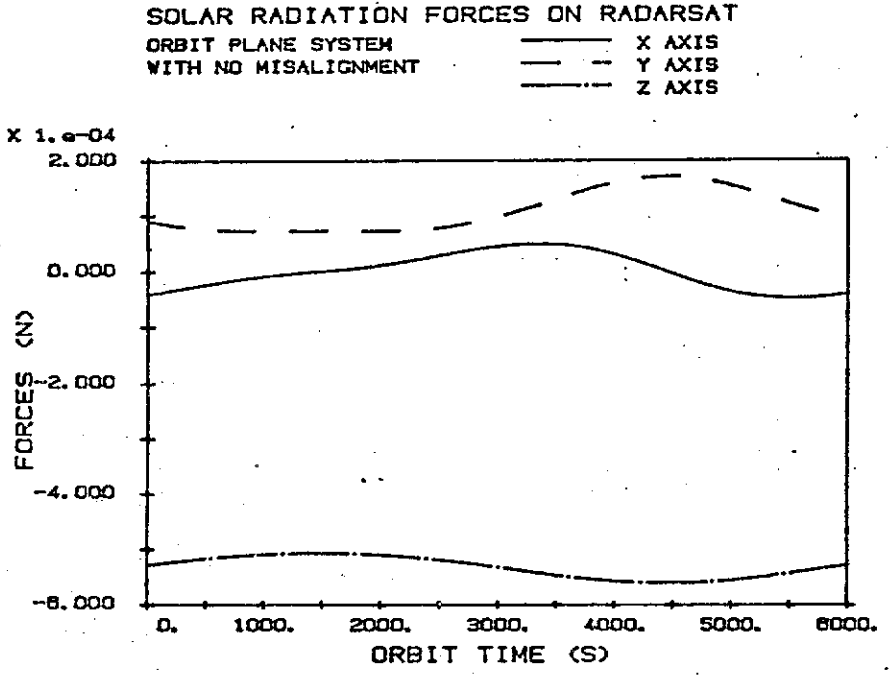


Fig A-5. Solar radiation forces on RADARSAT without misalignment.

AERODYNAMIC TORQUES ON RADARSAT
 SPACECRAFT SYSTEM WITH NO MISALIGNMENT

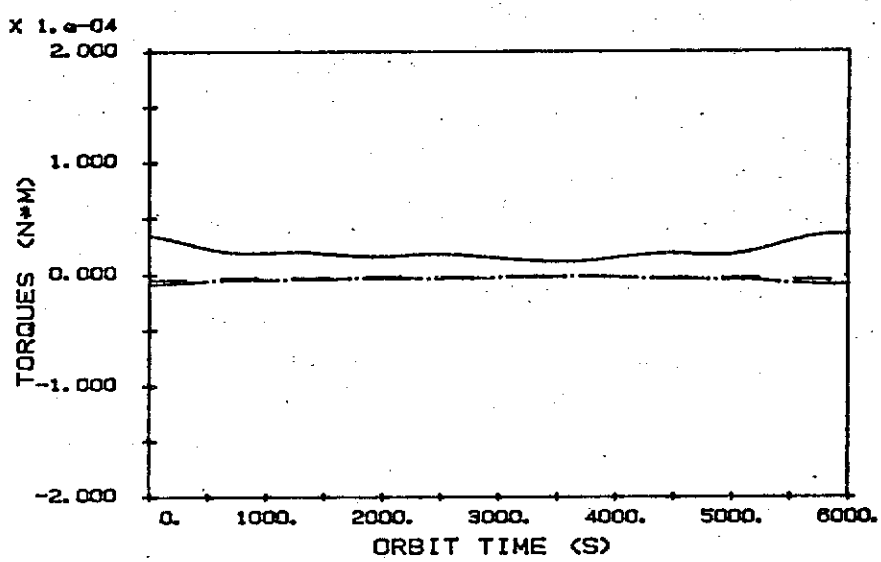


Fig. A-6. Aerodynamic torques on RADARSAT without misalignment.

SOLAR RADIATION TORQUES ON RADARSAT
 SPACECRAFT SYSTEM WITH NO MISALIGNMENT

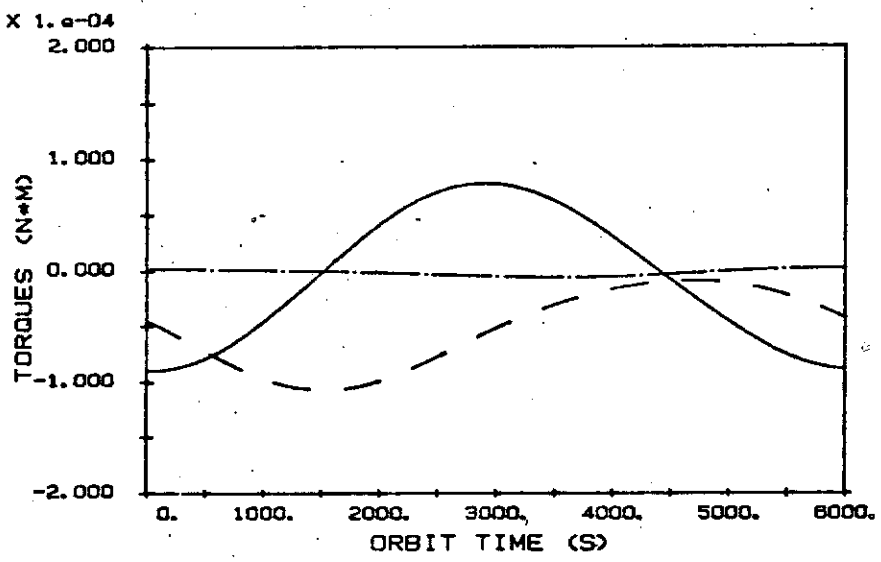


Fig A-7. Solar radiation torques on RADARSAT without misalignment.

AERODYNAMIC LINEAR IMPULSE ON RADARSAT
ORBITAL SYSTEM WITH NO MISALIGNMENT

— X AXIS
- - Y AXIS
- · - Z AXIS

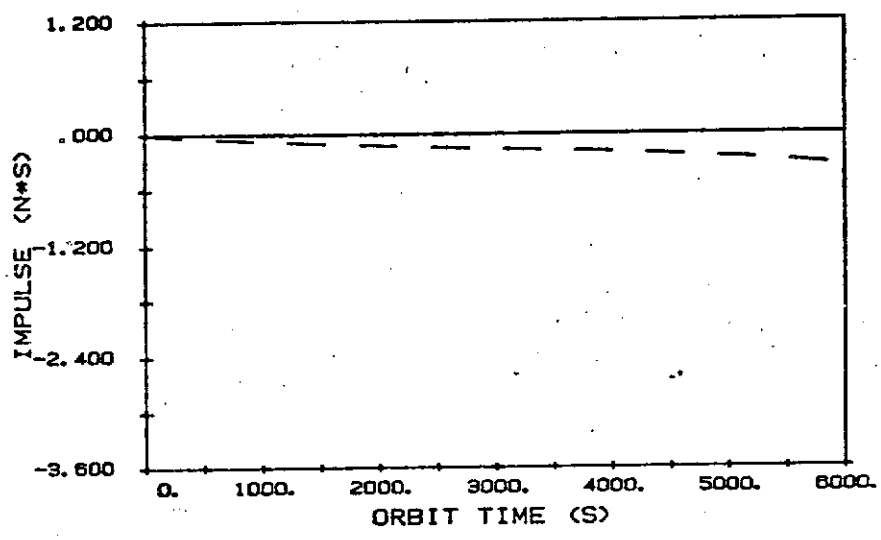


Fig. A-8. Aerodynamic linear impulse variation.

SOLAR RADIATION LINEAR IMPULSE ON RADARSAT
ORBIT PLANE SYSTEM WITH NO MISALIGNMENT

— X AXIS
- - Y AXIS
- · - Z AXIS

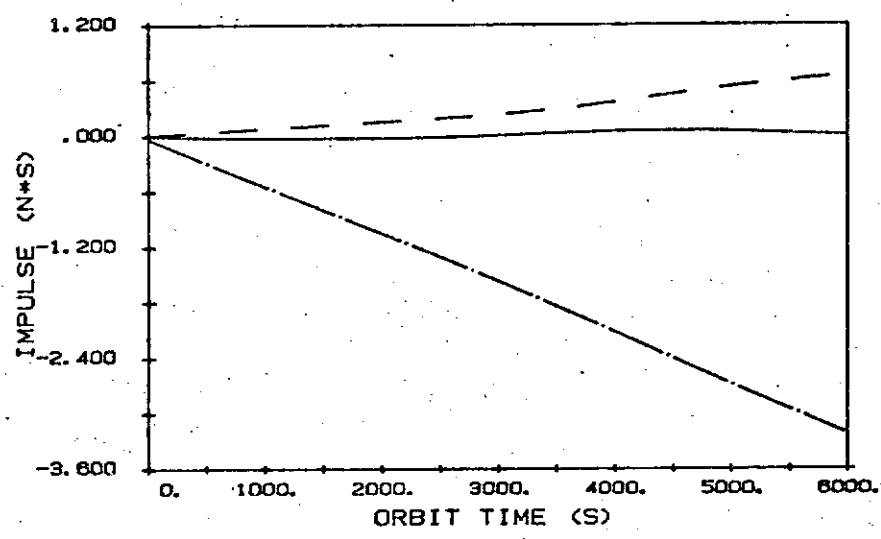


Fig A-9. Solar radiation linear impulse variation.

AERODYNAMIC ANGULAR IMPULSE ON RADARSAT

SPACECRAFT SYSTEM WITH NO MISALIGNMENT

— X AXIS
 - - - Y AXIS
 - · - · Z AXIS

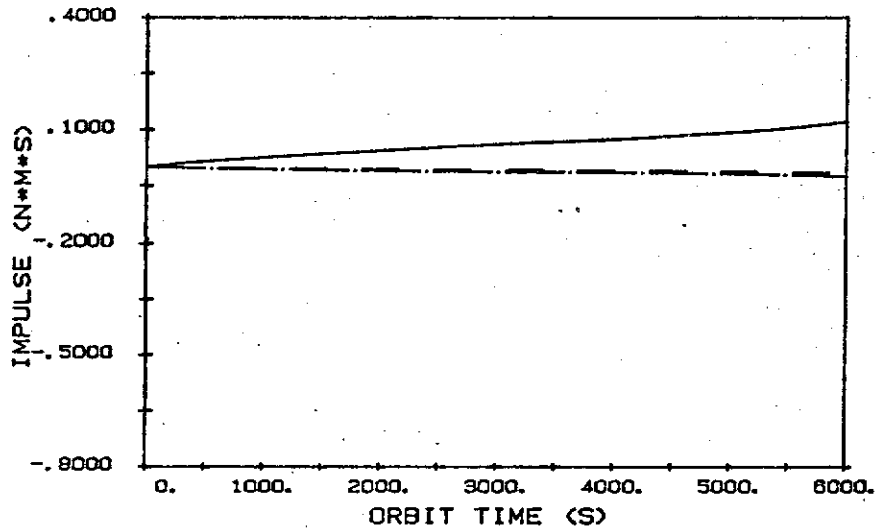


Fig. A-10. Aerodynamic angular impulse variation during a single orbit.

SOLAR RADIATION ANGULAR IMPULSE ON RADARSAT

SPACECRAFT SYSTEM WITH NO MISALIGNMENT

— X AXIS
 - - - Y AXIS
 - · - · Z AXIS

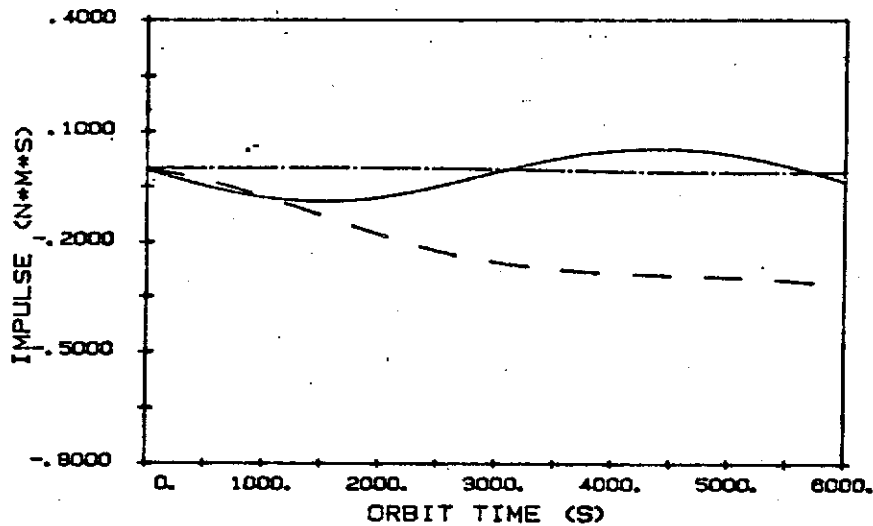


Fig A-11. Solar radiation angular impulse variation during a single orbit.

It was assumed, based on ref. [4], that for each solar array the misalignment angles could be in the range: 1.2 deg, 0.1 deg and 0.2 deg in the spacecraft X, Y and Z direction, respectively. These angles can be visualized (exaggerated) in Figure A-12. The RADARSAT center of mass offsets were considered equal to 0.1 m maximum for all three axes.

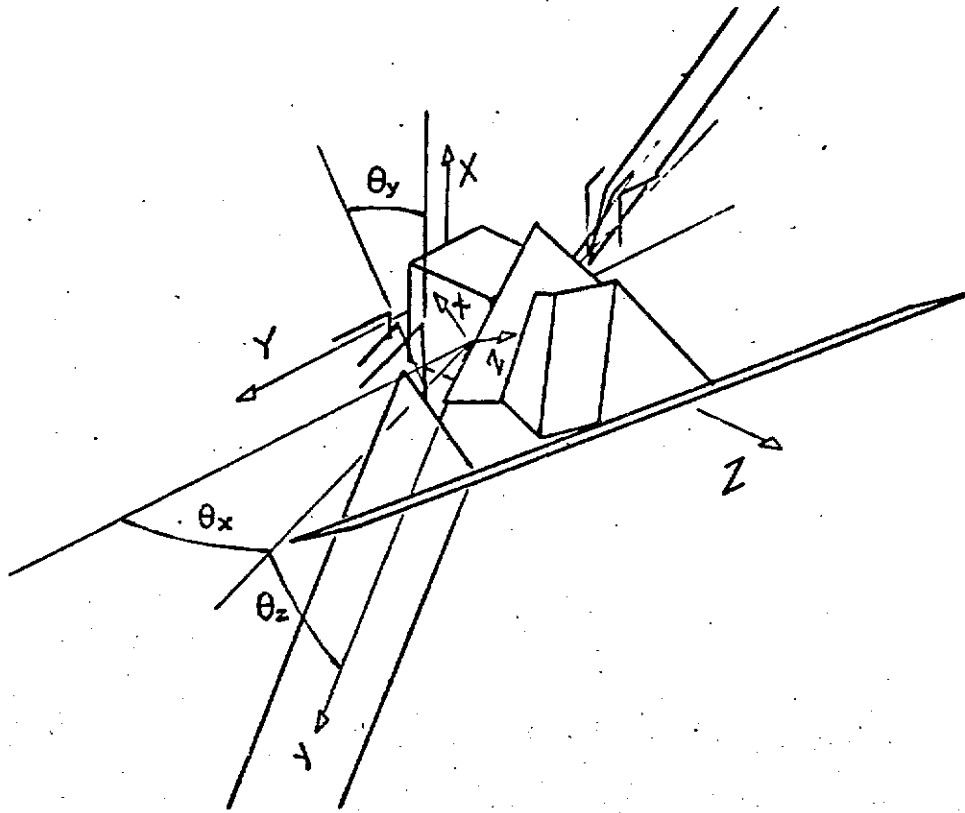


Figure A-12. RADARSAT misalignment angles (largely exaggerated).

To compare the results with and without misalignment, the angles were chosen in the possible range for the worst case, in order to maximize the resulting build-up angular impulse. As the torques and, consequently, the impulses are almost linear for small array angles, it was necessary to test only the maximum values. Table A-14 shows the altered cards of INELE.DAT file which caused the highest angular impulse build-up for the RADARSAT dawn-dusk configuration.

Table A-14. INELE.DAT cards with the misaligned angles and the center of mass offset.

BARICEN	-0.1	-0.1	2.46							
APPESYS	1	0.	-0.875	1.328190	0.	0.	0.	1.2	0.1	0.2
APPESYS	2	0.	0.875	1.328190	0.	0.	0.	-1.2	0.1	-0.2

The resulting torques and impulses due to the misalignments are shown in Figures A-13, A-14 (aerodynamic and solar radiation torques), A-15 and A-16 (aerodynamic and solar radiation angular impulses).

By comparing these results with Figures A-6 to A-9, it is seen that the misalignments cause a greater change in the solar radiation torque than in the aerodynamic. In fact, the solar radiation torque in the spacecraft's Y axis (roll) is strongly dependent on the center of mass position in the X axis. The center of mass offset in the X axis alone raises the angular build-up impulse from .33 to .65 Nms in the Y axis. The other misalignments are responsible for the increase in this value to .78 Nms, or 94% of the total solar radiation angular momentum build-up and 84% of the total angular momentum build-up (aerodynamic and solar radiation effects together). A change in the negative X axis direction of 0.1 m position of the spacecraft's center of mass causes an increasing of 35% in the angular momentum build-up.

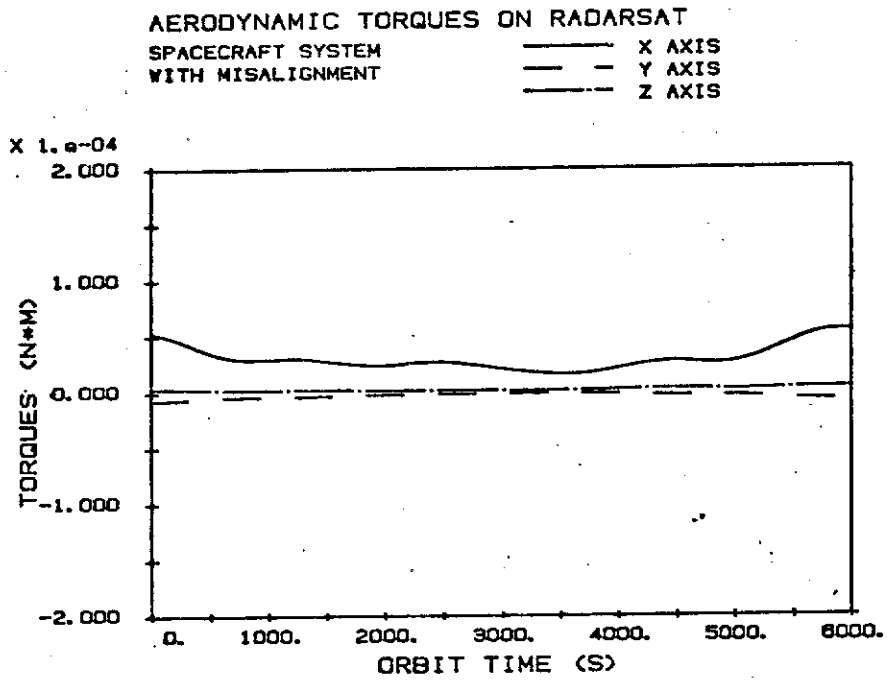


Fig. A-13. Aerodynamic torque on RADARSAT with misalignment included.

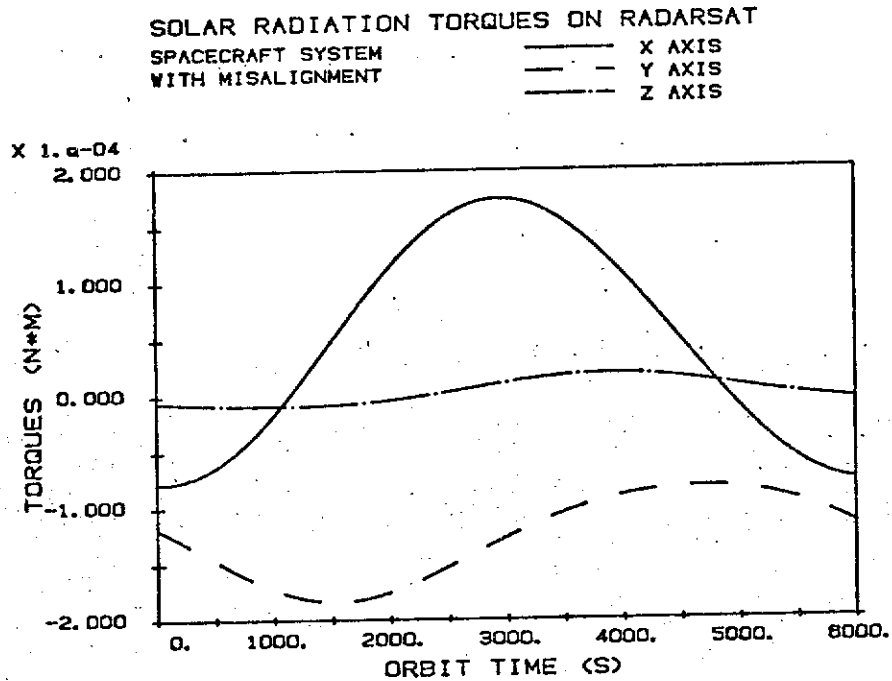


Fig. A-14. Solar radiation torques on RADARSAT with misalignments included.

AERODYNAMIC ANGULAR IMPULSE ON RADARSAT
SPACECRAFT SYSTEM WITH MISALIGNMENT

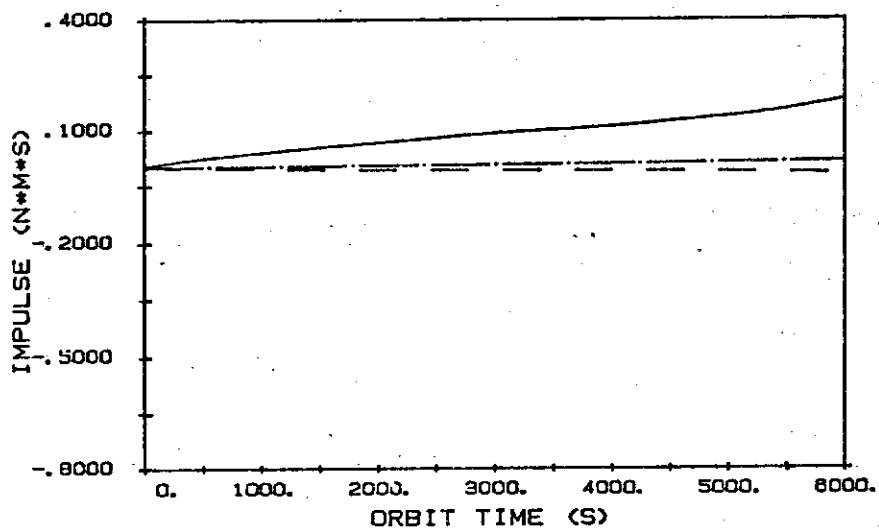


Fig. A-15. Aerodynamic angular impulse build-up with misalignments included.

SOLAR RADIATION ANGULAR IMPULSE ON RADARSAT
SPACECRAFT SYSTEM WITH MISALIGNMENT

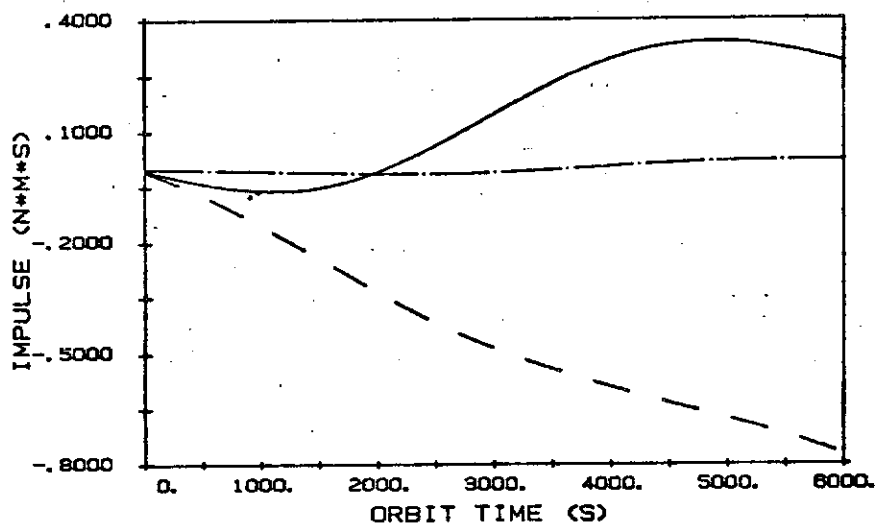


Fig. A-16. Solar radiation angular impulse build-up with misalignments included.

APPENDIX B
MAIN PROGRAM AND SUBROUTINES LISTING.

C SUNAN.DAT CONTAINS PRINTED VALUES OF THE SUN ASPECT
C
C ANGLE, DEFINED AS THE ANGLE BETWEEN THE
C SUN DIRECTION IN THE SPACECRAFT SYSTEM
C (FOR THE MAIN BODY) OR THE APPENDAGE
C SYSTEMS (FOR THE APPENDAGES), AND THE
C PLANE YZ (ALPHA). THE ANGLE BETA IS THE
C ANGLE BETWEEN THE PROJECTION OF THE SUN
C DIRECTION IN THE PLANE YZ AND THE Y AXIS.
C
C ORBMO.DAT CONTAINS ORBIT EVALUATIONS OF THE FORCES
C (MINIMUM AND MAXIMUM INTENSITIES AND
C MEAN VALUES) AND THE IMPULSES (MINIMUM,
C AND MAXIMUM VALUES ENCOUNTERED AT THAT
C ORBIT, ITS DIFFERENCE, THE IMPULSE BUILD-
C UP - DIFFERENCE BETWEEN THE ACCUMULATED
C IMPULSE AT THE END AND THE BEGINING OF
C THE ORBIT - AND THE ACCUMULATED IMPULSE).
C
C SUBROUTINES UTILISED BY THIS PROGRAM:
C RCPAR - HP 1000 INTERNAL ROUTINE.
C CONSTP - SETS CONSTANT VALUES (ORBIT.FTN FILE).
C ARELEM - READS THE SATGE.DAT FILE
C REAINP - READS SOME INELE.DAT FILE CARDS.
C DJM - MODIFIED JULIAN DATE (ORBIT.FTN FILE).
C VARELK - TIME VARIATIONS OF THE KEPLERIAN ELEMENTS
C (ORBIT.FTN FILE).
C ROTSYS - OBTAINS SOME ROTATION MATRIX.
C ROTXYZ - OBTAINS THE MATRIX FOR A XYZ ROTATION.
C UNIVEC - CALCULATES THE UNIT OF A VECTOR.
C STAVEC - CALCULATES THE POSITION AND VELOCITY OF A
C SATELLITE BY ITS KEPLERIAN ELEMENTS.
C ROTMAT - OBTAINS THE ROTATION MATRIX OF A ROTATION
C ABOUT A GIVEN AXIS.
C MATMUL - OBTAINS THE PRODUCT OF TWO MATRIX.
C SUN - SUN POSITION (ORBIT.FTN FILE).
C VECROT - OBTAINS THE COMPONENTS OF A VECTOR IN
C OTHER SYSTEM.
C EARSHA - VERIFIES IF THE SATELLITE IS IN THE EARTH
C SHADOW.
C SGMT - OBTAINS THE GREENWICH SIDERAL TIME AT OH
C (ORBIT.FTN FILE).
C GEOCEN - CALCULATES THE GEOCENTRIC COORDINATES.
C ASDAMO - OBTAINS THE ATMOSPHERIC PROPERTIES.
C ATMVEL - CALCULATES THE ATMOSPHERIC VELOCITY.
C AERODY - COMPUTES THE AERODYNAMIC FORCE AND TORQUE
C IN A FLAT PLATE.
C SOLPRE - COMPUTES THE SOLAR RADIATION FORCE AND
C TORQUE IN A FLAT PLATE.
C VEGINV - OBTAINS THE INVERSE ROTATION OF A VECTOR.
C

C
C
C
C

AUTHOR: VALDEMIR CARRARA MAY/87

IMPLICIT REAL*8 (A-H,O-Z)

PARAMETER (NSUR=50)

DIMENSION SU(2)

DIMENSION AH(3),AP(3),AS(3),AV(3),AW(3)

DIMENSION BA(3),BR(3),DA(3),DR(3),EA(3),EB(3)

DIMENSION FA(3),FR(3),F1(3),F2(3),F5(3),F6(3)

DIMENSION GA(3),GC(3),HA(3),HR(3),PV(3)

DIMENSION RG(3),RH(3),RV(3),SF(3),SP(3)

DIMENSION TA(3),TR(3),T1(3),T2(3),T5(3),T6(3)

DIMENSION US(3),UV(3),VN(3),WA(3),WU(3),WV(3)

DIMENSION AD(6),CA(6),CR(6),DE(6),EL(6),ER(6)

DIMENSION SA(6),SC(7),SV(6),TE(6)

DIMENSION AB(10),AM(40),AA(20),AL(50),SM(40),WM(10)

DIMENSION RA(3,3),RC(3,3),RI(3,3),RL(3,3),RN(3,3)

DIMENSION RP(3,3),RS(3,3),RT(3,3),RZ(3,3)

DIMENSION R1(3,3),R2(3,3)

DIMENSION PL(NSUR,3),PP(10,3),SL(NSUR,7),VL(NSUR,3)

DIMENSION WS(10,3),QR(10,3,3)

DIMENSION NF(11)

CHARACTER *20 FD(6),OD(6)

CHARACTER *10 FN(14)

CHARACTER *8 CQUE

CHARACTER RFIL*3,FILI*14

COMMON /CONSTA/ PI,PIT2,PID2,RAD,DEG

COMMON /CEARTH/ RE,GM,FLAT,TETP

COMMON /ANGSUN/ ALSU,DESU,ROSU,ELSU,ECLI

COMMON /CTSSUN/ ASTU,DSUN

DATA BOAV /8314.3298D0/

DATA SUNC /1.011D+17/

DATA AMAX /1.D+32/

DATA AMIN /1.D-32/

	DATA FD/	INERTIAL SYSTEM	,	ASC. NODE SYSTEM	,
1		ORBIT PLANE SYSTEM	,	SPACECRAFT SYSTEM	,
2			,		/
	DATA OD/	INERTIAL SYSTEM	,	ASC. NODE SYSTEM	,
1		ORBIT PLANE SYSTEM	,	ORBITAL SYSTEM	,
2		SPACECRAF SYS AT T=0	,	SPACECRAFT SYSTEM	/

C
C
C

OPEN FILES

```
FN(1) = 'SATGE.DAT'  
FN(2) = 'INELE.DAT'  
FN(3) = '  
FN(4) = 'PLASE'  
FN(5) = 'EXTSE'  
FN(6) = 'AEROF'  
FN(7) = 'AEROT'  
FN(8) = 'SORAF'  
FN(9) = 'SORAT'  
FN(10) = 'ORBPR'  
FN(11) = 'ADRAG'  
FN(12) = 'SUNAN'  
FN(13) = 'ORBMO'  
FN(14) = '
```

```
FILI = FN(1)  
CALL RCPAR(1I,FILI)  
OPEN(UNIT=3,FILE=FILI)
```

```
FILI = FN(2)  
CALL RCPAR(2I,FILI)  
OPEN(UNIT=4,FILE=FILI)
```

```
RFIL = 'DAT'
```

```
CALL RCPAR(3I,RFIL)
```

```
DO 10 I = 4, 13  
FILI = FN(I)//'.'//RFIL  
NFIL = I + 2  
OPEN(UNIT=NFIL,FILE=FILI)  
10 CONTINUE
```

C
C
C

COMMON VALUES INITIALIZATION

```
CALL CONSTP
```

C
C
C
C

```
READ THE SURFACE ELEMENTS AND ITS PROPERTIES FROM  
SATGE.DAT FILE
```

```
CALL ARELEM(NF,VL,PL,SL,AL)
```

C
C
C
READ THE INPUT ELEMENTS FROM INELE.DAT FILE

```
CALL REAINP('ORBIT ',4,6,EL)
PERI = EL(1)
APOG = EL(2)
SEMA = (EL(1) + EL(2))* .5D+03 + RE
EXCE = (EL(2) - EL(1))* .5E+03/SEMA
EL(1) = SEMA
EL(2) = DABS(EXCE)
PERI = PIT2*DSQRT(EL(1)/GM)*EL(1)
CALL REAINP('DATE ',4,8,AB)
AMO = AB(1)
DAY = AB(2)
YEA = AB(3)
HOU = AB(4)
AMI = AB(5)
SEC = AB(6)
DTI = AB(7)
TEN = AB(8)
CALL REAINP('SOLFLUX ',4,3,SF)
CALL REAINP('BARICEN ',4,3,GC)
CALL REAINP('MAINSYS ',4,7,AB)
IFRA = AB(1)
IF(IFRA.LT.1.OR.IFRA.GT.4) GOTO 920
```

```
DO 20 I = 1, 3
EB(I) = AB(I+1)
WV(I) = AB(I+4)
20 CONTINUE
```

```
CALL REAINP('PROAREA ',4,1,AB)
SBA = AB(1)
CALL REAINP('OUTSYS ',4,4,AB)
IOFA = AB(1)
IOTA = AB(2)
IOFR = AB(3)
IOTR = AB(4)
IF(IOFA.LT.1.OR.IOFA.GT.6) IOFA = 1
IF(IOTA.LT.1.OR.IOTA.GT.6) IOTA = 1
IF(IOFR.LT.1.OR.IOFR.GT.6) IOFR = 1
IF(IOTR.LT.1.OR.IOTR.GT.6) IOTR = 1
```

```
DJUL = DJM(YEA,AMO,DAY)
YFRA = (DJUL - DJM(YEA,1.DO,1.DO))/3.6525D02
DFRA = SEC + 60.DO*(AMI + 60:DO*HOU)
SBAS = SBA
```

C
C
C
C

WRITE HEADER OF THE ORBPR.DAT, ADRAG.DAT AND
SUNAN.DAT FILES

```
DO 30 I = 12, 14
WRITE(I,1000) EL,PERI/60.DO
WRITE(I,1001) AMO,DAY,YEA,DJUL,HOU,AMI,SEC,DTI,TEN
WRITE(I,1002) SF
WRITE(I,1003) GC
WRITE(I,1004) SBA
WRITE(I,1005) IFRA,FD(IFRA),EB,WV
30 CONTINUE
```

C
C
C

READ THE APPSYS COMMANDS FROM INELE.DAT FILE

```
NBOD = 0
NELE = 0

40 CONTINUE
REWIND 4
NBOD = NBOD + 1
IF(NBOD.GT.10) GOTO 910
IF(NF(NBOD+1).EQ.0) GOTO 80
NELE = NELE + NF(NBOD)

50 CONTINUE
READ(4,*,END=950) CQUE
IF(CQUE.EQ.'ENDDATA ') GOTO 950
IF(CQUE.NE.'APPESYS ') GOTO 50
BACKSPACE 4
READ(4,*) CQUE,INAP,AP,AW,EA
IF(INAP.NE.NBOD) GOTO 50

WRITE(12,1006) INAP,AP,AW,EA
WRITE(13,1006) INAP,AP,AW,EA
WRITE(14,1006) INAP,AP,AW,EA

DO 60 I = 1, 3
EA(I) = EA(I)*RAD
AW(I) = AW(I)*PIT2/60.DO
60 CONTINUE

CALL UNIVEC(AW,WA,WAPP)
CALL ROTXYZ(EA,R1)
WM(NBOD) = WAPP
```

```
DO 70 I = 1, 3
WS(NBOD,I) = WA(I)
PP(NBOD,I) = AP(I)
QR(NBOD,I,1) = R1(I,1)
QR(NBOD,I,2) = R1(I,2)
QR(NBOD,I,3) = R1(I,3)
70 CONTINUE

DO 78 I = 1, NF(NBOD+1)
INAP = NELE + I
DO 72 J = 1, 3
PL(INAP,J) = PL(INAP,J) - PP(NBOD,J)
72 CONTINUE

78 CONTINUE

GOTO 40

80 CONTINUE
NBAP = NBOD

WRITE( 8,3000) OD(IOFA)
WRITE(10,4000) OD(IOFR)
WRITE( 9,5000) OD(IOTA)
WRITE(11,6000) OD(IOTR)
WRITE(12,7000)
WRITE(13,8000)
WRITE(14,9000)
```

C
C
C

ORBIT AND ATTITUDE CONSTANTS

```
AM(1) = AMAX
AM(2) = AMAX
AM(3) = AMIN
AM(4) = AMIN
SM(1) = AMAX
SM(2) = AMAX
SM(3) = AMIN
SM(4) = AMIN

DO 85 I=5, 40
AM(I) = 0.DO
SM(I) = 0.DO
85 CONTINUE
```

```
EL(3) = EL(3)*RAD
EL(4) = EL(4)*RAD
EL(5) = EL(5)*RAD
EL(6) = EL(6)*RAD
EB(1) = EB(1)*RAD
EB(2) = EB(2)*RAD
EB(3) = EB(3)*RAD
WV(1) = WV(1)*PI/30.DO
WV(2) = WV(2)*PI/30.DO
WV(3) = WV(3)*PI/30.DO
```

```
CALL VARELK(EL,DE)
```

```
ER(1) = EL(1)
ER(2) = EL(2)
ER(3) = EL(3)
```

```
CALL ROTSYS(1,EL,RI)
CALL ROTSYS(2,EL,RN)
CALL ROTSYS(3,EL,RL)
```

```
IF(IFRA.NE.4) CALL ROTSYS(IFRA,EL,RS)
```

```
CALL ROTXYZ(EB,RA)
CALL UNIVVEC(WV,WU,WBOD)
```

```
ICON = 0
NORB = 0
TIST = 0.DO
```

```
DO 90 I = 1, 3
HA(I) = 0.DO
HR(I) = 0.DO
AH(I) = 0.DO
BA(I) = 0.DO
BR(I) = 0.DO
90 CONTINUE
```


C
C
C

ORBIT AND ATTITUDE PROPAGATION

```
100 CONTINUE
TIME   = ICON*DTI
ICON   = ICON + 1
DTIM   = DFRA + TIME
IF(TIME.GT.TEN) GOTO 900

ER(4)  = DMOD(EL(4)+DE(4)*TIME,PIT2)
ER(5)  = DMOD(EL(5)+DE(5)*TIME,PIT2)
ER(6)  = DMOD(EL(6)+DE(6)*TIME,PIT2)

CALL STAVEC(ER,SV,RP)

IF(IFRA.NE.4) GOTO 120
DO 110 I=1,3
DO 110 J=1,3
RS(I;J) = RP(I,J)
110 CONTINUE
120 CONTINUE

ANGL   = DMOD(WBOD*TIME,PIT2)
CALL ROTMAT(ANGL,WU,R1)
CALL MATMUL(RA,RS,RZ)
CALL MATMUL(R1,RZ,RT)
```

C
C
C

SUN POSITION

```
CALL SUN(DJUL,DTIM,SP)
SU(1)  = ALSU
SU(2)  = DESU
PRAD   = SUNC/ROSU/ROSU

CALL UNIVVEC(SP,RV,DIST)
CALL VECROT(RT,RV,US)

AA(1)  = DASIN (US(1))*DEG
AA(2)  = DATAN2(US(3),US(2))*DEG

CALL EARSHA(SV,SP,SHAD)
SUPR   = PRAD*SHAD
```

C
C
C

ATMOSPHERIC PROPERTY COMPUTATIONS

GWST = SGMT(DJUL,DTIM)

CALL GEOCEN(SV,SA)

CALL ASDAMO(SA,SU,SF,DJUL,DTIM,GWST,TE,AD,WMOL,RHOD)

CALL ATMVEL(SV,AV)

UV(1) = SV(4) - AV(1)

UV(2) = SV(5) - AV(2)

UV(3) = SV(6) - AV(3)

CALL UNIVVEC(UV,RV,VREL)

CALL VECROT(RT,RV,UV)

TEMP = TE(2)

SPRT = VREL/DSQRT(2.DO*BOAV*TEMP/WMOL)

AEPR = VREL*VREL*RHOD*.5DO

C
C
C

SPACECRAFT MAIN BODY LOOP

DO 140 J=1,3

FA(J) = 0.DO

TA(J) = 0.DO

FR(J) = 0.DO

TR(J) = 0.DO

140 CONTINUE

DO 170 I=1,NF(1)

DO 150 J = 1, 3

VN(J) = VL(I,J)

PV(J) = PL(I,J)

SC(J) = SL(I,J)

SC(J+3) = SL(I,J+3)

RG(J) = PV(J) - GC(J)

150 CONTINUE

SC(7) = SL(I,7)

AREA = AL(I)

TWTI = SC(7)/TEMP

C
C
C
AERODYNAMIC FORCE AND TORQUE ON A PLANE ELEMENT

CALL AERODY(AEPR,UV,VN,RG,SPRT,TWTI,SC,AREA,F1,T1)

C
C
C
SOLAR RADIATION FORCE AND TORQUE COMPUTATION

CALL SOLPRE(SUPR,US,VN,RG,SC,AREA,F5,T5)

DO 160 J=1,3

FA(J) = FA(J) + F1(J)

TA(J) = TA(J) + T1(J)

FR(J) = FR(J) + F5(J)

TR(J) = TR(J) + T5(J)

160 CONTINUE

170 CONTINUE

C
C
C
APPENDAGE LOOP

NBOD = 0

NELE = 0

190 CONTINUE

NBOD = NBOD + 1

IF(NF(NBOD+1).EQ.0) GOTO 400

NELE = NELE + NF(NBOD)

WAPP = WM(NBOD)

DO 200 J = 1, 3

AW(J) = WS(NBOD,J)

F1(J) = GC(J) - PP(NBOD,J)

R1(J,1) = QR(NBOD,J,1)

R1(J,2) = QR(NBOD,J,2)

R1(J,3) = QR(NBOD,J,3)

F2(J) = 0.DO

T2(J) = 0.DO

F6(J) = 0.DO

T6(J) = 0.DO

200 CONTINUE

```
ANGL = DMOD(WAPP*TIME,PIT2)
CALL ROTMAT(ANGL,AW,R2)
CALL MATMUL(R1,R2,RC)
CALL VECROT(RC,UV,RV)
CALL VECROT(RC,US,AS)
CALL VECROT(RC,F1,GA)
```

```
AA(2*NBOD+1) = DASIN(AS(1))*DEG
AA(2*NBOD+2) = DATAN2(AS(3),AS(2))*DEG
```

```
DO 300 I=1,NF(NBOD+1)
```

```
INAP = NELE + I
```

```
DO 210 J=1,3
```

```
VN(J) = VL(INAP,J)
```

```
PV(J) = PL(INAP,J)
```

```
SC(J) = SL(INAP,J)
```

```
SC(J+3) = SL(INAP,J+3)
```

```
RG(J) = PV(J) - GA(J)
```

```
210 CONTINUE
```

```
SC(7) = SL(INAP,7)
```

```
AREA = AL(INAP)
```

```
TWTI = SC(7)/TEMP
```

```
C
C
C      AERODYNAMIC FORCE AND TORQUE
```

```
CALL AERODY(AEPR,RV,VN,RG,SPRT,TWTI,SC,AREA,F1,T1)
```

```
C
C
C      SOLAR RADIATION FORCE AND TORQUE
```

```
CALL SOLPRE(SUPR,AS,VN,RG,SC,AREA,F5,T5)
```

```
DO 220 J=1,3
```

```
F2(J) = F2(J) + F1(J)
```

```
T2(J) = T2(J) + T1(J)
```

```
F6(J) = F6(J) + F5(J)
```

```
T6(J) = T6(J) + T5(J)
```

```
220 CONTINUE
```

```
300 CONTINUE
```

C
C
C
MAIN BODY + APPENDAGE

```
CALL VECINV(RC,F2,F1)
CALL VECINV(RC,T2,T1)
CALL VECINV(RC,F6,F5)
CALL VECINV(RC,T6,T5)
```

```
DO 310 J=1,3
FA(J) = FA(J) + F1(J)
TA(J) = TA(J) + T1(J)
FR(J) = FR(J) + F5(J)
TR(J) = TR(J) + T5(J)
```

```
310 CONTINUE
```

```
GOTO 190
```

```
400 CONTINUE
```

C
C
C
ROTATE THE FORCE AND TORQUE TO THE OUTPUT SYSTEM

```
CALL VECINV(RT,FA,F1)
CALL VECINV(RT,TA,T1)
CALL VECINV(RT,FR,F5)
CALL VECINV(RT,TR,T5)
```

```
IF(IOFA.EQ.1) CALL VECROT(RI,F1,F2)
IF(IOFA.EQ.2) CALL VECROT(RN,F1,F2)
IF(IOFA.EQ.3) CALL VECROT(RL,F1,F2)
IF(IOFA.EQ.4) CALL VECROT(RP,F1,F2)
IF(IOFA.EQ.5) CALL VECROT(RZ,F1,F2)
IF(IOFA.EQ.6) CALL VECROT(RT,F1,F2)
```

```
IF(IOFR.EQ.1) CALL VECROT(RI,F5,F6)
IF(IOFR.EQ.2) CALL VECROT(RN,F5,F6)
IF(IOFR.EQ.3) CALL VECROT(RL,F5,F6)
IF(IOFR.EQ.4) CALL VECROT(RP,F5,F6)
IF(IOFR.EQ.5) CALL VECROT(RZ,F5,F6)
IF(IOFR.EQ.6) CALL VECROT(RT,F5,F6)
```

```
IF(IOTA.EQ.1) CALL VECROT(RI,T1,T2)
IF(IOTA.EQ.2) CALL VECROT(RN,T1,T2)
IF(IOTA.EQ.3) CALL VECROT(RL,T1,T2)
IF(IOTA.EQ.4) CALL VECROT(RP,T1,T2)
IF(IOTA.EQ.5) CALL VECROT(RZ,T1,T2)
IF(IOTA.EQ.6) CALL VECROT(RT,T1,T2)
```

```
IF(IOTR.EQ.1) CALL VECROT(RI,T5,T6)
IF(IOTR.EQ.2) CALL VECROT(RN,T5,T6)
IF(IOTR.EQ.3) CALL VECROT(RL,T5,T6)
IF(IOTR.EQ.4) CALL VECROT(RP,T5,T6)
IF(IOTR.EQ.5) CALL VECROT(RZ,T5,T6)
IF(IOTR.EQ.6) CALL VECROT(RT,T5,T6)
```

C
C
C

LINEAR AND ANGULAR MOMENTUM COMPUTATIONS

```
DO 450 I = 1, 3
HA(I) = HA(I) + F2(I)*DTI
HR(I) = HR(I) + F6(I)*DTI
AH(I) = AH(I) + T2(I)*DTI
RH(I) = RH(I) + T6(I)*DTI
J      = 10*I
IF(HA(I).LT.AM(J+1)) AM(J+1) = HA(I)
IF(HA(I).GT.AM(J+3)) AM(J+3) = HA(I)
IF(AH(I).LT.AM(J+2)) AM(J+2) = AH(I)
IF(AH(I).GT.AM(J+4)) AM(J+4) = AH(I)
IF(HR(I).LT.SM(J+1)) SM(J+1) = HR(I)
IF(HR(I).GT.SM(J+3)) SM(J+3) = HR(I)
IF(RH(I).LT.SM(J+2)) SM(J+2) = RH(I)
IF(RH(I).GT.SM(J+4)) SM(J+4) = RH(I)
450 CONTINUE
```

C
C
C

AEROF, AEROT, SORAF AND SORAT TABLES

```
WRITE(1,1120) TIME
WRITE( 8,1320) TIME,F2,HA
WRITE(10,1320) TIME,F6,HR
WRITE( 9,1320) TIME,T2,AH
WRITE(11,1320) TIME,T6,RH
```

C
C
C

ORBPR.DAT FILE RESULTS

```

ANME      = ER(6)*DEG
VLCT      = DSQRT(SV(4)*SV(4)+SV(5)*SV(5)+SV(6)*SV(6))
VLCT      = VLCT*1.D-03
RAIO      = DSQRT(SV(1)*SV(1)+SV(2)*SV(2)+SV(3)*SV(3))
RAIO      = RAIO*1.D-03
ALON      = DMOD(SA(1)-GWST,PIT2)*DEG
IF(ALON.LT.0.DO) ALON      = ALON + 360.DO
ALAT      = SA(2)*DEG
ALTU      = SA(3)/1.D+03
SURA     = DMOD(SU(1),PIT2)*DEG
SUDL      = SU(2)*DEG
WRITE(12,1720) TIME,ANME,RAIO,VLCT,
1          ALON,ALAT,ALTU,SURA,SUDL

```

C
C
C

ADRAG.DAT FILE RESULTS

```

ALPH      = DASIN(UV(1))*DEG
BETA      = DMOD(DATAN2(UV(3),UV(2)),PIT2)*DEG
F1(1)     = FA(1)/AEPR/SBAS
F1(2)     = FA(2)/AEPR/SBAS
F1(3)     = FA(3)/AEPR/SBAS
CDAR      = FA(1)*UV(1)+FA(2)*UV(2)+FA(3)*UV(3)
CDAR      = -CDAR/AEPR/SBAS
WRITE(13,1820) TIME,TEMP,RHOD,SPRT,ALPH,BETA,F1,CDAR

```

C
C
C

SUNAN.DAT FILE RESULTS

```

WRITE(14,9020) TIME,(AA(I),I=1,2*NBAP)

```

C
C
C

ORBMO.DAT FILE RESULTS

```

FMOD      = DSQRT(F2(1)*F2(1)+F2(2)*F2(2)+F2(3)*F2(3))
TMOD      = DSQRT(T2(1)*T2(1)+T2(2)*T2(2)+T2(3)*T2(3))
IF(FMOD.LT.AM(1)) AM(1)      = FMOD
IF(FMOD.GT.AM(3)) AM(3)      = FMOD
IF(TMOD.LT.AM(2)) AM(2)      = TMOD
IF(TMOD.GT.AM(4)) AM(4)      = TMOD

```

```

FMOD  = DSQRT(F6(1)*F6(1)+F6(2)*F6(2)+F6(3)*F6(3))
TMOD  = DSQRT(T6(1)*T6(1)+T6(2)*T6(2)+T6(3)*T6(3))
IF(FMOD.LT.SM(1)) SM(1)  = FMOD
IF(FMOD.GT.SM(3)) SM(3)  = FMOD
IF(TMOD.LT.SM(2)) SM(2)  = TMOD
IF(TMOD.GT.SM(4)) SM(4)  = TMOD

```

```

DPER  = TIME - TIST

```

```

IF(DPER.LT.PERI) GOTO 880

```

```

TIST  = TIME

```

```

NORB  = NORB + 1

```

```

DO 850 I = 1, 3

```

```

AM(3+2*I) = (HA(I) - DA(I))/DPER

```

```

AM(4+2*I) = (AH(I) - BA(I))/DPER

```

```

SM(3+2*I) = (HR(I) - DR(I))/DPER

```

```

SM(4+2*I) = (RH(I) - BR(I))/DPER

```

```

J      = 10*I

```

```

AM(J+1) = AM(J+1) - DA(I)

```

```

AM(J+3) = AM(J+3) - DA(I)

```

```

AM(J+2) = AM(J+2) - BA(I)

```

```

AM(J+4) = AM(J+4) - BA(I)

```

```

SM(J+1) = SM(J+1) - DR(I)

```

```

SM(J+3) = SM(J+3) - DR(I)

```

```

SM(J+2) = SM(J+2) - BR(I)

```

```

SM(J+4) = SM(J+4) - BR(I)

```

```

AM(J+5) = AM(J+3) - AM(J+1)

```

```

AM(J+6) = AM(J+4) - AM(J+2)

```

```

SM(J+5) = SM(J+3) - SM(J+1)

```

```

SM(J+6) = SM(J+4) - SM(J+2)

```

```

AM(J+7) = HA(I) - DA(I)

```

```

AM(J+8) = AH(I) - BA(I)

```

```

SM(J+7) = HR(I) - DR(I)

```

```

SM(J+8) = RH(I) - BR(I)

```

```

AM(J+9) = HA(I)

```

```

AM(J+10) = AH(I)

```

```

SM(J+9) = HR(I)

```

```

SM(J+10) = RH(I)

```

```

DA(I) = HA(I)

```

```

BA(I) = AH(I)

```

```

DR(I) = HR(I)

```

```

BR(I) = RH(I)

```

```

850 CONTINUE

```



```
WRITE(15,3110) NORB,TIME  
WRITE(15,3120) OD(IOFA),OD(IOTA),AM
```

```
WRITE(15,3140) OD(IOFR),OD(IOTR),SM
```

```
AM(1) = AMAX  
AM(2) = AMAX  
AM(3) = AMIN  
AM(4) = AMIN  
SM(1) = AMAX  
SM(2) = AMAX  
SM(3) = AMIN  
SM(4) = AMIN
```

```
DO 860 I = 1, 3
```

```
J = 10*I  
AM(J+1) = HA(I)  
AM(J+3) = HA(I)  
AM(J+2) = AH(I)  
AM(J+4) = AH(I)  
SM(J+1) = HR(I)  
SM(J+3) = HR(I)  
SM(J+2) = RH(I)  
SM(J+4) = RH(I)
```

```
860 CONTINUE
```

```
880 CONTINUE
```

```
GOTO 100
```

C
C
C

TERMINATION

```
900 CONTINUE
    CLOSE(UNIT=3)
    CLOSE(UNIT=4)

    DO 902 I = 6, 15
    CLOSE(UNIT=I)
902 CONTINUE

STOP
```

C
C
C

ERROR MESSAGES

```
910. CONTINUE
    WRITE(1,*) ' TOO MANY APPENDAGES (>10) '
    GOTO 900

920 CONTINUE
    WRITE(1,*) ' ATTITUDE RELATED SYSTEM ',
1         ' INDEX OUT OF RANGE (1-4) '
    GOTO 900

950 CONTINUE
    WRITE(1,*) ' APPENDAGE SYSTEM DEFINITION CARD ',
1         ' NOT FOUND '
    GOTO 900
```

```

1000 FORMAT(' ORBITAL ELEMENTS:',/,
  1 ' SEMI MAJOR AXIS (M) ',F15.3,/,
  2 ' ECCENTRICITY ',F15.8,/,
  3 ' INCLINATION (DEG) ',F15.3,/,
  4 ' ASCENDING NODE (DEG) ',F15.3,/,
  5 ' PERIGEE ARG. (DEG) ',F15.3,/,
  6 ' MEAN ANOMALY (DEG) ',F15.3,/,
  7 ' ORBITAL PERIOD (MIN) ',F15.3,/)

1001 FORMAT(' START TIME:',/,
  1 ' MONTH ',I11,/,
  2 ' DAY ',I11,/,
  3 ' YEAR ',I11,/,
  4 ' MODIFIED JULIAN DATE ',F13.1,/,
  5 ' HOUR ',I11,/,
  6 ' MINUTES ',I11,/,
  7 ' SECONDS ',F15.3,/,
  8 ' TIME STEP (SECONDS) ',F15.3,/,
  9 ' CALCULATION TIME (S) ',F15.3,/)

1002 FORMAT(' SOLAR FLUX DATA:',/,
  1 ' SOLAR FLUX F10.7 ',F15.2,/,
  2 ' AVERAGED SOLAR FLUX ',F15.2,/,
  1 ' GEOMAG. ACTIVITY KP ',F15.2,/)

1003 FORMAT(' SATELLITE CENTER OF MASS IN BODY',
  1 ' FRAME COORDINATES:',/,
  2 ' X AXIS (METERS) ',F15.5,/,
  3 ' Y AXIS (METERS) ',F15.5,/,
  4 ' Z AXIS (METERS) ',F15.5,/)

1004 FORMAT(' REFERENCE AREA FOR DRAG COEFFICIENT:',/,
  1 ' SATELLITE AREA (M2) ',F15.5,/)

1005 FORMAT(' ATTITUDE RELATED FRAME:',/,
  1 ' FRAME NUMBER ',I9,' - ',A20,/,
  2 ' X-Y-Z ROTATION (DEG) ',3F15.5,/,
  3 ' ANGULAR VELOCITY (RPM) ',3F15.5,/)

1006 FORMAT(' APPENDAGE NUMBER ',I2,/,
  1 ' APPENDAGE POSITION (M) ',3F15.5,/,
  2 ' ANGULAR VELOCITY (RPM) ',3F15.5,/,
  3 ' X-Y-Z ROTATION (DEG) ',3F15.5,/)

3000 FORMAT(16X,'AERODYNAMIC FORCES (N) AND IMPULSES',
  1' (N*S)',/,26X,A20,/,
  2' TIME FORCES',26X,'IMPULSES',/,
  3' (S) X Y Z ',
  4'X Y Z')

4000 FORMAT(16X,'RADIATION FORCES (N) AND IMPULSES',
  1' (N*S)',/,26X,A20,/,
  2' TIME FORCES',26X,'IMPULSES',/,
  3' (S) X Y Z ',
  4'X Y Z')

5000 FORMAT(' AERODYNAMIC TORQUES (N*M) AND',
  1' ANGULAR IMPULSES (N*M*S)',/,26X,A20,/,
  2' TIME',15X,'TORQUES',21X,'ANGULAR IMPULSES',/,
  3' (S) X Y Z ',
  4'X Y Z')

```

```

6000 FORMAT('          RADIATION TORQUES (N*M) AND'
1' ANGULAR IMPULSES (N*M*S)',/,26X,A20,/,
2' TIME',15X,'TORQUES',21X,'ANGULAR IMPULSES',/,
3'   (S)      X      Y      Z      ',
4'X          Y          Z')
7000 FORMAT('          ORBIT DATA',/,
1' ORBIT MEAN GEOCENTR. SATEL. LONGI- LATI- ',
2'GEOCENTR. SUN R. SUN',/,
3' TIME ANOM. DISTANCE VELOC. TUDE TUDE ',
4' ALTITUDE ASCEN. DECLI.',/,
5'   (S) (DEG) (KM) (KM/S) (DEG) (DEG) ',
6'   (KM) (DEG) (DEG)')
8000 FORMAT(29X,'ATMOSPHERIC DATA',/,
1' ORBIT ATM. LOCAL SPEED ANG. ANG. ',
2'DRAG COMPONENTS DRAG',/,
3' TIME TEMP DENSITY RATIO ALPHA BETA ',
4'   COEFF.',/,
5'   (S) (K) (KG/M3) (DEG) (DEG) X',
6'   Y Z ')
9000 FORMAT(22X,'SUN ASPECT ANGLE (DEGREES)',/,
1' ORBIT SPACECRAFT APPENDAGE APPENDAGE ',
2'APPENDAGE APPENDAGE',/,
3' TIME MAIN BODY NUMBER 1 NUMBER 2 ',
4' NUMBER 3 NUMBER 4',/,
5'   (S) ALPHA BETA ALPHA BETA ALPHA BETA ALPHA',
6' BETA ALPHA BETA')
1120 FORMAT(20X,I6)
1320 FORMAT(I6,1X,5(E10.4,1X),E10.4)
1720 FORMAT(I6,1X,F6.2,1X,F9.3,1X,F6.3,1X,F6.2,1X,
1 F6.2,1X,F9.3,1X,F6.2,1X,F6.2)
1820 FORMAT(I6,1X,I4,1X,E9.3,1X,F5.2,1X,F6.2,1X,F6.2,1X,
1 3(F7.3,1X),F6.3)
9020 FORMAT(I6,10(1X,F5.1,1X,F5.1))
3110 FORMAT(' ORBIT NUMBER ',I4,' T = ',I6,' S',/)

```

3120 FORMAT(' AERODYNAMIC FORCES AND TORQUES:',/,

```

1'           ',A20,1X,A20,/,/,
2' AERODYNAMIC          FORCE          TORQUE',/,
3' MINIMUM             ',E12.6,' N    ',E12.6,' N*M ',/,
4' MAXIMUM             ',E12.6,' N    ',E12.6,' N*M ',/,
5' X MEAN              ',E12.6,' N    ',E12.6,' N*M ',/,
6' Y MEAN              ',E12.6,' N    ',E12.6,' N*M ',/,
7' Z MEAN              ',E12.6,' N    ',E12.6,' N*M ',/,
8' AEROD. IMPULSE     LINEAR          ANGULAR',/,
9' X MINIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
A' X MAXIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
1' X DIFF.            ',E12.6,' N*S   ',E12.6,' N*M*S',/,
2' X BUILD-UP         ',E12.6,' N*S   ',E12.6,' N*M*S',/,
3' X ACCUM.           ',E12.6,' N*S   ',E12.6,' N*M*S',/,
4' Y MINIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
5' Y MAXIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
6' Y DIFF.            ',E12.6,' N*S   ',E12.6,' N*M*S',/,
7' Y BUILD-UP         ',E12.6,' N*S   ',E12.6,' N*M*S',/,
8' Y ACCUM.           ',E12.6,' N*S   ',E12.6,' N*M*S',/,
9' Z MINIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
B' Z MAXIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
1' Z DIFF.            ',E12.6,' N*S   ',E12.6,' N*M*S',/,
2' Z BUILD-UP         ',E12.6,' N*S   ',E12.6,' N*M*S',/,
3' Z ACCUM.           ',E12.6,' N*S   ',E12.6,' N*M*S',/)

```

3140 FORMAT(' SOLAR RADIATION FORCES AND TORQUES:',/,

```

1'           ',A20,1X,A20,/,/,
2' SOLAR RADIATION     FORCE          TORQUE',/,
3' MINIMUM             ',E12.6,' N    ',E12.6,' N*M ',/,
4' MAXIMUM             ',E12.6,' N    ',E12.6,' N*M ',/,
5' X MEAN              ',E12.6,' N    ',E12.6,' N*M ',/,
6' Y MEAN              ',E12.6,' N    ',E12.6,' N*M ',/,
7' Z MEAN              ',E12.6,' N    ',E12.6,' N*M ',/,
8' S. RAD. IMPULSE    LINEAR          ANGULAR',/,
9' X MINIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
A' X MAXIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
1' X DIFF.            ',E12.6,' N*S   ',E12.6,' N*M*S',/,
2' X BUILD-UP         ',E12.6,' N*S   ',E12.6,' N*M*S',/,
3' X ACCUM.           ',E12.6,' N*S   ',E12.6,' N*M*S',/,
4' Y MINIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
5' Y MAXIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
6' Y DIFF.            ',E12.6,' N*S   ',E12.6,' N*M*S',/,
7' Y BUILD-UP         ',E12.6,' N*S   ',E12.6,' N*M*S',/,
8' Y ACCUM.           ',E12.6,' N*S   ',E12.6,' N*M*S',/,
9' Z MINIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
B' Z MAXIMUM          ',E12.6,' N*S   ',E12.6,' N*M*S',/,
1' Z DIFF.            ',E12.6,' N*S   ',E12.6,' N*M*S',/,
2' Z BUILD-UP         ',E12.6,' N*S   ',E12.6,' N*M*S',/,
3' Z ACCUM.           ',E12.6,' N*S   ',E12.6,' N*M*S',/)

```

END

SUBROUTINE ARELEM(NF,VU,VQ,SG,AE)

C
C THE SUBROUTINE ARELEM READS THE GRID ELEMENTS DATA
C FILE AND CALCULATES ITS NORMAL, AREA AND THE POSITION
C OF ITS CENTER OF PRESSURE. ALL THE ELEMENTS ARE STORED
C IN ARRAYS, TO BE USED IN THE MAIN PROGRAM.
C

C
C OUTPUTS:

C NF(10) ARRAY CONTAINING THE NUMBER OF SURFACE
C ELEMENTS OF THE SPACECRAFT BODY (NF(1)) AND
C EACH APPENDAGE (NF(2) ... NF(10) - MAXIMUM
C 9 APPENDAGES).

C VU(50,3) COMPONENTS OF THE SURFACE NORMAL, FOR EACH
C ELEMENT (MAXIMUM 50 ELEMENTS).

C VQ(50,3) POSITION, IN METERS, OF THE ELEMENT CENTER
C OF PRESSURE, RELATIVE TO THE SPACECRAFT
C SYSTEM.

C SG(50,7) SURFACE CHARACTERISTICS, RESPECTIVELY:

C TANGENTIAL AERODYNAMIC COEFFICIENT.

C NORMAL AERODYNAMIC COEFFICIENT.

C SOLAR SPECULAR REFLECTION COEFFICIENT.

C SOLAR DIFFUSE REFLECTION COEFFICIENT.

C SOLAR ABSORPTION COEFFICIENT.

C INFRA-RED EMMITANCE COEFFICIENT.

C ABSOLUTE TEMPERATURE OF THE ELEMENT IN K.

C AE(50) AREA OF EACH ELEMENT IN M**2.
C

C
C AUTHOR: VALDEMIR CARRARA, MAY 1987.
C

C IMPLICIT REAL*8 (A-H,O-Z)

C PARAMETER (NSUR=50, NGRD=50)

C DIMENSION GR(NGRD,3),VV(NSUR,3),VP(NSUR,3),AR(NSUR)

C DIMENSION SG(NSUR,7),VU(NSUR,3),VQ(NSUR,3),AE(NSUR)

C DIMENSION IG(NGRD),NS(NSUR),NF(11)

C DIMENSION EG(6),VN(3),PV(3),SC(7)

C DIMENSION KG(4)

C CHARACTER *8 CQ

C WRITE(6,1600)

C WRITE(7,1500)

C NG = 0

C
C
C

READ THE GRID ELEMENTS FROM SATGE.DAT FILE

```
100 CONTINUE
    READ(3,1400,END=200) CQ
1400 FORMAT(A8,2I8,4F8.2)
    IF(CQ.EQ.'ENDDATA ') GOTO 200
    IF(CQ.NE.'GRID ') GOTO 100
```

```
    BACKSPACE 3
    READ(3,1400) CQ,EG
    NG = NG + 1
    IF(NG.GT.NGRD) GOTO 920
    GR(NG,1) = EG(3)
    GR(NG,2) = EG(4)
    GR(NG,3) = EG(5)
    IG(NG) = EG(1)
    GOTO 100
```

```
200 CONTINUE
    IF(NG.LT.3) GOTO 900
    REWIND 3
    IELE = 0
```

C
C
C

READ THE SURFACE DEFINITION CARDS FROM SATGE.DAT

```
250 CONTINUE
    READ(3,1410,END=600) CQ
    IF(CQ.EQ.'ENDDATA ') GOTO 600
    IF(CQ.NE.'CTRIA3 ') GOTO 300
```

C
C
C

TRIANGLE ELEMENT COMPUTATIONS

```
    BACKSPACE 3
    READ(3,1410) CQ,EG
    IELE = IELE + 1
    IF(IELE.GT.NSUR) GOTO 930

    DO 270 I = 1, 3
    JG = 0
260 CONTINUE
    JG = JG + 1
    IF(JG.GT.NG) GOTO 270
    IF(IG(JG).EQ.EG(2+I)) KG(I) = JG
    GOTO 260
270 CONTINUE
```

```

XD21 = GR(KG(2),1) - GR(KG(1),1)
XD31 = GR(KG(3),1) - GR(KG(1),1)
YD21 = GR(KG(2),2) - GR(KG(1),2)
YD31 = GR(KG(3),2) - GR(KG(1),2)
ZD21 = GR(KG(2),3) - GR(KG(1),3)
ZD31 = GR(KG(3),3) - GR(KG(1),3)

```

```

XVTC = YD21*ZD31 - ZD21*YD31
YVTC = ZD21*XD31 - XD21*ZD31
ZVTC = XD21*YD31 - YD21*XD31

```

```

ARMO = DSQRT(XVTC*XVTC + YVTC*YVTC + ZVTC*ZVTC)
AREA = ARMO*.5D-06
VN(1) = XVTC/ARMO
VN(2) = YVTC/ARMO
VN(3) = ZVTC/ARMO

```

```

PV(1) = (GR(KG(1),1)+GR(KG(2),1)+GR(KG(3),1))/3.D3
PV(2) = (GR(KG(1),2)+GR(KG(2),2)+GR(KG(3),2))/3.D3
PV(3) = (GR(KG(1),3)+GR(KG(2),3)+GR(KG(3),3))/3.D3

```

```

NS(IELE) = EG(1)
AR(IELE) = AREA

```

```

DO 280 I = 1, 3
VV(IELE,I) = VN(I)
VP(IELE,I) = PV(I)

```

```
280 CONTINUE
```

```
GOTO 250
```

```
300 CONTINUE
```

```
IF(CQ.NE.'CQUAD4 ') GOTO 400
```

C
C
C

QUADRILATERAL ELEMENT COMPUTATIONS

```

BACKSPACE 3
READ(3,1410) CQ,EG
IELE = IELE + 1
IF(IELE.GT.NSUR) GOTO 930

```

```

DO 320 I = 1, 4
JG = 0

```

```
310 CONTINUE
```

```

JG = JG + 1
IF(JG.GT.NG) GOTO 320
IF(IG(JG).EQ.EG(2+I)) KG(I) = JG
GOTO 310

```

```
320 CONTINUE
```


XD21 = GR(KG(2),1) - GR(KG(1),1)
 XD31 = GR(KG(3),1) - GR(KG(1),1)
 YD21 = GR(KG(2),2) - GR(KG(1),2)
 YD31 = GR(KG(3),2) - GR(KG(1),2)
 ZD21 = GR(KG(2),3) - GR(KG(1),3)
 ZD31 = GR(KG(3),3) - GR(KG(1),3)

XVTC = YD21*ZD31 - ZD21*YD31
 YVTC = ZD21*XD31 - XD21*ZD31
 ZVTC = XD21*YD31 - YD21*XD31

ARMO = DSQRT(XVTC*XVTC + YVTC*YVTC + ZVTC*ZVTC)
 AREA = ARMO*.5D-06

VN(1) = XVTC/ARMO
 VN(2) = YVTC/ARMO
 VN(3) = ZVTC/ARMO

XD21 = GR(KG(3),1) - GR(KG(1),1)
 XD31 = GR(KG(4),1) - GR(KG(1),1)
 YD21 = GR(KG(3),2) - GR(KG(1),2)
 YD31 = GR(KG(4),2) - GR(KG(1),2)
 ZD21 = GR(KG(3),3) - GR(KG(1),3)
 ZD31 = GR(KG(4),3) - GR(KG(1),3)

XVTC = YD21*ZD31 - ZD21*YD31
 YVTC = ZD21*XD31 - XD21*ZD31
 ZVTC = XD21*YD31 - YD21*XD31

ARMO = DSQRT(XVTC*XVTC + YVTC*YVTC + ZVTC*ZVTC)
 AREA = AREA + ARMO*.5D-06

PV(1) = (GR(KG(1),1) + GR(KG(2),1) +
 1 GR(KG(3),1) + GR(KG(4),1))/4.D3
 PV(2) = (GR(KG(1),2) + GR(KG(2),2) +
 1 GR(KG(3),2) + GR(KG(4),2))/4.D3
 PV(3) = (GR(KG(1),3) + GR(KG(2),3) +
 1 GR(KG(3),3) + GR(KG(4),3))/4.D3

NS(IELE) = EG(1)
 AR(IELE) = AREA

DO 340 I = 1, 3
 VV(IELE,I) = VN(I)
 VP(IELE,I) = PV(I)

340 CONTINUE

GOTO 250

400 CONTINUE

GOTO 250

C
C
C

PLASE.DAT FILE OUTPUTS

```
600 CONTINUE
    DO 605 J = 1, IELE
        WRITE(6,1610) J,NS(J),(VV(J,I),I=1,3),
1          (VP(J,I),I=1,3),AR(J)
605 CONTINUE
```

C
C
C

READ THE BODYAP COMMANDS FROM THE INELE.DAT FILE

```
IELE = 0
NBOD = -1
IFLA = 1
```

```
610 CONTINUE
    IF(IFLA.EQ.0) GOTO 660
    IELA = 0
    NBOD = NBOD + 1
    IF(NBOD.GE.10) GOTO 950
    IFLA = 0
    REWIND 4
```

```
620 CONTINUE
    READ(4,*,END=610) CQ
    IF(CQ.EQ.'ENDDATA ') GOTO 610
    IF(CQ.NE.'BODYAP ') GOTO 620
    BACKSPACE 4
    READ(4,*) CQ,INBO,NUEL,SC
    IF(INBO.NE.NBOD) GOTO 620
    IFLA = 1
    NUEA = IABS(NUEL)
    IELE = IELE + 1
    IELA = IELA + 1
    NF(NBOD+1) = IELA
    SNUE = 1.DO
    IF(NUEL.LT.0) SNUE = -1.DO
    NG = 0
```

```
630 CONTINUE
    NG = NG + 1
    IF(NG.GT.NSUR) GOTO 940
    IF(NUEA.NE.NS(NG)) GOTO 630

    AE(IELE) = AR(NG)
```

```
DO 640 I = 1, 3
  SG(IELE,I) = SC(I)
  SG(IELE,I+3) = SC(I+3)
  VQ(IELE,I) = VP(NG,I)
  VU(IELE,I) = SNUE*VV(NG,I)
640 CONTINUE
  SG(IELE,7) = SC(7)
```

C
C
C

EXTSE.DAT FILE OUTPUTS

```
WRITE(7,1510) NBOD,IELA,(VU(IELE,I),I=1,3),
1 (VQ(IELE,I),I=1,3),AE(IELE)
GOTO 620

660 CONTINUE

RETURN
```

C
C
C

ERROR MESSAGES

```
900 CONTINUE
  WRITE(1,*) ' TOO FEW GRID ELEMENTS IN DATA FILE'
  GOTO 990
920 CONTINUE
  WRITE(1,*) ' TOO MANY GRID ELEMENTS (>50) ',
1 ' IN DATA FILE'
  GOTO 990
930 CONTINUE
  WRITE(1,*) ' TOO MANY SURFACE ELEMENTS (>50)'
  GOTO 990
940 CONTINUE
  WRITE(1,*) ' SURFACE NUMBER NOT FOUND'
  GOTO 990
950 CONTINUE
  WRITE(1,*) ' TOO MANY APPENDAGES (>10)'
  GOTO 990

990 CONTINUE
  STOP
```

1410 FORMAT(A8,6I8)

1500 FORMAT(18X,'SPACECRAFT EXTERNAL SURFACE ELEMENTS',/,
1' BODY EXTERNAL NORMAL VECTOR ELEMENT ',
2' POSITION (M) AREA',/,
3' NUBR X Y Z X ',
4' Y Z (M2)')

1510 FORMAT(I2,I3,3F9.5,4E10.4)

1600 FORMAT(24X,'SPACECRAFT SURFACE ELEMENTS',/,
1' INT IDENT EXT. NORMAL VECTOR ELEMENT ',
2' POSITION (M) AREA',/,
3' NBR NUMBR X Y Z X ',
4' Y Z (M2)')

1610 FORMAT(I4,I6,3F7.4,3E10.4,E10.4)

END

```
SUBROUTINE REAINP(CHAT,NFIL,NUBR,AY)
```

```
C THE SUBROUTINE REAINP READS THE INPUT ELEMENTS GIVEN  
C BY ITS NAME AND RETURNS WITH THE ELEMENTS IN AY.  
C
```

```
C INPUTS:
```

```
C CHAT NAME OF THE INPUT ELEMENTS.  
C NFIL NUMBER OF THE FILE.  
C NUBR NUMBER OF ELEMENTS THAT MUST BE READ.  
C
```

```
C OUTPUTS:
```

```
C AY ARRAY CONTAINING THE ELEMENTS.  
C
```

```
C AUTHOR: VALDEMIR CARRARA, MAY 87.  
C
```

```
IMPLICIT REAL*8 (A-H,O-Z)  
CHARACTER *8 CHAT,CQUE
```

```
DIMENSION AY(1)  
REWIND NFIL
```

```
100 CONTINUE  
READ(NFIL,*,END=200) CQUE  
IF(CQUE.EQ.'ENDDATA ') GOTO 200  
IF(CQUE.NE.CHAT) GOTO 100  
BACKSPACE NFIL  
READ(NFIL,*) CQUE,(AY(I),I=1,NUBR)  
RETURN
```

```
.200 CONTINUE  
WRITE(1,*) CHAT,' DATA CARD NOT FOUND'  
STOP  
END
```

FILE AUROT.FTN

FTN77,J
\$CDS ON

SUBROUTINE AERODY(AEPR,UV,VN,RG,SPRT,TWTI,SC,AREA,FC,
1 TC)

C
C THE SUBROUTINE AERODY OBTAINS THE AERODYNAMIC FORCE
C AND TORQUE COEFFICIENTS FOR AN EXTERNAL SURFACE PLANE
C ELEMENT. THE SUBROUTINE AERODY USES THE FREE MOLECULAR
C FLOW THEORY TO COMPUTE THE ELEMENTARY FORCES AND TOR-
C QUES.
C

INPUTS:

C AEPR AERODYNAMIC PRESSURE:
C HALF OF THE LOCAL ATMOSPHERIC DENSITY TIMES
C THE SQUARE OF THE SATELLITE VELOCITY, TO
C CALCULATE THE FORCES OR THE UNIT TO CALCULATE
C THE COEFFICIENTS.
C UV UNIT VECTOR OF THE SATELLITE VELOCITY IN THE
C BODY FRAME COORDINATES.
C VN UNIT VECTOR NORMAL TO THE ELEMENT.
C RG POSITION OF THE ELEMENT CENTRE OF PRESSURE
C RELATIVE TO THE GRAVITY CENTRE OF SPACECRAFT,
C IN BODY FRAME COORDINATES, IN METERS.
C SPRT SPEED RATIO (MAGNITUDE OF THE SATELLITE
C VELOCITY DIVIDED BY THE MOST PROBABLE VELOCITY
C OF THE MOLECULES).
C TWTI RATIO BETWEEN THE SURFACE ABSOLUTE TEMPERATURE
C OF THE ELEMENT AND THE LOCAL TEMPERATURE.
C SC ARRAY THAT CONTAINS THE SURFACE CHARACTERISTICS
C OF THE ELEMENT:
C SC(1) TANGENTIAL MOMENTUM COEFFICIENT.
C SC(2) NORMAL MOMENTUM EXCHANGE COEFFICIENT.
C AREA AREA OF THE ELEMENT IN METERS**2.
C

OUTPUTS:

C FC FORCE COEFFICIENTS OF THE SATELLITE IN THE
C BODY FRAME.
C TC TORQUE COEFFICIENTS OF THE SATELLITE IN THE
C BODY FRAME.
C

SUBROUTINES CALLED IN THIS SUBPROGRAM:

C DERF
C

AUTHOR:

C VALDEMIR CARRARA MAR/87
C

```
IMPLICIT REAL*8 (A-H,O-Z)
```

```
DIMENSION UV(3),VN(3),RG(3),SC(7),FC(3),TC(3)
```

```
COMMON /CONSTA/ PI,PIT2,PID2,RAD,DEG  
DATA SQPI /1.7724538509D0/
```

```
XDIR = UV(1)  
YDIR = UV(2)  
ZDIR = UV(3)
```

```
XNOR = VN(1)  
YNOR = VN(2)  
ZNOR = VN(3)
```

```
TWTS = DSQRT(TWTI)  
SIGM = SC(1)  
SIGP = SC(2)
```

```
CONE = XNOR*XDIR + YNOR*YDIR + ZNOR*ZDIR
```

```
IF(DABS(CONE).GT.1.D0) CONE = DSIGN(1.D0,CONE)
```

```
SINE = DSQRT(1.D0 - CONE*CONE)
```

```
IF(SINE.EQ.0.D0) GOTO 200
```

```
XTOX = (XDIR - XNOR*CONE)/SINE  
YTOX = (YDIR - YNOR*CONE)/SINE  
ZTOX = (ZDIR - ZNOR*CONE)/SINE
```

```
200 CONTINUE
```

```
ESCO = SPRT*CONE  
ERF1 = 1.D0 + DERF(ESCO)  
ESC2 = ESCO*ESCO  
EXSE = DEXP(-ESC2)  
APAS = AEPR*AREA/SPRT
```

```
PAER = -(EXSE*((2.D0-SIGM)/SQPI*ESCO +  
1 .5D0*SIGP*TWTS) +  
2 ERF1*((2.D0-SIGP)*(.5D0+ESC2) +  
3 .5D0*SIGP*TWTS*SQPI*ESCO))*APAS/SPRT  
TAER = -SIGM*SINE*(EXSE/SQPI + ESCO*ERF1)*APAS
```

```
FC(1) = PAER*XNOR + TAER*XTOX  
FC(2) = PAER*YNOR + TAER*YTOX  
FC(3) = PAER*ZNOR + TAER*ZTOX  
TC(1) = FC(3)*RG(2) - FC(2)*RG(3)  
TC(2) = FC(1)*RG(3) - FC(3)*RG(1)  
TC(3) = FC(2)*RG(1) - FC(1)*RG(2)
```

```
RETURN  
END
```

SUBROUTINE SOLPRE(RAPR,UV,VN,RG,SC,AREA,FC,TC)

THE SUBROUTINE SOLPRE OBTAINS THE SOLAR RADIATION FORCE AND TORQUE FOR A SURFACE ELEMENT OF THE SATELLITE.

INPUTS:

RAPR RADIATION PRESSURE AT THE LOCAL, IN N/M**2.
UV(3) UNIT VECTOR OF THE SOLAR RADIATION DIRECTION,
IN THE BODY FRAME COORDINATES.
VN(3) UNIT VECTOR NORMAL TO THE SURFACE ELEMENT.
RG(3) POSITION OF THE ELEMENT CENTRE OF PRESSURE
RELATIVE TO THE SPACECRAFT GRAVITY CENTRE,
IN THE BODY FRAME COORDINATES, IN METERS.
SC(6) SURFACE CHARACTERISTICS:
SC(3) SPECULAR REFLECTION COEFFICIENT.
SC(4) DIFFUSE REFLECTION COEFFICIENT.
SC(5) SURFACE ABSORPTION COEFFICIENT.
SC(6) SURFACE EMISSIVITY OF THE ELEMENT.
SC(7) ABSOLUTE TEMPERATURE OF THE ELEMENT.
AREA AREA OF THE ELEMENT IN M**2.

OUTPUTS:

FC RADIATION FORCE ON THE ELEMENT, IN BODY FRAME
COORDINATES, IN NEWTONS.
TC RADIATION TORQUE ON THE ELEMENT, IN N*M, IN
BODY FRAME COORDINATES.

AUTHOR:

VALDEMIR CARRARA MAR/87

IMPLICIT REAL*8 (A-H,O-Z)

DIMENSION UV(3),VN(3),RG(3),SC(7),FC(3),TC(3)

COMMON /CONSTA/ PI,PIT2,PID2,RAD,DEG

DATA STEC /1.891208D-19/

XDIR = UV(1)
YDIR = UV(2)
ZDIR = UV(3)

XNOR = VN(1)
YNOR = VN(2)
ZNOR = VN(3)

CONE = XNOR*XDIR + YNOR*YDIR + ZNOR*ZDIR

IF(DABS(CONE).GT.1.DO) CONE = DSIGN(1.DO,CONE)

RENI = SC(6)
TEMP = SC(7)

IF(CONE.GT.0.DO) GOTO 200
EMMI = TEMP*TEMP
PSOL = -2.DO*AREA*RENI*STEC*EMMI*EMMI/3.DO
TSOL = 0.DO
GOTO 300

200 CONTINUE

SIES = SC(3)
SIDI = SC(4)
SIAB = SC(5)
EMMI = TEMP*TEMP
EMMI = RENI*STEC*EMMI*EMMI
CON2 = CONE*CONE

IF(CON2.GT.1.DO) CON2 = 1.DO

SINE = DSQRT(1.DO - CON2)

IF(SINE.EQ.0.DO) GOTO 250
XTOX = (XDIR - XNOR*CONE)/SINE
YTOX = (YDIR - YNOR*CONE)/SINE
ZTOX = (ZDIR - ZNOR*CONE)/SINE

250 CONTINUE

PSOL = -((2.DO/3.DO*SIDI +
1 (SIAB+2.DO*SIES+SIDI)*CONE)*CONE*RAPR +
2 2.DO/3.DO*EMMI)*AREA
TSOL = -(SIAB+SIDI)*CONE*SINE*RAPR*AREA

300 CONTINUE

FC(1) = PSOL*XNOR + TSOL*XTOX
FC(2) = PSOL*YNOR + TSOL*YTOX
FC(3) = PSOL*ZNOR + TSOL*ZTOX
TC(1) = FC(3)*RG(2) - FC(2)*RG(3)
TC(2) = FC(1)*RG(3) - FC(3)*RG(1)
TC(3) = FC(2)*RG(1) - FC(1)*RG(2)

RETURN
END

SUBROUTINE GEOCEN(GT,GC)

THE SUBROUTINE GEOCEN OBTAINS THE GEOCENTRIC COORDINATES OF A POINT GIVEN BY ITS TERRESTRIAL GEOCENTRIC COORDINATES.

INPUTS:

GV(3) GEOCENTRIC COORDINATES IN METERS.

OUTPUTS:

GC(1) LONGITUDE IN RADIANS.

GC(2) GEOCENTRIC LATITUDE IN RADIANS.

GC(3) GEOCENTRIC ALTITUDE IM METERS.

AUTHOR: VALDEMIR CARRARA, APR/87

IMPLICIT REAL*8 (A-H,O-Z)
DIMENSION GT(3),GC(3)

COMMON /CONSTA/ PI,PIT2,PID2,RAD,DEG
COMMON /CEARTH/ RE,GM,FLAT,TETP

GVXY = GT(1)*GT(1) + GT(2)*GT(2)

GXYS = DSQRT(GVXY)

GC(1) = DATAN2(GT(2)/GXYS,GT(1)/GXYS)

GC(2) = DATAN2(GT(3),GXYS)

GVZZ = GT(3)*GT(3)

GVM2 = GVXY + GVZZ

FACT = 1.DO - FLAT

FACT = 1.DO/FACT/FACT - 1.DO

GC(3) = (DSQRT(GVM2) - RE/DSQRT(1.DO+FACT*GVZZ/GVM2))

RETURN

END

```
SUBROUTINE ROTMAT(ANGL,WV,RM)
```

```
THIS SUBROUTINE OBTAINS THE ROTATION MATRIX GIVEN THE  
VECTOR AND THE ANGLE OF ROTATION ABOUT THIS VECTOR.
```

```
INPUTS:
```

```
  ANGL  ANGLE OF ROTATION IN RADIANS.  
  WV    UNIT VECTOR OF THE ROTATION DIRECTION.
```

```
OUTPUTS:
```

```
  RM    ROTATION MATRIX.
```

```
AUTHOR:  VALDEMIR CARRARA,  APR/87
```

```
IMPLICIT REAL*8 (A-H,O-Z)  
DIMENSION WV(3),RM(3,3)
```

```
COAN    = DCOS(ANGL)  
SIAN    = DSIN(ANGL)  
COM1    = 1.DO - COAN
```

```
RM(1,1) = COM1*WV(1)*WV(1) + COAN  
RM(1,2) = COM1*WV(1)*WV(2) + SIAN*WV(3)  
RM(1,3) = COM1*WV(1)*WV(3) - SIAN*WV(2)  
RM(2,1) = COM1*WV(2)*WV(1) - SIAN*WV(3)  
RM(2,2) = COM1*WV(2)*WV(2) + COAN  
RM(2,3) = COM1*WV(2)*WV(3) + SIAN*WV(1)  
RM(3,1) = COM1*WV(3)*WV(1) + SIAN*WV(2)  
RM(3,2) = COM1*WV(3)*WV(2) - SIAN*WV(1)  
RM(3,3) = COM1*WV(3)*WV(3) + COAN
```

```
RETURN  
END
```

```
SUBROUTINE MATMUL(R1,R2,RS)
```

```
THE SUBROUTINE MATMUL CALCULATES THE PRODUCT OF TWO  
3X3 INPUT MATRIX, R1 AN R2.
```

```
OUTPUTS:
```

```
RS = R1*R2
```

```
AUTHOR: VALDEMIR CARRARA, APR/87
```

```
IMPLICIT REAL*8 (A-H,O-Z)
```

```
DIMENSION R1(3,3),R2(3,3),RS(3,3)
```

```
DO 20 I=1,3
```

```
DO 20 J=1,3
```

```
RS(I,J) = R1(I,1)*R2(1,J) + R1(I,2)*R2(2,J) +  
1 R1(I,3)*R2(3,J)
```

```
20 CONTINUE
```

```
RETURN
```

```
END
```

```
SUBROUTINE VECROT(RM,VI,VO)
```

```
THE SUBROUTINE VECROT CALCULATES THE COORDINATES OF A  
VECTOR IN OTHER REFERENCE SYSTEM, GIVEN THE ROTATION  
MATRIX BETWEEN THE FRAMES.
```

```
INPUTS:
```

```
RM      ROTATION MATRIX.  
VI      VECTOR IN THE FIRST SYSTEM.
```

```
OUTPUTS:
```

```
VO      ROTATED VECTOR (VO = RM*VI).
```

```
AUTHOR:  VALDEMIR CARRARA, APR/87
```

```
IMPLICIT REAL*8 (A-H,O-Z)  
DIMENSION RM(3,3),VI(3),VO(3)
```

```
DO 20 I=1,3  
VO(I) = RM(I,1)*VI(1) + RM(I,2)*VI(2) + RM(I,3)*VI(3)  
20 CONTINUE
```

```
RETURN  
END
```

SUBROUTINE STAVEC(EL,SV,RM)

THIS SUBROUTINE OBTAINS THE STATE VECTOR (CARTESIAN POSITION AND VELOCITY) OF A SATELLITE, GIVEN THE ORBITAL KEPLERIAN ELEMENTS.

INPUTS:

EL(6) KEPLERIAN ELEMENTS:
 EL(1) SEMI-MAJOR AXIS, IN METERS.
 EL(2) ECCENTRICITY
 EL(3) INCLINATION IN RADIANS.
 EL(4) RIGHT ASCENSION OF ASCENDING NODE IN RAD.
 EL(5) PERIGEE ARGUMENT IN RAD.
 EL(6) MEAN ANOMALY IN RAD.

OUTPUTS:

SV POSITION (SV(1), SV(2) AND SV(3)) IN METERS AND VELOCITY (SV(4), SV(5) AND SV(6)) IN METERS PER SECOND OF THE SPACECRAFT IN THE INERTIAL FRAME.
 RM ROTATION MATRIX FROM THE INERTIAL FRAME TO THE ORBIT FRAME (X UPWARD, Y IN THE VELOCITY DIRECTION AND Z NORMAL TO THE ORBIT PLANE).

SUBROUTINES CALLED IN THIS SUBPROGRAM:

KEPLER

AUTHOR: VALDEMIR CARRARA APR/87

IMPLICIT REAL*8 (A-H,O-Z)
 DIMENSION EL(6),SV(6),RM(3,3)

COMMON /CONSTA/ PI,PIT2,PID2,RAD,DEG
 COMMON /CEARTH/ RE,GM,FLAT,TETP

COSI = DCOS(EL(3))
 SINI = DSIN(EL(3))
 CORA = DCOS(EL(4))
 SIRA = DSIN(EL(4))
 COPA = DCOS(EL(5))
 SIPA = DSIN(EL(5))

A11 = COPA*CORA - SIPA*SIRA*COSI
 A12 = COPA*SIRA + SIPA*CORA*COSI
 A13 = SIPA*SINI
 A21 = -SIPA*CORA - COPA*SIRA*COSI
 A22 = -SIPA*SIRA + COPA*CORA*COSI
 A23 = COPA*SINI

ECCE = EL(2)
ANOM = EL(6)

CALL KEPLER(ECCE, ANOM, ECAN)

COEA = DCOS(ECAN)
SIEA = DSIN(ECAN)

SECC = DSQRT(1.00 - ECCE*ECCE)
ECEA = 1.00 - ECCE*COEA
VELO = DSQRT(GM/EL(1))/ECEA
CEME = COEA * ECCE

XORB = EL(1)*CEME
YORB = EL(1)*SECC*SIEA
XDOT = VELO*SECC*COEA
YDOT = VELO*SIEA

SV(1) = A11*XORB + A21*YORB
SV(2) = A12*XORB + A22*YORB
SV(3) = A13*XORB + A23*YORB
SV(4) = A21*XDOT - A11*YDOT
SV(5) = A22*XDOT - A12*YDOT
SV(6) = A23*XDOT - A13*YDOT

COTA = CEME/ECEA
SITA = SECC*SIEA/ECEA

RM(1,1) = A11*COTA + A21*SITA
RM(1,2) = A12*COTA + A22*SITA
RM(1,3) = A13*COTA + A23*SITA
RM(2,1) = A21*COTA - A11*SITA
RM(2,2) = A22*COTA - A12*SITA
RM(2,3) = A23*COTA - A13*SITA
RM(3,1) = SINI*SIRA
RM(3,2) = -SINI*CORA
RM(3,3) = COSI

RETURN
END

SUBROUTINE ROTXYZ(EA,RS)

THE SUBROUTINE ROTXYZ CALCULATES THE ROTATION MATRIX
GIVEN THE EULER ANGLES OF A X-Y-Z ROTATION.

INPUTS:

EA EULER ANGLES, IN RADIANS, ABOUT THE X, Y
AND Z AXIS, RESPECTIVELY.

OUTPUTS:

RS ROTATION MATRIX.

SUBROUTINES CALLED IN THIS SUBPROGRAM:

ROTMAX

ROTMAY

ROTMZ

MATMUL

AUTHOR: VALDEMIR CARRARA, APR/87

IMPLICIT REAL*8 (A-H,O-Z)
DIMENSION EA(3),RS(3,3),R1(3,3),R2(3,3)

CALL ROTMAX(EA(1),RS)
CALL ROTMAY(EA(2),R1)
CALL MATMUL(R1,RS,R2)
CALL ROTMAZ(EA(3),R1)
CALL MATMUL(R1,R2,RS)

RETURN
END

SUBROUTINE ROTMAX(ANGL,RS)

C
C
C
C
C
C
C
C
C
C

THIS SUBROUTINE OBTAINS THE ROTATION MATRIX ABOUT THE
X AXIS, GIVEN THE ANGLE OF ROTATION ABOUT THIS AXIS.

INPUTS:

ANGL ANGLE OF ROTATION IN RADIAN.

OUTPUTS:

RS ROTATION MATRIX.

AUTHOR: VALDEMIR CARRARA, APR/87

IMPLICIT REAL*8 (A-H,O-Z)
DIMENSION RS(3,3)

COAN = DCOS(ANGL)
SIAN = DSIN(ANGL)

RS(1,1) = 1.DO
RS(1,2) = 0.DO
RS(1,3) = 0.DO
RS(2,1) = 0.DO
RS(2,2) = COAN
RS(2,3) = SIAN
RS(3,1) = 0.DO
RS(3,2) = -SIAN
RS(3,3) = COAN

RETURN
END

SUBROUTINE ROTMAY(ANGL,RS)

THIS SUBROUTINE OBTAINS THE ROTATION MATRIX ABOUT THE
Y AXIS, GIVEN THE ANGLE OF ROTATION ABOUT THIS AXIS.

INPUTS:

ANGL ANGLE OF ROTATION IN RADIANS.

OUTPUTS:

RS ROTATION MATRIX.

AUTHOR: VALDEMIR CARRARA, APR/87

IMPLICIT REAL*8 (A-H,O-Z)
DIMENSION RS(3,3)

COAN = DCOS(ANGL)
SIAN = DSIN(ANGL)

RS(1,1) = COAN
RS(1,2) = 0.DO
RS(1,3) = -SIAN
RS(2,1) = 0.DO
RS(2,2) = 1.DO
RS(2,3) = 0.DO
RS(3,1) = SIAN
RS(3,2) = 0.DO
RS(3,3) = COAN

RETURN
END

C
C
C
C
C
C
C
C
C
C

SUBROUTINE ROTMAZ(ANGL,RS)

THIS SUBROUTINE OBTAINS THE ROTATION MATRIX ABOUT THE
Z AXIS, GIVEN THE ANGLE OF ROTATION ABOUT THIS AXIS.

INPUTS:

ANGL ANGLE OF ROTATION IN RADIANS.

OUTPUTS:

RS ROTATION MATRIX.

AUTHOR: VALDEMIR CARRARA, APR/87

IMPLICIT REAL*8 (A-H,O-Z)
DIMENSION RS(3,3)

COAN = DCOS(ANGL)
SIAN = DSIN(ANGL)

RS(1,1) = COAN
RS(1,2) = SIAN
RS(1,3) = 0.DO
RS(2,1) = -SIAN
RS(2,2) = COAN
RS(2,3) = 0.DO
RS(3,1) = 0.DO
RS(3,2) = 0.DO
RS(3,3) = 1.DO

RETURN
END

```
SUBROUTINE UNIVVEC(VI,VO,DMDL)
```

```
THE SUBROUTINE UNIVVEC OBTAINS THE UNIT VECTOR AND ITS  
MAGNITUDE FROM THE INPUT VECTOR VI.
```

```
INPUTS:
```

```
VI      INPUT VECTOR.
```

```
OUTPUTS:
```

```
VO      UNIT VECTOR WITH THE SAME DIRECTION OF VI.  
DMDL    MAGNITUDE OF VI.
```

```
AUTHOR:  VALDEMIR CARRARA,  APR/87
```

```
IMPLICIT REAL*8(A-H,O-Z)  
DIMENSION VI(3),VO(3)
```

```
DMDL = DSQRT(VI(1)*VI(1)+VI(2)*VI(2)+VI(3)*VI(3))
```

```
IF(DMDL.EQ.0.DO) RETURN
```

```
VO(1) = VI(1)/DMDL
```

```
VO(2) = VI(2)/DMDL
```

```
VO(3) = VI(3)/DMDL
```

```
RETURN
```

```
END
```

```
SUBROUTINE VECINV(RM,VI,VO)
```

```
C
C THE SUBROUTINE VECINV CALCULATES THE COORDINATES OF A
C VECTOR IN THE FIRST REFERENCE SYSTEM, GIVEN THE
C MATRIX FROM THE FIRST TO THE SECOND FRAME AND THE
C VECTOR COORDINATES IN THE SECOND FRAME.
```

```
C
C INPUTS:
```

```
C   RM      ROTATION MATRIX.
C   VI      VECTOR IN THE SECOND SYSTEM.
```

```
C
C OUTPUTS:
```

```
C   VO      ROTATED VECTOR (VO = (RM**-1)*VI).
```

```
C
C AUTHOR:   VALDEMIR CARRARA, APR/87
```

```
C
C IMPLICIT REAL*8 (A-H,O-Z)
C DIMENSION RM(3,3),VI(3),VO(3)
```

```
DO 20 I=1,3
VO(I) = RM(1,I)*VI(1)+RM(2,I)*VI(2)+RM(3,I)*VI(3)
20 CONTINUE
```

```
RETURN
END
```

SUBROUTINE ROTSYS(IFRA,EL,RS).

THE SUBROUTINE ROTSYS OBTAINS THE ROTATION MATRIX FROM THE INERTIAL FRAME (X=VERNAL EQUINOX) TO THE FRAME DEFINED BY IFRA.

INPUTS:

IFRA ATTITUDE RELATED FRAME:
1- INERTIAL FRAME (X=VERNAL POINT)
2- X=ORBIT ASCENDING NODE, Z=NORTH POLE
3- X=ORBIT ASCENDING NODE, Z=ORBIT NORMAL
EL(6) KEPLERIAN ELEMENTS:
EL(1) SEMI-MAJOR AXIS, IN METERS.
EL(2) ECCENTRICITY
EL(3) INCLINATION IN RADIANS.
EL(4) RIGHT ASCENSION OF ASCENDING NODE IN RAD.
EL(5) PERIGEE ARGUMENT IN RAD.
EL(6) MEAN ANOMALY IN RAD.

OUTPUTS:

RS ROTATION MATRIX FROM THE INERTIAL FRAME TO THE DESIRED FRAME.

SUBROUTINES CALLED IN THIS SUBPROGRAM:

ROTMX
ROTMZ
MATMUL

AUTHOR: VALDEMIR CARRARA APR/87

IMPLICIT REAL*8 (A-H,O-Z)
DIMENSION EL(6),RS(3,3)
DIMENSION V1(3),V2(3),V3(3),R1(3,3),R2(3,3)

DATA V1 /1.DO,0.DO,0.DO/
DATA V2 /0.DO,1.DO,0.DO/
DATA V3 /0.DO,0.DO,1.DO/

IF(IFRA.LT.0.OR.IFRA.GT.4) GOTO 999

DO 20 I=1,3
RS(1,I) = V1(I)
RS(2,I) = V2(I)
RS(3,I) = V3(I)
20 CONTINUE

```
IF(IFRA.EQ.1) GOTO 100
CALL ROTMAZ(EL(4),RS)

IF(IFRA.EQ.2) GOTO 100
CALL ROTMAX(EL(3),R1)
DO 30 I=1,3
DO 30 J=1,3
R2(I,J) = RS(I,J)
30 CONTINUE
CALL MATMUL(R1,R2,RS)

100 CONTINUE
RETURN

999 CONTINUE
WRITE(1,*) ' FRAME SELECTOR OUT OF RANGE (1-4) -> ',
1 IFRA
STOP
END
```


SUBROUTINE ATMVEL(SV,AV)

THIS SUBROUTINE OBTAINS THE VELOCITY OF THE ATMOSPHERE.

INPUTS:

SV POSITION (SV(1), SV(2) AND SV(3)) IN METERS OF THE POINT IN THE INERTIAL FRAME.

OUTPUTS:

AV VELOCITY OF THE ATMOSPHERE IN METERS PER SECOND.

AUTHOR: VALDEMIR CARRARA APR/87

IMPLICIT REAL*8 (A-H,O-Z)
DIMENSION SV(6),AV(3)

COMMON /CONSTA/ PI,PIT2,PID2,RAD,DEG
COMMON /CEARTH/ RE,GM,FLAT,TETP

AV(1) = -SV(2)*TETP
AV(2) = SV(1)*TETP
AV(3) = 0.DO

RETURN
END

SUBROUTINE EARSHA(SA,SU,SHAD)

THE SUBROUTINE EARSHA VERIFIES IF THE SPACECRAFT IS OR
ISN'T IN THE EARTH SHADOW.

INPUTS:

SA(3) SPACECRAFT COORDINATES IN THE INERTIAL SYS-
TEM, IN METERS.
SU(3) COORDINATES OF THE SUN IN THE INERTIAL FRA-
ME, IN METERS.

OUTPUTS:

SHAD 0: SPACECRAFT IS IN THE EARTH SHADOW
1: SPACECRAFT IS ILLUMINATED BY THE SUN
BETWEEN 0 AND 1: SPACECRAFT IS IN THE EARTH
PENUMBRA.

AUTHOR: VALDEMIR CARRARA: APR/87

IMPLICIT REAL*8 (A-H,O-Z)

DIMENSION SA(6),SU(3)

COMMON /CONSTA/ PI,PIT2,PID2,RAD,DEG
COMMON /CEARTH/ RE,GM,FLAT,TETP

DATA RSUN /0.6953D+09/

DSUN = DSQRT(SU(1)*SU(1)+SU(2)*SU(2)+SU(3)*SU(3))
XSUN = SU(1)/DSUN
YSUN = SU(2)/DSUN
ZSUN = SU(3)/DSUN
RCOB = SA(1)*XSUN + SA(2)*YSUN + SA(3)*ZSUN

IF(RCOB.GE.0.DO) GOTO 200

RADI = RSUN/DSUN
XSVS = SA(2)*ZSUN - SA(3)*YSUN
YSVS = SA(3)*XSUN - SA(1)*ZSUN
ZSVS = SA(1)*YSUN - SA(2)*XSUN
PSVS = (DSQRT(XSVS*XSVS + YSVS*YSVS + ZSVS*ZSVS) -
RE)/RCOB

1

IF(ABS(PSVS).GE.RADI) GOTO 100

UDIV = PSVS/RADI

SHAD = (DACOS(UDIV) - UDIV*DSQRT(1.DO-UDIV*UDIV))/PI

RETURN

100 CONTINUE

IF(PSVS.LT.0.DO) GOTO 200

SHAD = 0.DO

RETURN

200 CONTINUE

SHAD = 1.DO

RETURN

END

FILE ANATM.FTN

FTN77,J
\$CDS ON

SUBROUTINE ASDAMO(SA,SU,SF,RJUD,DAFR,GSTI,TE,AD,WMOL,
1 RHOD)

C
C THE SUBROUTINE ASDAMO GIVES THE DENSITY, MOLECULAR
C WEIGHT AND TEMPERATURE OF THE UPPER ATMOSPHERE, USING
C THE ANALITICAL STATIC VERSION (LAFONTAINE -HUGHES) AND
C THE DYNAMIC ATMOSPHERIC MODEL (JACCHIA 1977).

INPUTS:

C SA(1) RIGHT ASCENTION OF THE POINT, IN RADIANS.
C SA(2) DECLINATION (GEOCENTRIC LATITUDE) OF THE
C POINT, IN RADIANS (-PI TO PI).
C SA(3) GEOCENTRIC ALTITUDE IN METERS, BETWEEN THE
C RANGE 90,000-2,000,000.
C SU(1) RIGHT ASCENTION OF THE SUN AT THE DATE, IN
C RADIANS (0 TO 2*PI).
C SU(2) SUN DECLINATION IN RADIANS (-PI TO PI).
C SF(1) DAILY SOLAR FLUX AT 10.7 CM, ADJUSTED FOR
C THE DISTANCE BETWEEN THE EARTH AND THE SUN
C AT THE DATE, IN 1E-22 W/M/M/HZ.
C SF(2) AVERAGED DAILY FLUX AS DEFINED BY JACCHIA,
C IN 1E-22 W/M/M/HZ, ADJUSTED FOR THE EARTH-
C SUN DISTANCE.
C SF(3) 3-HOURLY PLANETARY GEOMAGNETIC INDEX KP, AT
C THE TIME DAFR - TAU, WHERE TAU IS GIVEN BY
C JACCHIA'S 1977 MODEL.
C RJUD MODIFIED JULIAN DATE (JULIAN DATE-2400000).
C DAFR TIME (UT) OF THE DAY, IN SECONDS.
C GSTI GREENWICH SIDERAL TIME, IN RADIANS, AT THE
C TIME DAFR OF THE DATE RJUD (0 TO 2*PI).

OUTPUTS:

C TE(1) MEAN EXOSPHERIC TEMPERATURE ABOVE THE POINT
C AS DEFINED BY EQUATION 20 IN THE JACCHIA'S
C 1977 MODEL, IN KELVIN.
C TE(2) LOCAL TEMPERATURE AROUND THE POINT, IN
C KELVIN
C AD(1) LOGARITHM BASE 10 OF THE HE NUMBER-DENSITY.
C AD(2) LOGARITHM BASE 10 OF THE O2 NUMBER-DENSITY.
C AD(3) LOGARITHM BASE 10 OF THE N2 NUMBER-DENSITY.
C AD(4) LOGARITHM BASE 10 OF THE AR NUMBER-DENSITY.
C AD(5) LOGARITHM BASE 10 OF THE O NUMBER-DENSITY.
C AD(6) LOGARITHM BASE 10 OF THE H NUMBER-DENSITY.
C WMOL MEAN MOLECULAR WEIGHT OF THE ATMOSPHERE AT
C THE POINT, IN KG/KGMOL.

C RHOD MEAN MASS DENSITY OF THE ATMOSPHERE AT THE
C POINT, IN KG/M/M/M.
C OBS:
C HE HELIUM
C O2 MOLECULAR OXYGEN
C N2 MOLECULAR NITROGEN
C AR ARGON
C O ATOMIC OXYGEN
C H ATOMIC HYDROGEN

SUBROUTINES CALLED IN THIS SUB-PROGRAM:

C ASMADE
C DIVARA
C GEOACA
C SEALAT
C SEMIAN

REFERENCES:

- C [1] LAFONTAINE, J.; HUGHES, P. AN ANALYTIC VER-
C SION OF JACCHIA'S 1977 MODEL ATMOSPHERE.
C "CELESTIAL MECHANICS" 29(3-26) 1983.
C [2] JACCHIA, L. G. "THERMOSPHERIC TEMPERATURE,
C DENSITY AND COMPOSITION: NEW MODELS." CAM-
C BRIDGE, MA, SAO 1977. (SAO SPECIAL REPORT
C 375).

C AUTHOR: VALDEMIR CARRARA APR/87

C IMPLICIT REAL*8 (A-H,O-Z)

C DIMENSION SA(3),SU(2),SF(3),TE(2),AD(6)
C DIMENSION AL(6),WM(6)

C COMMON /STADE/ AC(6),TH(6)

C DATA AVOG /6.02217D+26/

C DATA WM /4.0026,31.9988,28.0134,39.948,15.9994,

1 1.00797/

C YEFR = (RJUD-.332825D05+DAFR/.86400D05)/.3652422D03

C TYFR = YEFR - IDINT(YEFR)

C GEAC = SF(3)

C RLAT = SA(2)

C RLON = SA(1) - GSTI

C HEIG = SA(3)

C ALTU = HEIG/1.D03

C SUDC = SU(2)

```
CALL ASMADE(HEIG,SF,TE,AL,WMOL,RHOD)

TEXO = TE(1)

CALL DIVARA(TEXO,SA,SU,WMOL,AL)

DO 100 I = 1, 6
AD(I) = AL(I)
100 CONTINUE

TQUT = TH(6)

CALL GEOACA(TQUT,GEAC,RLAT,RLON,ALTU,AL)

DO 200 I = 1, 6
AD(I) = AD(I) + AL(I) - AC(I)
200 CONTINUE

CALL SEALAT(TYFR,SUDC,RLAT,ALTU,AL)

DO 300 I = 1, 6
AD(I) = AD(I) + AL(I)
300 CONTINUE

CALL SEMIAN(TYFR,ALTU,ALCO)

DO 400 I = 1, 6
AD(I) = AD(I) + ALCO
400 CONTINUE

WEIG = 0.DO
ANUT = 0.DO

DO 500 I = 1, 6
ANAC = 10.DO**AD(I)
ANUT = ANUT + ANAC
WEIG = WEIG + WM(I)*ANAC
500 CONTINUE

WMOL = WEIG/ANUT
RHOD = WEIG/AVOG

RETURN
END
```

SUBROUTINE ASMADE(ALTU,SF,TE,AL,WMOL,RHOD)

THE SUBROUTINE CALCULATES THE ATMOSPHERIC DENSITY FOR HEIGHTS FROM 100 TO 2000 KM, USING THE ANALYTIC STATIC MODEL OF LAFONTAINE-HUGHES.

INPUTS:

ALTU ALTITUDE OF THE POINT IN METERS.
 SF(1) DAILY OBSERVED SOLAR FLUX AT 10.7 CM, IN
 1E-22 W/M/M/HZ.
 SF(2) AVERAGED DAILY FLUX, AS DEFINED BY JACCHIA,
 IN 1E-22 W/M/M/HZ.

OUTPUTS:

TE(1) MEAN EXOSPHERIC TEMPERATURE ABOVE THE POINT
 AS DEFINED BY EQUATION 20 IN THE JACCHIA'S
 1977 MODEL, IN KELVIN.
 TE(2) LOCAL TEMPERATURE AROUND THE POINT, IN
 KELVIN
 AL(1) LOGARITHM BASE 10 OF THE HE NUMBER-DENSITY.
 AL(2) LOGARITHM BASE 10 OF THE O2 NUMBER-DENSITY.
 AL(3) LOGARITHM BASE 10 OF THE N2 NUMBER-DENSITY.
 AL(4) LOGARITHM BASE 10 OF THE AR NUMBER-DENSITY.
 AL(5) LOGARITHM BASE 10 OF THE O NUMBER-DENSITY.
 AL(6) LOGARITHM BASE 10 OF THE H NUMBER-DENSITY.
 WMOL MEAN MOLECULAR WEIGHT OF THE ATMOSPHERE AT
 THE POINT, IN KG/KGMOL.
 RHOD MEAN MASS DENSITY OF THE ATMOSPHERE AT THE
 POINT, IN KG/M/M/M.

OBS:

HE HELIUM
 O2 MOLECULAR OXYGEN
 N2 MOLECULAR NITROGEN
 AR ARGON
 O ATOMIC OXYGEN
 H ATOMIC HYDROGEN

SUBROUTINES CALLED IN THIS SUB-PROGRAM:

AMOWEI
 TEMLO

REFERENCES:

- [1] LAFONTAINE, J.; HUGHES, P. AN ANALYTIC VER-
 SION OF JACCHIA'S 1977 MODEL ATMOSPHERE.
 "CELESTIAL MECHANICS" 29(3-26) 1983.
 [2] JACCHIA, L. G. "THERMOSPHERIC TEMPERATURE,
 DENSITY AND COMPOSITION: NEW MODELS." CAM-
 BRIDGE, MA, SAO 1977. (SAO SPECIAL REPORT
 375).

C
C
C

AUTHOR: VALDEMIR CARRARA APR/87

IMPLICIT REAL*8 (A-H,O-Z)

DIMENSION SF(3),TE(2),AL(6)
DIMENSION WM(6),CP(7)

DATA AVOG /6.02217D+26/

DATA PI /3.14159265359D0/

DATA WM /4.0026,31.9988,28.0134,39.948,15.9994,
1 1.00797/

DATA TO /188.D0/

DATA ZX /125.D0/

DATA ZO /90.0D0/

DATA BE /5.5D-5/

FLUX = SF(1)

FBAR = SF(2)

THAF = 5.48D0*(FBAR**.8D0) + 101.8D0*(FLUX**.4D0)

HEIG = ALTU/1.D3

CALL AMOWEI(THAF,HEIG,AL,WMOL,RHOD)

POTE = 0.0045D0*(THAF - TO)

ARGU = POTE/DSQRT(1.D0 + POTE*POTE)

TX = TO + 110.5D0*DATANH(ARGU)

GX = 1.9D0*(TX - TO)/(ZX - ZO)

CP(1) = 2.D0*(TX - TO)/PI

CP(2) = GX/CP(1)

CP(3) = 1.7D0*CP(2)

CP(4) = 2.D0*(THAF - TX)/PI

CP(5) = GX/CP(4)

CP(6) = BE*CP(5)

CP(7) = TX

TE(1) = THAF

TE(2) = TEMLO(HEIG,CP)

RETURN

END

SUBROUTINE DIVARA(TEXO,SA,SU,WMOL,AL)

THE SUBROUTINE CALCULATES THE DIURNAL VARIATIONS IN THE ATMOSPHERE DENSITY, AS PROPOSED BY JACCHIA, USING THE ANALYTIC STATIC MODEL BY LAFONTAINE-HUGHES.

INPUTS:

TEXO MEAN EXOSPHERIC TEMPERATURE, IN KELVIN.
 SA(1) RIGHT ASCENTION OF THE POINT, IN RADIANS.
 SA(2) DECLINATION (GEOCENTRIC LATITUDE) OF THE POINT, IN RADIANS (-PI TO PI).
 SA(3) GEOCENTRIC ALTITUDE IN METERS, BETWEEN THE RANGE 90,000-2,000,000.
 SU(1) RIGHT ASCENTION OF THE SUN. AT THE DATE, IN RADIANS (0 TO 2*PI).
 SU(2) SUN DECLINATION IN RADIANS (-PI TO PI).
 WMOL MEAN MOLECULAR WEIGHT AT THE LOCAL REGION, IN KG/KGMOL.

OUTPUTS:

AL(1) LOGARITHM BASE 10 OF THE HE NUMBER-DENSITY.
 AL(2) LOGARITHM BASE 10 OF THE O2 NUMBER-DENSITY.
 AL(3) LOGARITHM BASE 10 OF THE N2 NUMBER-DENSITY.
 AL(4) LOGARITHM BASE 10 OF THE AR NUMBER-DENSITY.
 AL(5) LOGARITHM BASE 10 OF THE O NUMBER-DENSITY.
 AL(6) LOGARITHM BASE 10 OF THE H NUMBER-DENSITY.
 WMOL MEAN MOLECULAR WEIGHT OF THE ATMOSPHERE AT THE POINT, IN KG/KGMOL.
 RHOD MEAN MASS DENSITY OF THE ATMOSPHERE AT THE POINT, IN KG/M/M/M.

OBS:

HE HELIUM
 O2 MOLECULAR OXYGEN
 N2 MOLECULAR NITROGEN
 AR ARGON
 O ATOMIC OXYGEN
 H ATOMIC HYDROGEN

COMMON OUTPUTS:

COMMON /STADE/ AC(6),TH(6)
 AC SAME AS AL, BUT USING THE PSEUDO EXOSPHERIC TEMPERATURE FOR THE HYDROGEN.
 TH(1) EXOSPHERIC TEMPERATURE FOR HE, IN K.
 TH(2) EXOSPHERIC TEMPERATURE FOR O2, IN K.
 TH(3) EXOSPHERIC TEMPERATURE FOR N2, IN K.
 TH(4) EXOSPHERIC TEMPERATURE FOR AR, IN K.
 TH(5) EXOSPHERIC TEMPERATURE FOR O, IN K.
 TH(6) EXOSPHERIC TEMPERATURE FOR H, IN K.

C SUBROUTINES CALLED IN THIS SUB-PROGRAM:
C AMOWEI
C

REFERENCES:

- C [1] LAFONTAINE, J.; HUGHES, P. AN ANALYTIC VER-
C SION OF JACCHIA'S 1977 MODEL ATMOSPHERE.
C "CELESTIAL MECHANICS" 29(3-26) 1983.
C [2] JACCHIA, L. G. "THERMOSPHERIC TEMPERATURE,
C DENSITY AND COMPOSITION: NEW MODELS." CAM-
C BRIDGE, MA, SAO 1977. (SAO SPECIAL REPORT
C 375).
C

C AUTHOR: VALDEMIR CARRARA APR/87
C

IMPLICIT REAL*8 (A-H,O-Z)

DIMENSION SA(3),SU(2),AL(6)
DIMENSION AN(6),WM(6),CP(7)

COMMON /STADE/ AC(6),TH(6)

DATA AVOG. /6.02217D+26/
DATA PI /3.14159265359D0/
DATA RAD /1.7453292778D-2/

DATA WM /4.0026,31.9988,28.0134,39.948,15.9994,
1 1.00797/

RLAT = SA(2)
DELT = SU(2)
SOAN = SA(1) - SU(1)
ALTU = SA(3)/1.D3
ARGU = RLAT*RLAT/1.5707963D0
COAR = DCOS(ARGU)
COLA = DCOS(RLAT)
E1 = 2.D0 + .COAR*COAR
TEAU = 1.D0 + 0.3666069D0*DELT*DSIN(RLAT)

DO 100 I = 1, 5
BISH = RAD*(27.D0*(WMOL/WM(I) - 1.D0) - 35.D0)
ARGU = SOAN + BISH
FHAG = 0.08D0*DCOS(3.D0*ARGU - 1.3089969D0) +
1 DABS(DCOS(0.5D0*ARGU))**E1
TH(I) = (0.24D0*COLA*(FHAG - 0.5D0) + TEAU)*TEXO
100 CONTINUE

ARGU = SOAN - 1.0471976D0
FHAG = 0.08D0*DCOS(3.D0*ARGU - 1.3089969D0) +
1 DABS(DCOS(0.5D0*ARGU))**E1
TH(6) = TEXO*(TEAU + 0.24*COLA*(FHAG - 0.5D0))

```
DO 200 I = 1, 6
TEMP   = TH(I)
CALL AMOWEI(TEMP,ALTU,AN,WEIG,RHOD)
AL(I)  = AN(I)
200 CONTINUE

DO 300 I = 1, 6
AC(I)  = AN(I)
300 CONTINUE

RETURN
END
```

FUNCTION TEMLO(ALTU,C)

THE SUBROUTINE TEMLO EVALUATES THE TEMPERATURE AT THE LOCAL REGION.

INPUTS:

ALTU ALTITUDE OF THE POINT, IN KM (90 TO 2000).
C ARRAY THAT CONTAINS THE PARAMETERS USED IN THE EVALUATION OF EQUATIONS 3 AND 4, AS GIVEN BY JACCHIA.

OUTPUTS:

TEMLO LOCAL TEMPERATURE AS DEFINED IN EQUATIONS 3 AND 4, IN KELVIN.

REFERENCES:

[1] JACCHIA, L. G. "THERMOSPHERIC TEMPERATURE, DENSITY AND COMPOSITION: NEW MODELS." CAMBRIDGE, MA, SAO 1977. (SAO SPECIAL REPORT 375).

AUTHOR: VALDEMIR CARRARA APR/87

IMPLICIT REAL*8 (A-H,O-Z)

DIMENSION C(7)

DATA ZX /125.DO/
DATA ZO /90.OO0/
DATA TO /188.DO/

HIGX = ALTU - ZX
HIGO = ALTU - ZO
TEMLO = TO

IF(HIGX.GT.0.DO) GOTO 100

IF(HIGO.EQ.0.DO) RETURN

AUXI = HIGX/HIGO

TEMLO = C(7) +

1 C(1)*DATAN(C(2)*HIGX + C(3)*HIGX*AUXI*AUXI)

RETURN

100 CONTINUE

TEMLO = C(7) +

1 C(4)*DATAN(C(5)*HIGX + C(6)*HIGX*HIGX*HIGX)

RETURN

END

SUBROUTINE AMOWEI (TEXO,ALTU,AN,WMOL,RHOD)

THIS SUBROUTINE CALCULATES MEAN MOLECULAR WEIGHT OF THE ATMOSPHERE IN THE LOCAL REGION, USING THE LAFONTAINE-HUGHES VERSION FOR THE JACCHIA'S 1977 MODEL.

INPUTS:

TEXO MEAN EXOSPHERIC TEMPERATURE, IN KELVIN.
ALTU GEOMETRICAL HEIGHT WHERE THE MOLECULAR WEIGHT WILL BE CALCULATED, IN KM (90-2000).

OUTPUTS:

AN(1) LOGARITHM BASE 10 OF THE HE NUMBER-DENSITY.
AN(2) LOGARITHM BASE 10 OF THE O2 NUMBER-DENSITY.
AN(3) LOGARITHM BASE 10 OF THE N2 NUMBER-DENSITY.
AN(4) LOGARITHM BASE 10 OF THE AR NUMBER-DENSITY.
AN(5) LOGARITHM BASE 10 OF THE O NUMBER-DENSITY.
AN(6) LOGARITHM BASE 10 OF THE H NUMBER-DENSITY.
WMOL MEAN MOLECULAR WEIGHT OF THE ATMOSPHERE AT THE POINT, IN KG/KGMOL.
RHOD MEAN MASS DENSITY OF THE ATMOSPHERE AT THE POINT, IN KG/M/M/M.

OBS:

HE HELIUM
O2 MOLECULAR OXYGEN
N2 MOLECULAR NITROGEN
AR ARGON
O ATOMIC OXYGEN
H ATOMIC HYDROGEN

REFERENCES:

- [1] LAFONTAINE, J.; HUGHES, P. AN ANALYTIC VERSION OF JACCHIA'S 1977 MODEL ATMOSPHERE. "CELESTIAL MECHANICS" 29(3-26) 1983.
[2] JACCHIA, L. G. "THERMOSPHERIC TEMPERATURE, DENSITY AND COMPOSITION: NEW MODELS." CAMBRIDGE, MA, SAO 1977. (SAO SPECIAL REPORT 375).

AUTHOR: VALDEMIR CARRARA APR/87

IMPLICIT REAL*8 (A-H,O-Z)

DIMENSION AN(6)
DIMENSION CC(6),QI(6),AP(6),WM(6)
DIMENSION CE(9),AA(9),BE(9),BN(6),CN(6)

DATA PI /3.14159265359D0/
DATA CONV /2.302585092994D0/
DATA AVOG /6.02217D+26/

DATA CC /28.575551D0,-0.472715D0,-0.118328D0,
 1 0.054405D0,0.009942D0,-0.005744D0/
 DATA QI /0.78110D0,0.20955D0,0.009343D0,.5242D-5,
 1 0.D0;0.D0/
 DATA AP /0.D0,0.D0,0.D0,-0.38D0,-0.25D0,0.D0/
 DATA WM /28.0134,31.9988,39.948,4.0026,1.00797,
 1 15.9994/
 DATA RA /6356.766D0/
 DATA GA /9.806650D0/
 DATA CTGP /8.314320D0/
 DATA ZX /125.0D0 /
 DATA ZO /90.0D0 /
 DATA ZH /100.0D0 /
 DATA TO /188.0D0 /
 DATA RHOI /3.43D-06 /
 DATA EMEI /28.960D0 /
 DATA EPSI /1.D-06 /

AKXI = 2.D0*(RA + ZX)/(ZX - ZO)
 VH = AKXI*(ZH - ZO)/(RA + ZH) - 1.D0
 WO = (1.D0 - VH)/(1.D0 + VH)
 W1 = WO + 1.D0

CE(1) = CC(1) + WO*(CC(2) + WO*(CC(3) + WO*CC(4)))
 CE(2) = W1*(CC(2) + WO*(2.D0*CC(3) + 3.D0*WO*CC(4)))
 CE(3) = W1*W1*(CC(3) + 3.D0*WO*CC(4))
 CE(4) = W1*W1*W1*CC(4)
 CE(5) = 0.D0
 CE(6) = 0.D0
 CE(7) = 0.D0
 CE(8) = 0.D0
 CE(9) = 0.D0

ARGU = 0.0045D0*(TEXO - TO)
 AUXI = ARGU/DSQRT(1.D0 + ARGU*ARGU)
 TX = TO + 110.5D0*DATANH(AUXI)
 TAUO = (TX + TO)/(TX - TO)
 TAU1 = 2.D0*TO*TX/(TO - TX)
 GDxE = 1.9D0*(TX - TO)/(ZX - ZO)
 GDxE = 0.475D0*TO*(RA + ZX)/TX/(RA + ZO)

AA(5) = 0.06205282D0*DLOG(TEXO + 213.9884D0) -
 1 0.6286968D0
 AA(6) = 0.06555110D0*DLOG(TEXO - 329.6454D0) -
 1 0.1520990D0
 AA(1) = AA(5) - GDxE - TAUO
 AA(2) = AA(6) - GDxE + 1.5D0
 AA(3) = GDxE - 2.D0*AA(5)
 AA(4) = GDxE - 2.D0*AA(6) - 0.5D0
 AA(7) = 0.D0
 AA(8) = 0.D0
 AA(9) = 0.D0

```

Z      = ZH
IF(ALTU.LT.ZH) Z      = ALTU

VE     = AKXI*(Z - ZO)/(RA + Z) - 1.DO
TAUC   = AA(1) + VE*(AA(2) + VE*(AA(3) + VE*(AA(4) +
1      VE*(AA(5) + VE*AA(6))))))
TCHA   = TAU1/TAUC
GAVE   = GA*RA*RA/AKXI/(RA + ZO)

DO 220 I = 1, 9
BE(I)  = 0.DO

DO 200 J = 1, I
BE(I)  = BE(I) + CE(J)*AA(I+1-J)
200 CONTINUE
220 CONTINUE

BEAU   = 0.DO
EMCH   = 0.DO
VHEI   = 1.DO
VEEI   = 1.DO

DO 250 I = 1, 9
VEEI   = VEEI*VE
BEAU   = BEAU + BE(I)*(VEEI + 2.DO*(DBLE(I) -
1      2.DO*DBLE(INT(I/2)))) - 1.DO)/DBLE(I)
EMCH   = EMCH + CE(I)*VHEI
VHEI   = VEEI
250 CONTINUE

BETA   = EMCH/TCHA
ARGU   = DEXP(-GAVE*BEAU/CTGP/TAU1)
RHOD   = BETA*ARGU*TO*RHOI/EMEI
AUXI   = AVOG*RHOD
DENU   = AUXI/EMCH
ARGU   = AUXI/EMEI

BN(1)  = QI(1)*ARGU
BN(3)  = QI(3)*ARGU
BN(4)  = QI(4)*ARGU
CN(2)  = DENU*(EMCH*(QI(2) + 1.DO)/EMEI - 1.DO)
CN(6)  = 2.DO*DENU*(1.DO - EMCH/EMEI)
AUXI   = DLOG(CN(2))/CONV -
1      0.07DO*(1.DO + DTANH(0.18DO*(Z - 111.DO)))
BN(2)  = 10.DO**AUXI
AUXI   = DLOG(CN(6))/CONV -
1      0.24DO*DEXP(-.009DO*(Z-97.7DO))*(Z-97.7DO))
BN(6)  = 10.DO**AUXI
BN(5)  = 1.DO

IF(ALTU.LE.ZH) GOTO 800

```



```

Z      = ZX
IF(ALTU.LT.ZX) Z      = ALTU

VE     = AKXI*(Z - ZO)/(RA + Z) - 1.DO
TCHA   = AA(1) + VE*(AA(2) + VE*(AA(3) + VE*(AA(4) +
1      VE*(AA(5) + VE*AA(6))))))
THAU   = AA(1) + VH*(AA(2) + VH*(AA(3) + VH*(AA(4) +
1      VH*(AA(5) + VH*AA(6))))))
THAU   = TAU1/THAU
TCHA   = TAU1/TCHA
USAR   = THAU/TCHA
BN(2)  = CN(2)
AUXI   = 0.DO
VEEI   = VE
VHEI   = VH

DO 380 I = 1, 6
AUXI   = AUXI + AA(I)*(VEEI - VHEI)/DBLE(I)
VEEI   = VEEI*VE
VHEI   = VHEI*VH
380 CONTINUE

DO 400 I = 1, 4
ARGU   = -AUXI*WM(I)*GAVE/CTGP/TAU1
USAR   = USAR**(1.DO + AP(I))
BN(I)  = BN(I)*USAR*DEXP(ARGU)
400 CONTINUE

ARGU   = -AUXI*WM(6)*GAVE/CTGP/TAU1

CN(6)  = CN(6)*THAU/TCHA*DEXP(ARGU)
CN(2)  = BN(2)
AUXI   = DLOG(CN(2))/CONV -
1      0.07DO*(1.DO + DTANH(0.18DO*(Z - 111.DO)))
BN(2)  = 10.DO**AUXI
AUXI   = DLOG(CN(6))/CONV -
1      0.24DO*DEXP(-.009DO*(Z-97.7DO)*(Z-97.7DO))
BN(6)  = 10.DO**AUXI

IF(ALTU.LE.ZX) GOTO 800

DCAP   = DEXP(8.042617DO)*TEXO**(-0.4197668DO) +
1      22.58421DO -
2      0.05719352DO*TEXO*DEXP(-TEXO/1187.417DO)
GAMF   = (TEXO - TX)/GDEX/(RA + ZX)
ALFA   = GAMF*DCAP/(GAMF*DCAP - 1.DO)
ASUT   = DCAP/ALFA/(TX - TEXO)
BSUT   = -2.DO*ASUT*TEXO - DCAP
CSUT   = TEXO*(TEXO*ASUT + DCAP)

```

```

ARGU = ((ALFA/EPSI*(TX-TEXO) +
1     1.DO)/(1.DO-ALFA)**(1.DO/DCAP)
XINF = DLOG(ARGU)
ZINF = (RA*XINF + ZX)/(1.DO - XINF)
HAGA = ALTU

```

```
IF(HAGA.GE.ZINF) HAGA = ZINF
```

```

XOFZ = (HAGA - ZX)/(RA + HAGA)
AUXI = DEXP(-DCAP*XOFZ)
ARGU = ALFA*(TX - TEXO)*AUXI
TCHA = ARGU/(AUXI + ALFA - 1.DO) + TEXO
GAMA = GA*RA*RA/CTGP/(RA + ZX)/CSUT
BN(2) = CN(2)

```

```

ARGU = (ASUT*TX + BSUT)*TX + CSUT
ABBA = ((ASUT*TCHA + BSUT)*TCHA + CSUT)/ARGU
DTDO = TCHA - TEXO
DTD1 = TEXO - TX
DTD2 = ALFA/(1.DO - ALFA)
AAAA = DTD1*DTD2/DTDO + 1.DO/(1.DO - ALFA)

```

```

DO 600 I = 1, 4
GAAU = GAMA*WM(I)
AUXI = BN(I)*(TX/TCHA)**(1.DO + AP(I) + GAAU)
BN(I) = AUXI*(AAAA**(GAAU*BSUT/2.DO/DCAP))*
1     (ABBA**(GAAU/2.DO))

```

```
600 CONTINUE
```

```

CN(2) = BN(2)
AUXI = DLOG(CN(2))/CONV -
1     0.07DO*(1.DO + DTANH(0.18DO*(HAGA - 111.DO)))
BN(2) = 10.DO**AUXI
AUXI = CN(6)*AAAA**(GAMA*WM(6)*BSUT/2.DO/DCAP)
AUXI = AUXI*ABBA**(GAMA*WM(6)/2.DO)
CN(6) = AUXI*(TX/TCHA)**(1.DO + AP(6) + GAMA*WM(6))
AUXI = DLOG(CN(6))/CONV - 0.24DO*
1     DEXP(-.009DO*(HAGA - 97.7DO)*(HAGA - 97.7DO))
BN(6) = 10.DO**AUXI

```

```

HDEN = 10.DO**(TEXO**(-.25DO)*28.9DO + 5.94DO)
X500 = (500.DO - ZX)/(RA + 500.DO)

```

```

AUXI = DEXP(-X500*DCAP)
T500 = ALFA*(TX - TEXO)*AUXI/(AUXI+ALFA-1.DO) + TEXO
AUXI = (TCHA*ASUT + BSUT)*TCHA + CSUT
AUXI = AUXI/((ASUT*T500 + BSUT)*T500 + CSUT)
ARGU = GAMA*WM(5)*BSUT/2.DO*(XOFZ - X500)
AUXI = HDEN*AUXI**(GAMA*WM(5)/2.DO)*DEXP(ARGU)
BN(5) = AUXI*(T500/TCHA)**(1.DO + AP(5) + GAMA*WM(5))

```

```
IF(HAGA.EQ.ALTU) GOTO 750
X1      = 1.DO/(1.DO + ZINF/RA)
X2      = 1.DO/(1.DO + ALTU/RA)
COEF    = GA*RA/CTGP/TEXO*(X2 - X1)
BN(2)   = CN(2)
BN(6)   = CN(6)

DO 700 I = 1, 6
BN(I)   = BN(I)*DEXP(COEF*WM(I))
700 CONTINUE

AUXI    = DLOG(BN(2))/CONV -
1       0.07DO*(1.DO + DTANH(0.18DO*(ALTU - 111.DO)))
BN(2)   = 10.DO**AUXI
AUXI    = DLOG(BN(6))/CONV - 0.24DO*
1       DEXP(-.009DO*(ALTU - 97.7DO)*(ALTU - 97.7DO))
BN(6)   = 10.DO**AUXI

750 CONTINUE

DO 760 I = 1, 6
IF(BN(I).LT:1.D-30) BN(I) = 1.D-30
760 CONTINUE

800 CONTINUE
WMOL    = 0.DO
ARGU    = 0.DO

DO 850 I = 1, 6
AUXI    = BN(I)
ARGU    = ARGU + AUXI
WMOL    = WMOL + AUXI*WM(I)
850 CONTINUE

RHOD    = WMOL/AYOG
WMOL    = WMOL/ARGU

AN(1)   = DLOG(BN(4))/CONV
AN(2)   = DLOG(BN(2))/CONV
AN(3)   = DLOG(BN(1))/CONV
AN(4)   = DLOG(BN(3))/CONV
AN(5)   = DLOG(BN(6))/CONV
AN(6)   = DLOG(BN(5))/CONV

RETURN
END
```

SUBROUTINE GEOACA(TQUT,GEAC,RLAT,RLON,ALTU,DN)

THE SUBROUTINE GEOACA OBTAINS THE VARIATION OF THE ATMOSPHERE NUMBER-DENSITY, DUE TO THE GEOMAGNETIC ACTIVITY, USING THE ANALYTIC STATIC MODEL:

INPUTS:

TQUT EXOSPHERIC TEMPERATURE FOR THE HYDROGEN ABOVE THE POINT, IN KELVIN.
GEAC THE PLANETARY THREE-HOUR INDEX, KP, IN THE RANGE 0. TO 9.0
RLAT DECLINATION (GEOCENTRIC LATITUDE) OF THE POINT, IN RADIANS (-PI TO PI).
RLON LOCAL LONGITUDE, IN RADIANS (-PI TO PI).
ALTU GEOCENTRIC ALTITUDE IN KM, IN THE RANGE 90. TO 2000..

OUTPUTS:

DN ARRAY WITH THE COMMON LOGARITHM (BASE 10), OF THE VARIATION IN THE NUMBER DENSITY OF THE HE (HELIUM), O2 (MOLECULAR OXYGEN), N2 (MOLECULAR NITROGEN), AR (ARGON), O (ATOMIC OXYGEN) AND H (ATOMIC HYDROGEN), RESPECTIVELY, DUE TO THE GEOMAGNETIC ACTIVITY.

SUBROUTINES CALLED IN THIS SUB-PROGRAM:

AMOWEI

REFERENCES:

- [1] LAFONTAINE, J.; HUGHES, P. AN ANALYTIC VERSION OF JACCHIA'S 1977 MODEL ATMOSPHERE. "CELESTIAL MECHANICS" 29(3-26) 1983.
[2] JACCHIA, L. G. "THERMOSPHERIC TEMPERATURE, DENSITY AND COMPOSITION: NEW MODELS." CAMBRIDGE, MA, SAO 1977. (SAO SPECIAL REPORT 375).

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IMPLICIT REAL*8 (A-H,O-Z)

DIMENSION DN(6)

DIMENSION GM(6)

DATA GM /-6.30D-05,1.03D-05,0.D0,3.07D-05,-4.85D-05,
1 0.D0/

```
AUXI = 57.5DO*GEAC*(1.DO + 0.027DO*DEXP(0.4DO*GEAC))
SILA = DSIN(RLAT)
SENO = 0.9792*SILA +
1    0.2028*DCOS(RLAT)*DCOS(RLON - 5.0789081DO)
SEN2 = SENO*SENO
COL2 = 1.DO - SEN2
DELT = AUXI*SEN2*SEN2
TEMP = TQUT + DELT

CALL AMOWEI(TEMP,ALTU, DN, WMOL, RHOD)

DELT = 0.01DO*DELT
DEZH = 5000.DO*DLOG(DSQRT(1.DO + DELT*DELT) + DELT)
DELE = 5.2D-04*AUXI*COL2*COL2

DO 100 I = 1, 6
DN(I) = DN(I) + GM(I)*DEZH + DELE
100 CONTINUE

RETURN
END
```


SUBROUTINE SEMIAN(TYFR,ALTU,ALCO)

THE SUBROUTINE SEMIAN GIVES THE CORRECTION FACTOR ALCO FOR THE ATMOSPHERE NUMBER DENSITY, DUE TO THE SEMIANNUAL EFFECT.

INPUTS:

TYFR FRACTION OF THE TROPIC YEAR, IN THE RANGE 0. TO 1., STARTING ON JAN. 1ST.
ALTU GEOCENTRIC ALTITUDE IN KM, IN THE RANGE 90. TO 2000..

OUTPUTS:

ALCO THE SEMIANNUAL VARIATION OF THE ATMOSPHERE NUMBER DENSITY.

REFERENCES:

[1] JACCHIA, L. G. "THERMOSPHERIC TEMPERATURE, DENSITY AND COMPOSITION: NEW MODELS." CAMBRIDGE, MA, SAO 1977. (SAO SPECIAL REPORT 375).

AUTHOR: VALDEMIR CARRARA APR/87

IMPLICIT REAL*8 (A-H,O-Z)

DATA PITW /6.28318530718D0/

AUX1 = 0.04D0*ALTU*ALTU/1.D+04 + 0.05D0
FOFT = AUX1*DEXP(-0.25D-02*ALTU)
AUX1 = (0.5D0+0.5D0*DSIN(PITW*TYFR+6.04D0))*1.65D0
TAUC = 0.0954D0*(AUX1 - 0.5D0) + TYFR
AUX1 = DSIN(2.D0*PITW*TAUC + 4.26D0)*(1.D0 +
0.467D0*DSIN(PITW*TAUC + 4.14D0))
GOFT = AUX1*0.382D0 + 0.0284D0
ALCO = FOFT*GOFT

RETURN
END